

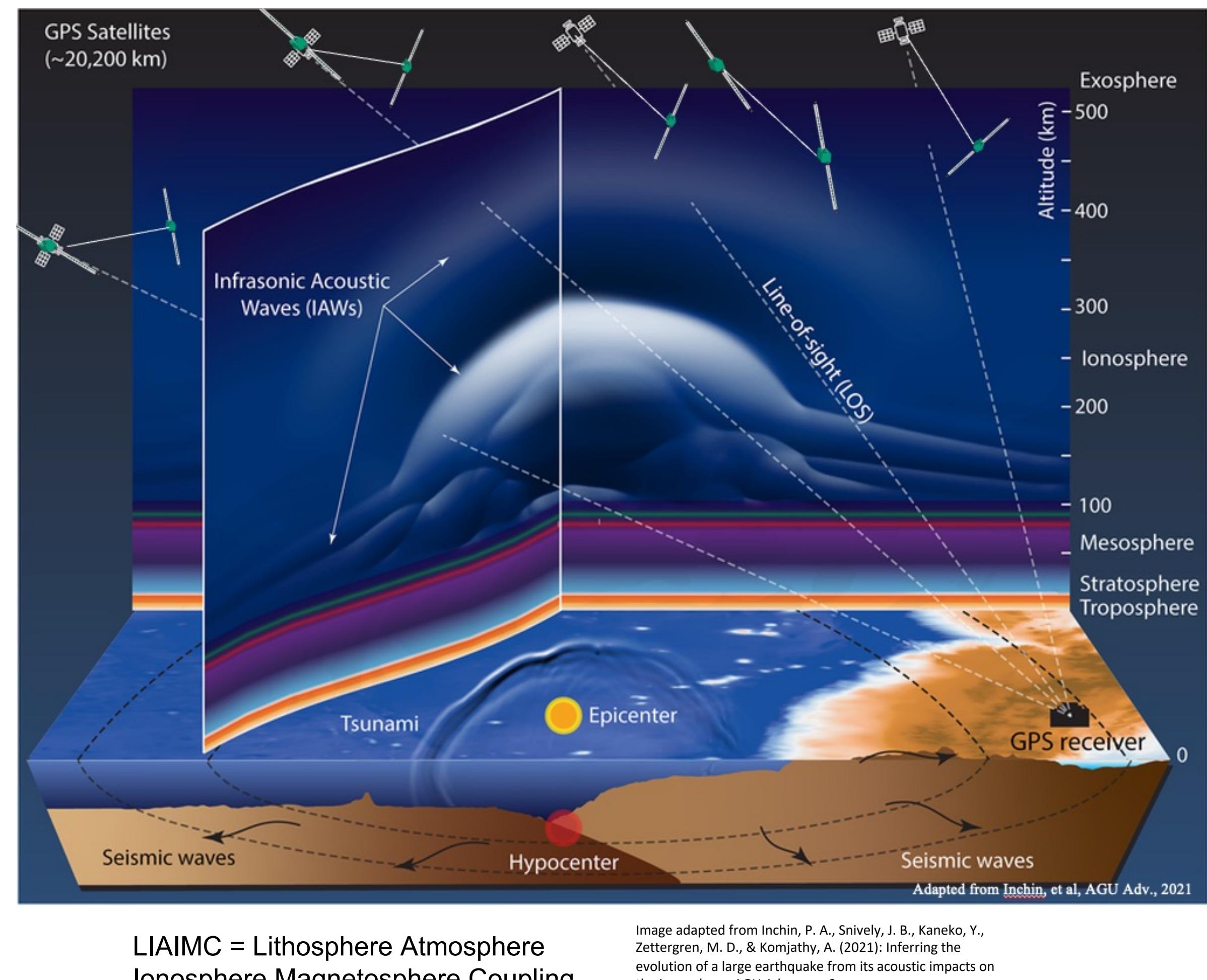
PULSAR: Planetary pULSe-tAkeR

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Innovation

- PULSAR leverages the coupling mechanism between the lithosphere, atmosphere, ionosphere, magnetosphere, and the plasmasphere of planets.
- PULSAR operates as a large baseline variable sensitivity detector network that can be reconfigured on demand.
- PULSAR enables new distributed measurements of coupled physical events occurring at different spatial and temporal scales operating as a planet-wide detector.
- Like a doctor taking vitals and monitoring the health state of a patient, PULSAR literally "takes the pulse of the planet".**
- Phase I applies PULSAR to the Earth geospace and Phase II and beyond extends PULSAR applicability to other planets with an ionosphere, such as Venus, Mars, Titan, and the Gas Giants.



