



Low coherence wavefront probe for nanometer level free-form metrology

NASA SBIR Phase II: NNX17CG04C

April 2017 Start

December 2018 Finish

- The Problem: Metrology of Freeforms
- Free Form Metrology Technical Limitations
- Probe Technology
- Progress to Date
- Commercial Impact
- Äpre Instruments background

Free form optics have known advantages.
Lack of metrology limits manufacture

- Basic Metrology Requirements
 - Low Measurement Uncertainty (<50 nm rms 2σ)
 - Measure “Any” Shape (includes aspheres & spheres)
 - High speed (<5 minutes)
 - Surface slopes $\leq 60^\circ$
 - Form & Mid-Spatial Frequencies (0.1 to 4 mega points)
- Nominally <150 mm X 150 mm parts
 - Up to 500 mm X 500 mm parts

NASA Project Goals

- Free form 2D profile map
- 2 nm 2σ RMS measurement uncertainty
 - $10^3 < \text{CMM}$, and $10^2 < \text{State of the Art}$
- Non-contact
- Large acceptance angle: $>30^\circ$
- 50 mm X 50 mm measurement area

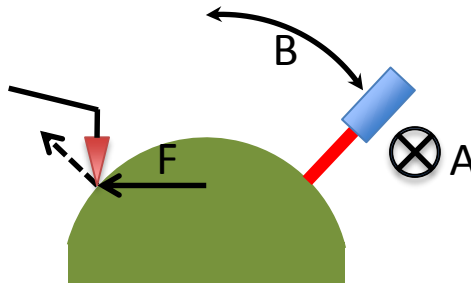
Today's Free form Metrology Technical Limitations

Touch Probes

- Pros
 - Any roughness
 - $\leq 60^\circ$ slopes
 - Low cost (probe)
 - X,Y,Z metrology frame
- Cons
 - Surface Damage
 - Slow (10's of minutes)
 - Low image resolution
 - Limited accuracy
 - Ball sphericity
 - Side loads
 - Z-sensitivity cosine error

Interferometer Probe

- Pros
 - Non-contact
 - nm Z-sensitivity
 - Normal to surface
 - Promise high image resolution
- Cons
 - Polished only
 - Moderately Slow (1 kHz)
 - 2° slope limitation
 - \$ > contact probe \$
 - Limited accuracy
 - X,Y,Z,A,B metrology frame required
 - Retrace errors – off normal



