# 10032

# Ilmenite Basalt 3.1 grams



Figure 1: Lunar basalt fragment 10032. NASA S75-31697. Cube is 1 cm.

#### **Introduction**

10032 was collected as part of the contingency sample from the regolith in front of the LM (Sutton and Schaber 1971). It is a fine grained ilmenite basalt (figure 1).

The age of 10032 has been determined as 3.6 b.y. with an exposure to cosmic rays of 140 m.y.

## **Petrography**

Beaty and Albee (1978) studied the mineral chemistry in 10032 and group it other fine-grained vesicular specimen in the high K suite of Apollo 11 basalts. It is ilmenite rich and has a grain size about 100 micron. High K glass is found in the intersticies.

# **Mineralogy**

**Olivine:** The most magnesium olivine in 10032 is  $Fo_{69}$  (Beaty and Albee 1978).

**Pyroxene:** Some of the large augite grains in 10032 have pigeonite cores (figure 2).

**Plagioclase:** Plagioclase in 10032 is An<sub>78-70</sub>.

*Ilmenite:* Ilmenite has about 1 % MgO.

# **Mineralogical Mode of 10032**

	Beaty and
	<b>Albee 1978</b>
Olivine	0.11
Pyroxene	52.61
Plagioclase	22.16
Ilmenite	16.45
mesostasis	6.66
silica	0.94
troilite	0.72
phosphate	0.24

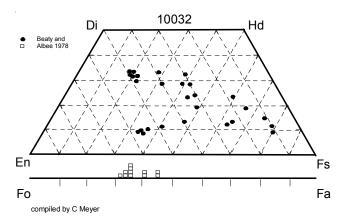


Figure 2: Pyroxene and olivine composition of 10032 (from Beaty and Albee 1978).

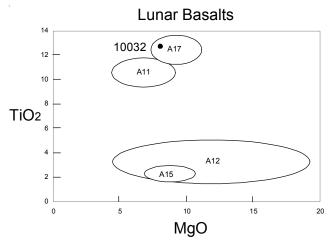


Figure 3: Composition of 10032 compared with other Apollo basalts..

**Phosphate:** The phosphate has 2.3 % fluorine.

# **Chemistry**

The composition of 10032 has been determined by calculation based on mineral mode and mineral composition (table 1, figure 3).

# Radiogenic age dating

The age of 10032 was determined aboput 3.6 b.y. by the Ar/Ar plateau technique (Guggisberg et al. 1979) (figure 4).

# Cosmogenic isotopes and exposure ages

Guggisberg et al. (1979) determined and <sup>37</sup>Ar/<sup>38</sup>Ar age of 139 b.y.

#### **Other Studies**

The total organic carbon content of 10032 was determined by hydrogen flame ionization pyrolysis (Ponnamperuma et al. 1970).

Funkhouser et al. (1970) and Bogard et al. (1971) reported the abundance and isotopic composition of rare gasses from 10032.

# **Processing**

Apollo 11 samples were originally described and cataloged in 1969 and "re-cataloged" by Kramer et al. (1977). There are 7 thin sections.

**List of Photo #s for 10032** S75-31697

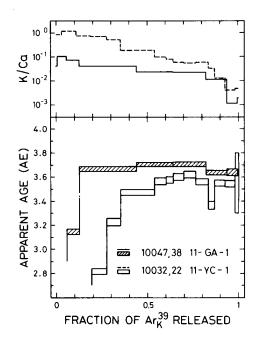


Figure 4: Argon release pattern (Guggisberg et al. 1979).

## **Summary of Age Data for 10032**

Ar/Ar plateau

Guggisberg et al. 1979  $3.58 \pm 0.06$  b.y. (poorly defined)

Table 1. Chemical composition of 10032.

Sc ppm V Cr Co Ni Cu Zn Ga Ge ppb As Se Rb Sr Y Zr Nb Mo Ru Rh Pd ppb Ag ppb Cd ppb In ppb Sn ppb Sb ppb Te ppb Cs ppm Ba La Ce Pr Nd Sm Eu Gd Tb Dy Ho Er Tm Yb Lu Hf Ta W ppb Re ppb Os ppb Ir ppb Au ppb Au ppb Th ppm U ppm technique: (a) elec. Probe + point count	reference weight SiO2 % TiO2 Al2O3 FeO MnO MgO CaO Na2O K2O P2O5 S % sum	Beaty 78 39.49 13.06 7.08 19.37 0.27 8.81 10.21 0.45 0.28 0.11 0.36	(a)
gao. (a) oloo. I lobo · bollic coult	V Cr Co Ni Cu Zn Ga Ge ppb As Se Rb Sr Y Zr Nb Mo Ru Rh ppb ppb ppb Te ppb Cs ppm Ba La Ce Pr Nd Sm Eu GTb Dy Ho Er Tm Yb Lu Hf Ta W eppb bt ppb Pt ppb Lu Hf Ta W eppb Lu ppm U ppm U ppm	(a) elec.	Probe + point count

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