Vision: take "vitals" of the Solar System in real-time

Almost all planets and many moons in the Solar System are geophysically active

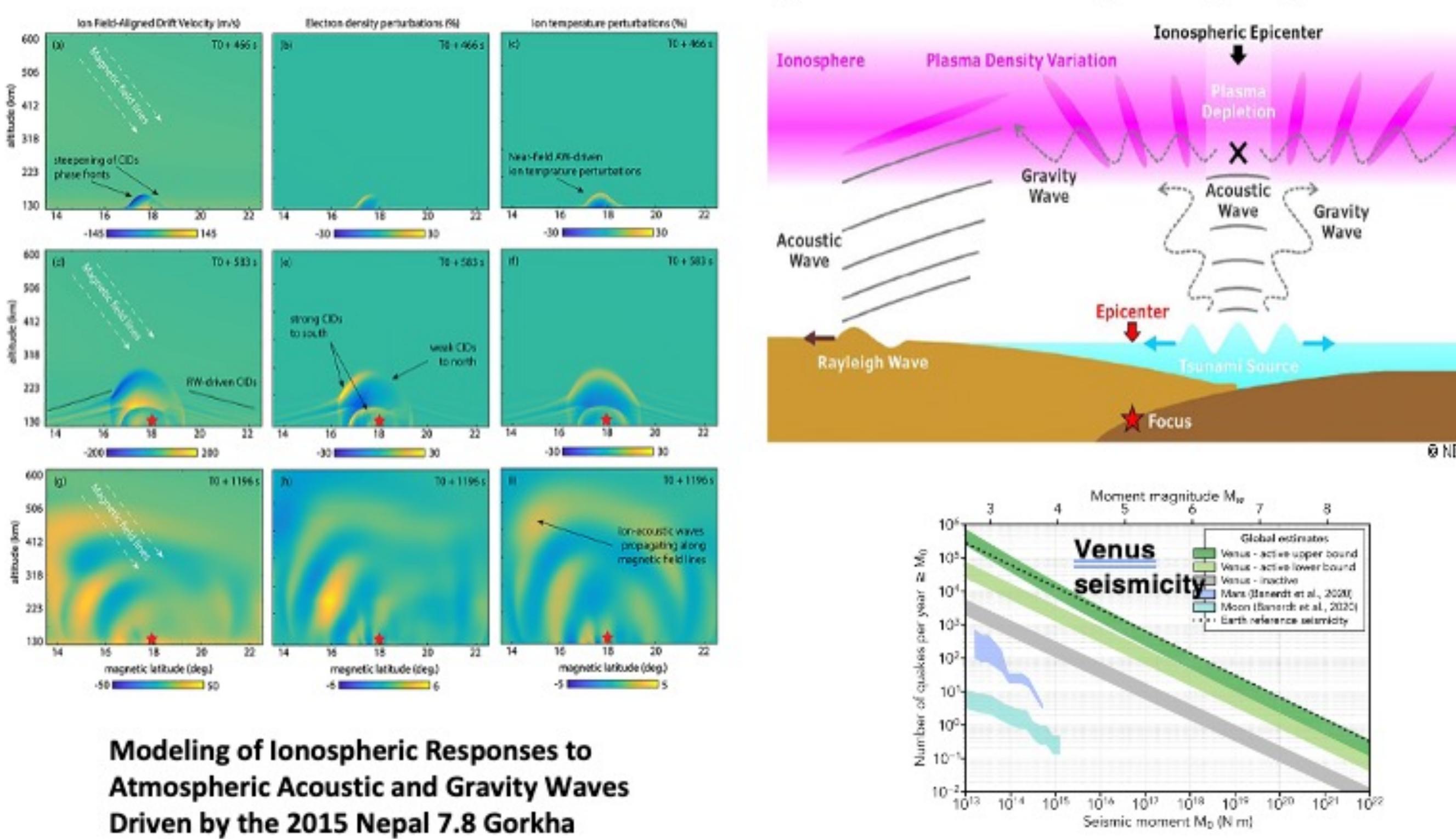
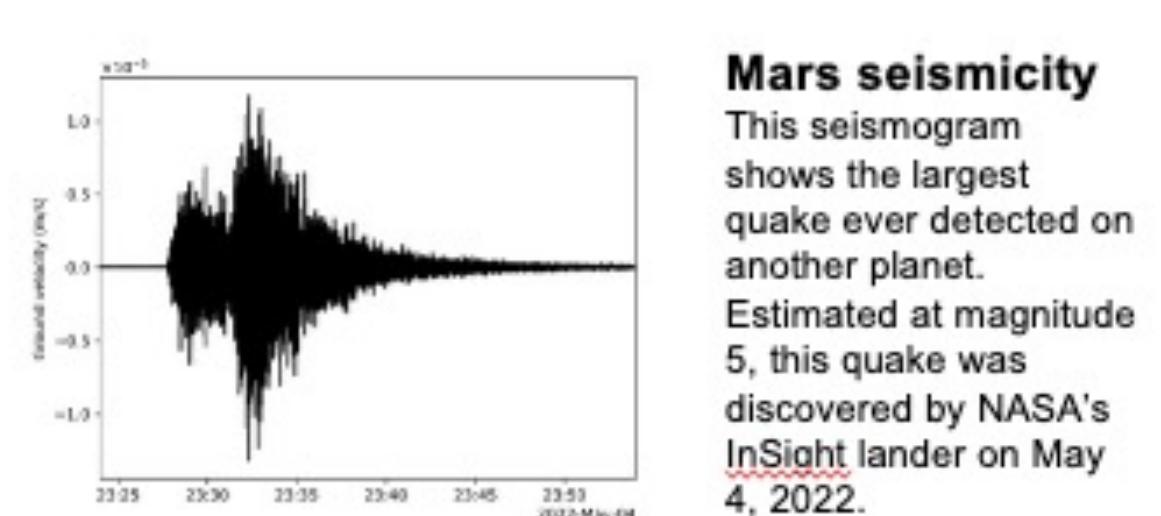
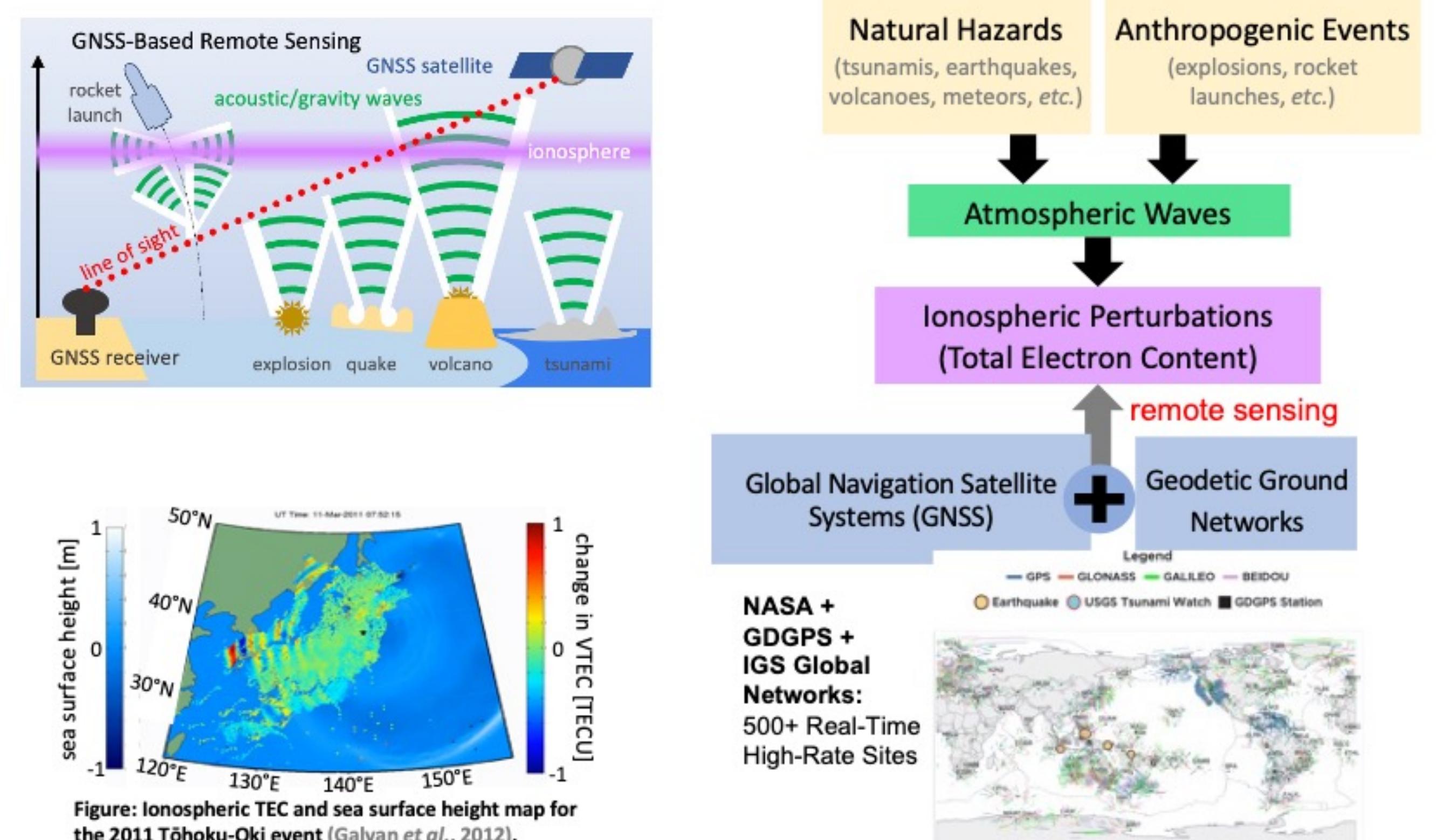