Ultimate Probe Combines Best of Both

Ultimate Probe

Simple X,Y,Z Metrology Frame
Steep Angles



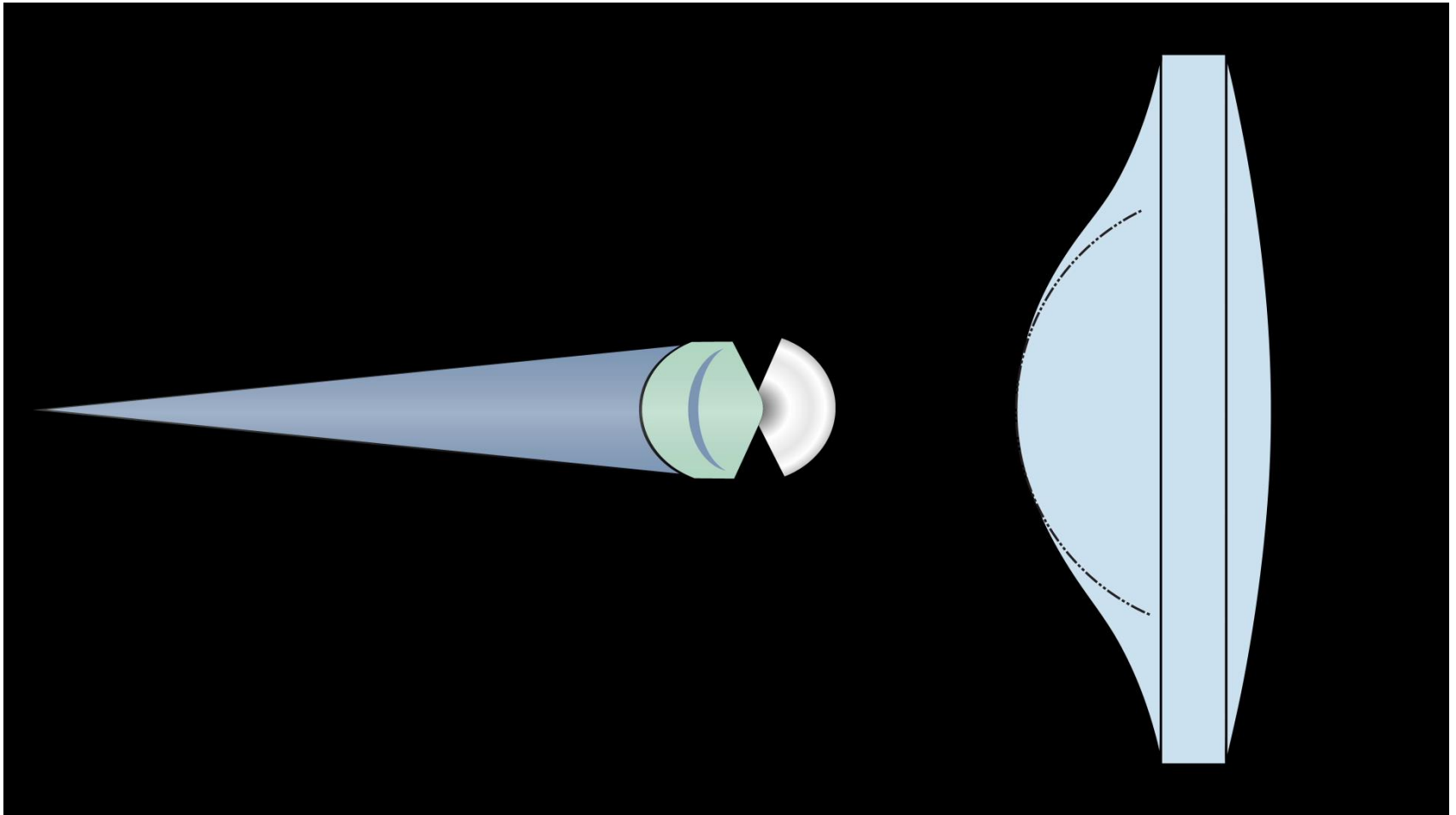
Non-Contact
4 KHz data acquisition
Normal to surface measurement
Nanometer level sensitivity

