14055 Regolith Breccia 111 grams



Figure 1: Photos of 14055, showing two sides. Sample is about 4 cm round. NASA S71-29572 and 29565.



Figure 2: Map of Apollo 14 showing location of 14055.

Transcript from station E

LMP This is a big crater. It's 40, 50 meters across. It has a fairly sharp crater in the south edge of it, which is -

CC Okay, that looks like it may be the one by E.

LMP - 20, 30 feet across. Yes, I think that's it Fredo. And it's – no, it's at least 50 or 60 feet deep.

CDR Why don't we just grab a couple from right here.

LMP Yes, Okay.

CDR That baby came apart. Very soft.

LMP Yes, it's falling apart as you pick it up, very crumbly, isn't it?

CDRVery, ver soft rock – rim of that crater, plus anotherone very close to us with crystals in it, now, going into bagLMP15-N.





Figure 3: 14055 is not a Fra Mauro breccia, but rather a vitric matrix breccia made from the A14 soil.

Introduction

From the transcript, it is clear that the astronauts put two friable samples in bag 15N, but when it was opened in the LRL there were seven fragments with a lot of "residue". 14055 and 14058 are more coherent than the other fragments (figure 1). They are also blocky, subangular to subrounded rocks lightly covered with glass-lined zap pits. The samples are friable, finegrained clastic rocks with 5 to 15 % of subrounded light-colored clasts in a medium-gray matrix.



Figure 4: Photomicrograph of thin section 14055,9. Section is 1 cm long. S71-43119.

Stoffler et al. (1979) consider 14055 a classic regolith breccia. Fruland (1983) and Simon et al. (1989) included 14055 in their suite of regolith breccias from Apollo 14.

The Apollo 14 regolith breccias (vitric matrix breccias) are slightly more aluminous than the Fra Mauro breccias (crystalline matrix breccias).

Petrography

Both rocks collected at station E were friable and returned in the same bag (15N). Swann et al. (1977) reported that 14058 may be part of 14055, while 14056, 14057, 14059, 14060 and 14061 are very friable and may be pieces of the other sample. Fruland (1983) and Simon et al. (1989) included 14055 is the suite of regolith breccias. It is very like 14042, 14047 and 14049.

Mineralogical Mode for 14055

-	
	Drozd et al. 1977
Mineral fragment	s 28.7 %
Lithic fragments	14.9
Colored glass	5.5
Agglutinate glass	49.5
Colorless glass	0.6
Chondrules	0
Devitrified glass	0.6

Mineralogical Mode for 14055

	Simonds et	Simon et	Drozd et
	al. 1977	al. 1989	al. 1976
Matrix	72 %	55.9	
Clasts			
Plagioclase	1	4.1	
Mafic	0.5	5.9	
Breccia	9		
Glass	5.5	8.6	6.7
Granulite	11	1.4	
Agglutinate		8.9	49.5

14055 originally had a thin layer of brown glass on one side, which has since mostly rubbed off.

Simon et al. (1989) and Drozd et al. (1976) reported a high percentage of agglutinates in 14055, while Simonds et al. (1977) found a lot of fine-grained matrix material (figure 3). The sample Drozd et al. studied may have included the glass splash on the outer surface.

Unusual fragments of ropy glass are seen in thin sections (figures 5 and 8).

Chemistry

The chemical composition of 14055 is remarkably similar to that of the Apollo 14 regolith (as measured on 14163) – see table.

Other Studies

Drozd et al. (1975) determined Kr and Xe concentration and isotopic ratios. Hart et al. (1972) studied the solar flare tracks. Horz et al. (1972) studied the micrometeorite craters.

> Next page Figure 5: Photomicrogaph of thin section 14055,42 shoing large fragment of ropy glass. Photo by C. Meyer.



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Figure 6: Composition of 14055 compared with other Apollo 14 breccias and bulk soil 14163.



Figure 7: Normalized rare-earth-element diagram for 14055 compared with KREEP.

Processing

14055 was returned in ALSRC 1006, which was sealed. It was originally designated as a "posterity" sample – to be studied later. There were no surface photos taken at station E. Twedell et al. (1978) made maps of the surface of 14055 (figure 9). There are 14 thin sections.



Figure 8: Ropy glass fragment in 14055. Scale is 2.8 mm across.