Illustration for coupling mechanisms for solid Earth, ocean, and ionosphere. Epicentral acoustic after 8 minutes (AGW_{epic}), gravity (IGW_{trans}) after 1 hour, and Rayleigh (surface or AW_{Rayleg}) wave reaching ionosphere in about 10 minutes, detectable by available GPS receivers. Teleseismic distances can be > 1000 km.



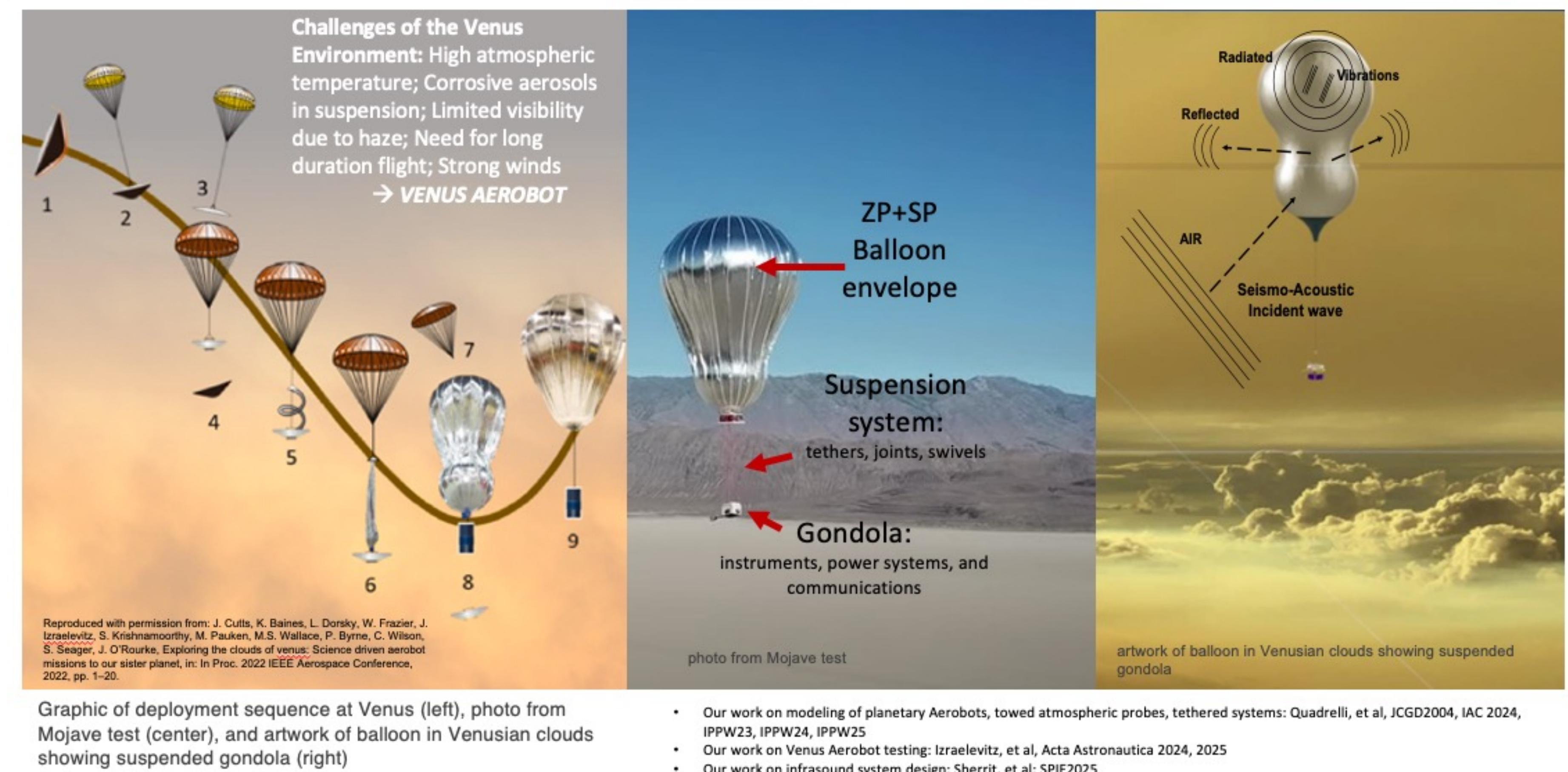
<https://www.nasa.gov/missions/inSight/nasainsight-records-monster-quake-on-mars/>