Three Core Technologies

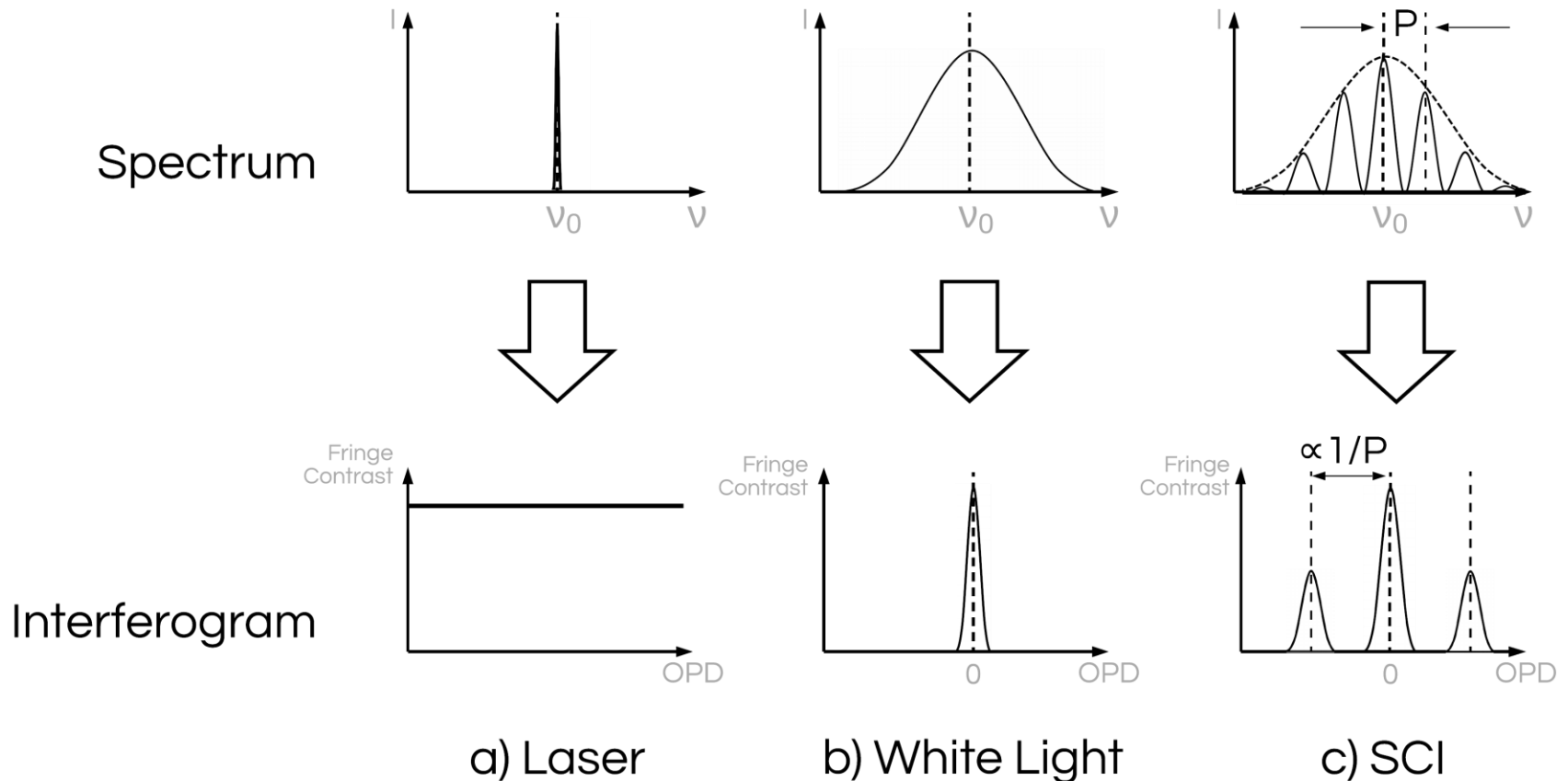
- Spectrally Controlled Interferometry (SCI)¹
- Divergent Spherical Wavefront Fizeau with localized interference¹
- Signal Detection/Analysis