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reference weight	Simon89 154 ma		14163 Laul80	Carlson I.C. and Walton W.J.A. (1978) Apollo 14 Rock Samples. Curators Office. JSC 14240	
SIO2 % TIO2 Al2O3 FeO MnO MgO CaO Na2O K2O P2O5 S %	1.73 17.4 10.4 0.126 10 10 0.73 0.59	(a) (a) (a) (a)	47.3 1.6 17.8 10.5 0.135	Chao E.C.T., Minkin J.A. and Best J.B. (1972) Apollo 14 breccias: General characteristics and classification. <i>Proc. 3rd Lunar Sci. Conf.</i> 645-659.	
		(a) (a) (a) (a)	9.6 11.4 0.7 0.55	Drozd R., Hohenberg C. and Morgan C. (1975) Krypton and xenon in Apollo 14 samples: Fission and neutron capture effects in gas-rich samples. <i>Proc.</i> 6 th <i>Lunar Sci. Conf.</i> 1857- 1877.	
Sc ppm V Cr Co Ni Cu Zn Ga Ge ppb As Se Rb Sr Y Zr Nb Mo	21.8 47 1280	(a) (a) (a)	21.7 45 1368	Drozd R.J., Kennedy B.M., Morgan C.J., Podosek F.A. and Taylor G.J. (1976) The excess fission xenon problem in lunar samples. <i>Proc.</i> 7 th <i>Lunar Sci. Conf.</i> 599-623.	
	32 390	(a) (a)) 33) 350	Hart H.R., Comstock G.M. and Fleischer R.L. (1972) The particle track record of Fra Mauro. <i>Proc.</i> 3 rd <i>Lunar Sci. Conf.</i> 2831-2844.	
	19 120	(a) (a)	170	Hörz F., Morrison D.A. and Hartung J.B. (1972) The surface orientation of some Apollo 14 rocks. <i>Modern Geology</i> 3 , 93-104.	
	730	(a)		LSPET (1971) Preliminary examination of lunar samples from Apollo 14. <i>Science</i> 173 , 681-693.	
Ru Rh Pd ppb				Fruland R.M. (1983) Regolith Breccia Workbook. JSC 19045	
Cd ppb In ppb Sn ppb Sb ppb Te ppb				Simon S.B., Papike J.J., Shearer C.K., Hughes S.S. and Schmitt R.A. (1989) Petrology of Apollo 14 regolith breccias and ion microprobe studies of glass beads. <i>Proc. 19th Lunar</i> <i>Planet. Sci. Conf.</i> 1-17.	
Cs ppm Ba La Ce Pr	0.63 910 71.4 179	(a) (a) (a) (a)	800 66.7 170	Simonds C.H., Phinney W.C., Warner J.L., McGee P.E., Geeslin J., Brown R.W. and Rhodes J.M. (1977) Apollo 14 revisited, or breccias aren't so bad after all. <i>Proc.</i> 8 th Lunar	
Nd Sm Eu	107 31 2.7	(a) (a) (a)	100 29.1 2.45	<i>Sci. Conf.</i> 1869-1893.	
Gd Tb Dy Ho Er Tm Yb Lu Hf Ta W ppb Re ppb Os ppb Ir ppb Pt ppb Au ppb Th ppm U ppm technique:	6 36	(a) (a)	5.9 36 8.6	Grain size statistics, composition and provenance of fragmental particles in some Apollo 14 breccias. <i>Proc.</i> 7 th <i>Lunar Sci. Conf.</i> 1965-1985.	
	22.1 2.9 20.6 3	(a) (a) (a) (a)	3.2 21.2 3 22.5 2.9	Stöffler D., Knoll HD. and Maerz U. (1979) Terrestrial and lunar impact breccias and the classification of lunar rocks. <i>Proc.</i> 10 th Lunar Planet. Sci. Conf. 639-675.	
	5	(a)		Sutton R.L., Hait M.H. and Swann G.A. (1972) Geology of the Apollo 14 landing site. <i>Proc.</i> 3 rd Lunar Sci. Conf. 27-38.	
	6.2 12 3.1 <i>(a) INAA</i>	(a) (a) (a)	13.3 3.5		

References for 14055

Table 1. Chemical composition of 14055.



Swann G.A., Trask N.J., Hait M.H. and Sutton R.L. (1971a) Geologic setting of the Apollo 14 samples. *Science* **173**, 716-719.

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Wilshire H.G. and Jackson E.D. (1972) Petrology and stratigraphy of the Fra Mauro Formation at the Apollo 14 site. U.S. Geol. Survey Prof. Paper **785**.



Figure 9: Map of original surface of 14055 showing large glass splash (Twedell et al. 1978). S77-23605.