Much work is being done in our geospace



- <https://guardian.jpl.nasa.gov>
 (GNSS-based Upper Atmospheric Realtime Disaster Information and Alert Network)
- GUARDIAN is a near-real-time (NRT) ionospheric monitoring software for natural hazards warning.
 - GUARDIAN uses total electron content (TEC) time series to allow users to explore ionospheric TEC perturbations due to natural and anthropogenic events on Earth and characterize potential natural hazards.
- Martire, L., Krishnamoorthy, S., Vergados, P., Romano, L.L., Solisguy, B., Merello, K., Anderson, J.L., Komjathy, A., Bar-Sever, Y.E. (2023) The GUARDIAN system - a GNSS upper atmospheric real-time disaster information and alert network, *GPS Solutions* 27(32).

...and for Venus LIAIMC exploration with Aerobots

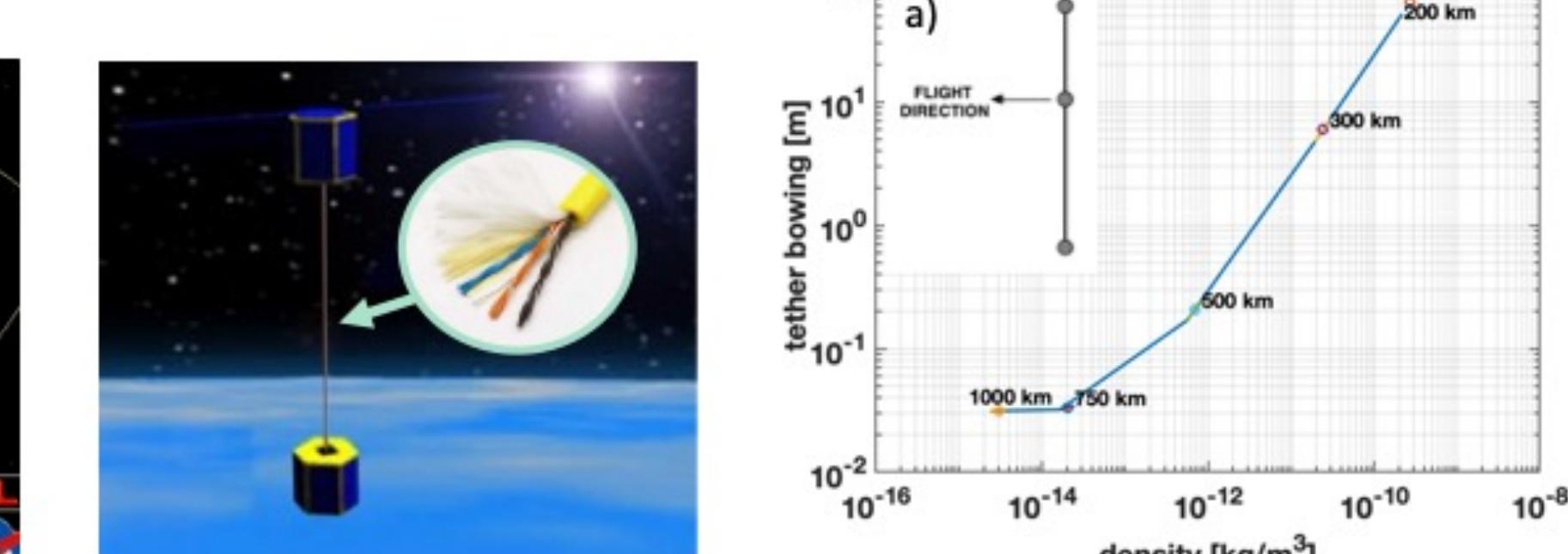


- Our work on modeling of planetary Aerobots, towed atmospheric probes, tethered systems: Quadrelli, et al, ICGD2004, IAC 2024, IPPV23, IPPV24, IPPV25
- Our work on Venus Aerobot testing: Izrailevitz, et al, Acta Astronautica 2024, 2025
- Our work on infrasound system design: Sherri, et al: SPIE2025

How does PULSAR work



- Instrument miniaturization and systems integration and multifunctionality are enablers
- If sufficient data can be returned and large baselines can be implemented, new scientific observations are possible
- Instrument resolution (strength, direction) is dictated by spatial sampling
- High sensitivity in the frequency band of interest is dictated by temporal sampling
- Many nodes leads to graceful degradation.
- Tethered constellations, lead to variable and very large detector sizes (even 1000's of km), at a fraction of the mass
→ the longer the tether, the more sensitive the detector, becoming a low frequency acoustic gradiometer