¹ US & International Patents and Patents Pending

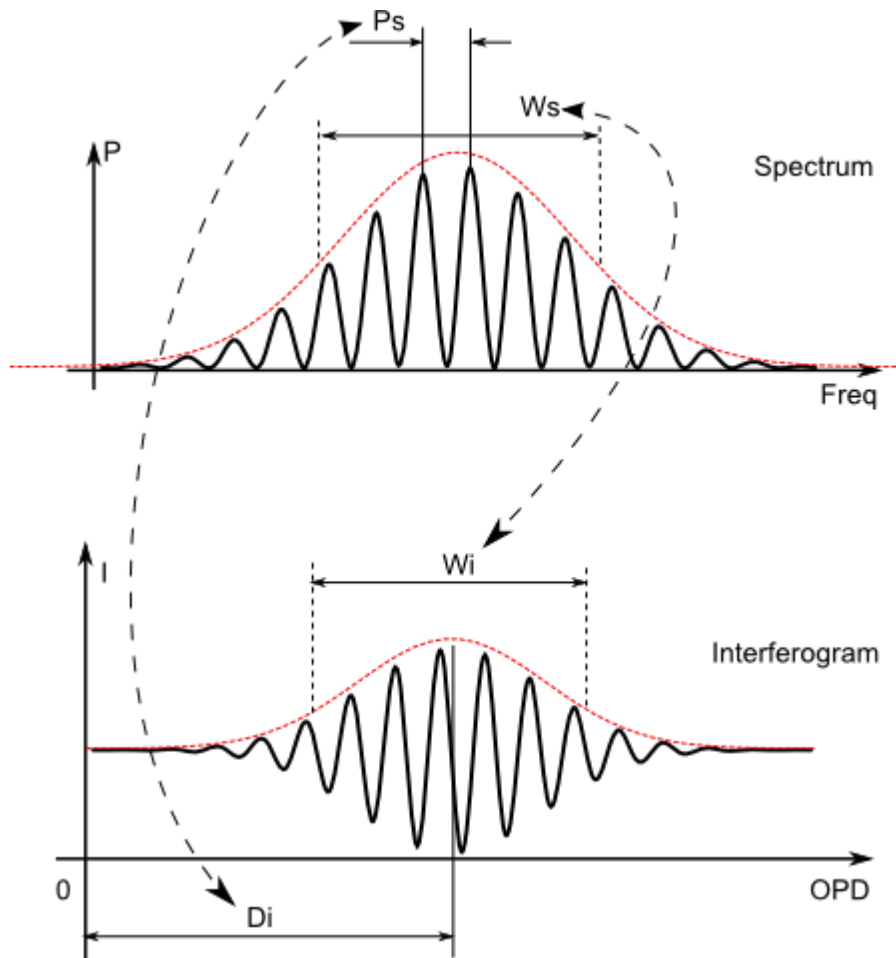
The V-Probe



Source Spectrum Drives Interference Fringe Formation



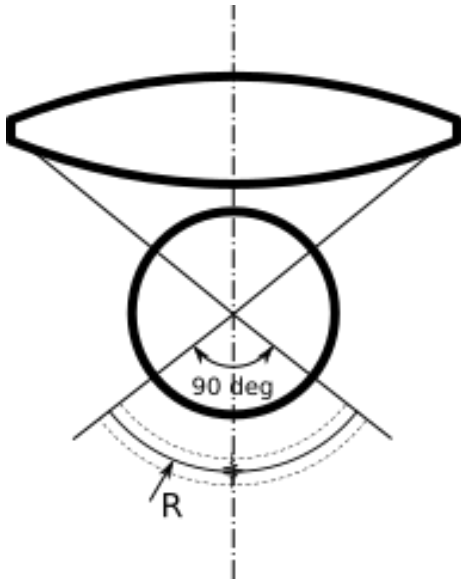
Three SCI Control Parameters



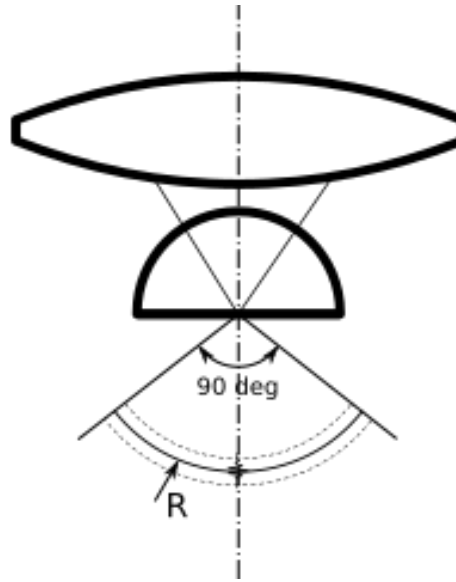
- $W_s \propto 1/W_i$
- $P_s \propto 1/D_i$
- $\phi_s = \phi_i$

KEY POINT: Fixing the source period (P_s) under a broad band source (W_i) creates a fixed narrow region of interference in front of a simple Fizeau interferometer