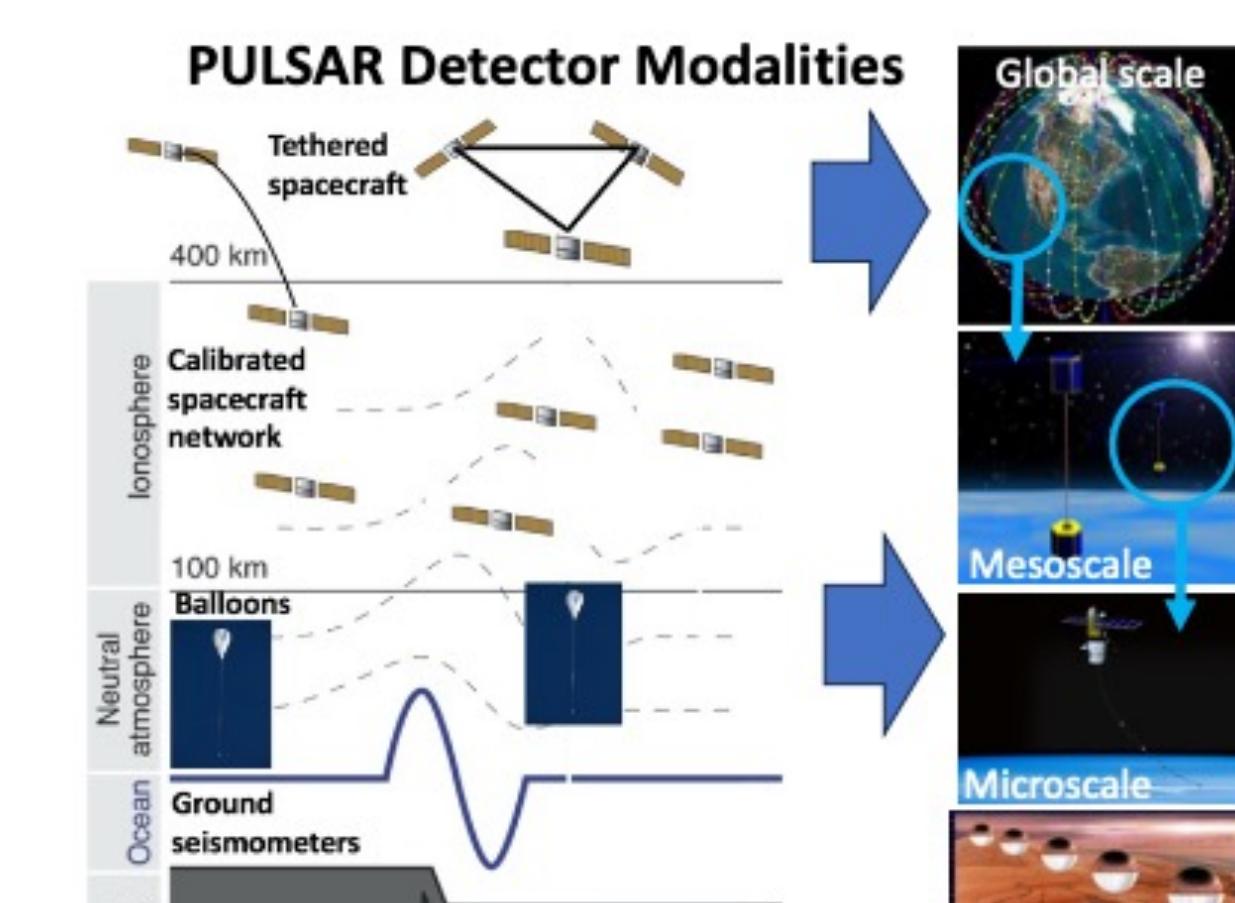


- The TEC changes are related to the pressure and density changes generated by the surface wave as they propagate upwards.
- For a surface displacement at the centimeter level, generated by events >Mag7, the pressure at H= 300 km altitude (F-layer of ionosphere) may be detected resulting in TEC changes at about 0.5 TEC units.
- This corresponds to about 8 cm range changes between the satellite and receivers, which is within the capability of existing radio-frequency metrology.
- Ionospheric signals from GNSS measurements can routinely discern the signal from the constant background noise level at approx. 0.05 TEC by using time of arrival, coherence, wave amplitude and period.

Phase I studies the feasibility of PULSAR to Earth, limiting the study to a seismic lithospheric trigger event of the geospace as a pilot example.

Phase II: Address challenges of non-terrestrial applications:

- No GNSS
- Planetary LIAIMC phenomena still largely unexplored
- Missions are infrequent
- Long-duration, distant, limited communication/data
- System resources are very limited
- No opportunity for maintenance



Science Goal	Observables	Instrument	Data Product
Orbital constellation to study near-Earth EM, plasma, and particle environment	Wave phenomena, ducting, and generation mechanisms of ULF, VLF, HF emissions.	Main UFD, antenna options: three orthogonal dipoles, two V-shaped dipoles, two crossed dipoles, Mass spectrometer, inlet discrimination along the satellite velocity vector. Magnetometer. Long-period probe. Energetic particle detector.	Waveform data (time series) of electric and magnetic fields, ULF/VLF/HF emission characteristics.
A planetary seismology mission using one or more balloons	Seismic event magnitude and location	An array of micro barometers can be deployed below a balloon. Another tether to detect seismic infrasound in range 3-10 Hz at 0.01 – 1 Pa	Vector infrasound time series of pressure levels at balloon receivers

Science traceability matrix

Phase I Study

Global: Reconfigurability as multi-objective optimization problem, Formation flight, GN&C, Cent./Decent. ops, Up-down link, HPC, Data storage

Mesoscale: Electrodynamic (bare) tethers: As it flies over the epicentral region, it senses peaks up to about $-10 V, 0.01 A$.

Microscale: Sensitivity analysis to detect impulsive density disturbance. Angle range: micro-degrees; Tether tension: nano-strain sensitivity; Radial acceleration: micro-g's → GRACE-level IMU

Distributed phased arrays: End-fire vs. formation flying arrays Efficiency and bandwidth of electrically small antennas at ELF/ULF/VLF frequency bands. Phased array coherence. Timing and signal synchronization. Polarization

- Next steps:
- Update Science Traceability Matrix
 - Continue analyses
 - Prepare final report