Spherical Diverging SCI Fizeau

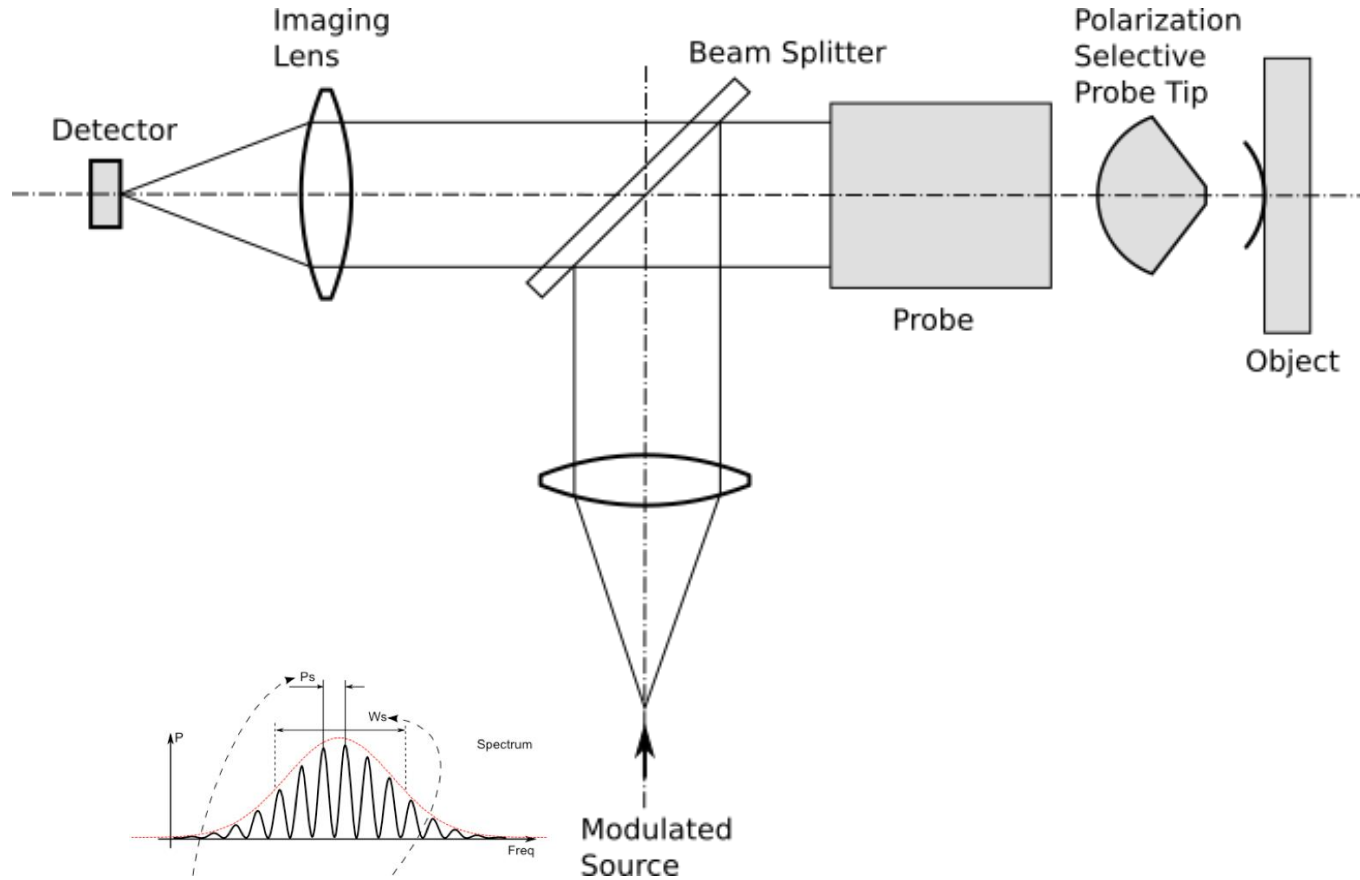


Fizeau Reference:
Ball Lens Surface

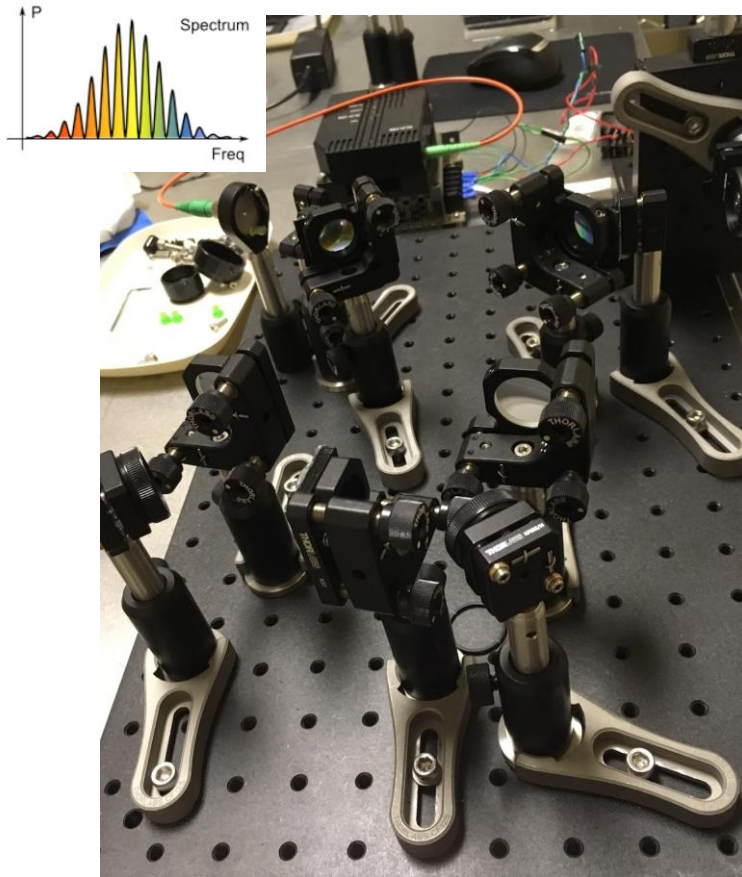


Fizeau Reference:
Diffraction Limited Point

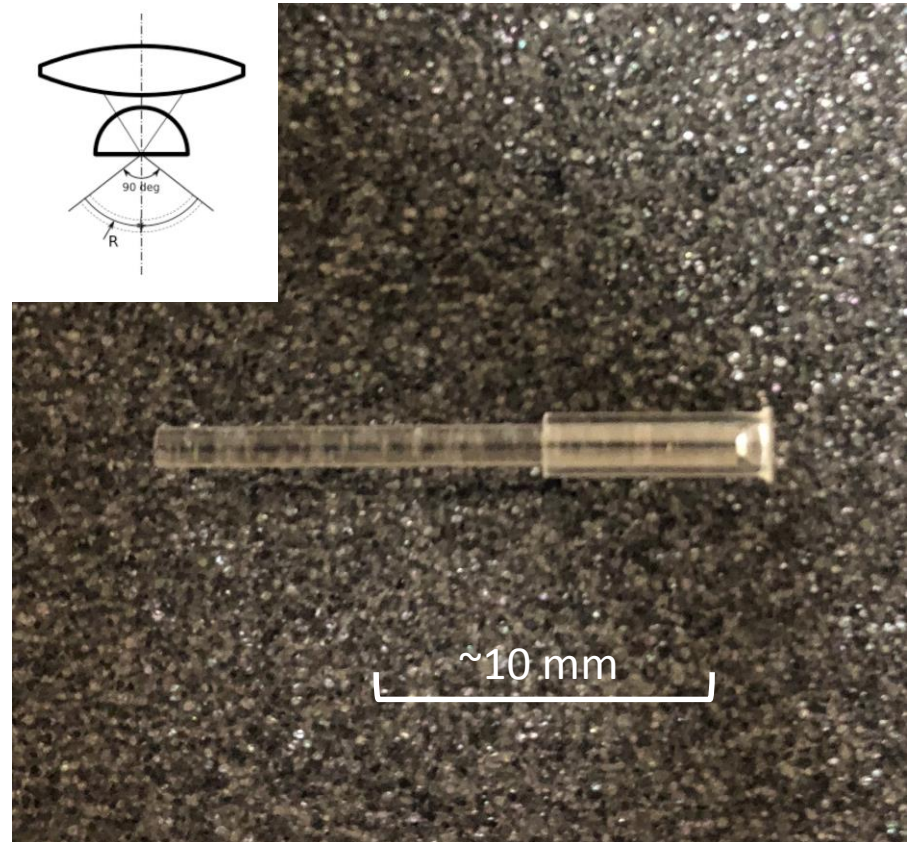
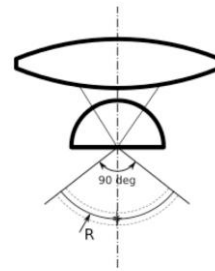
Probe Schematic: Fizeau Interferometer



Project Status: November 2018 (90% through Phase II)



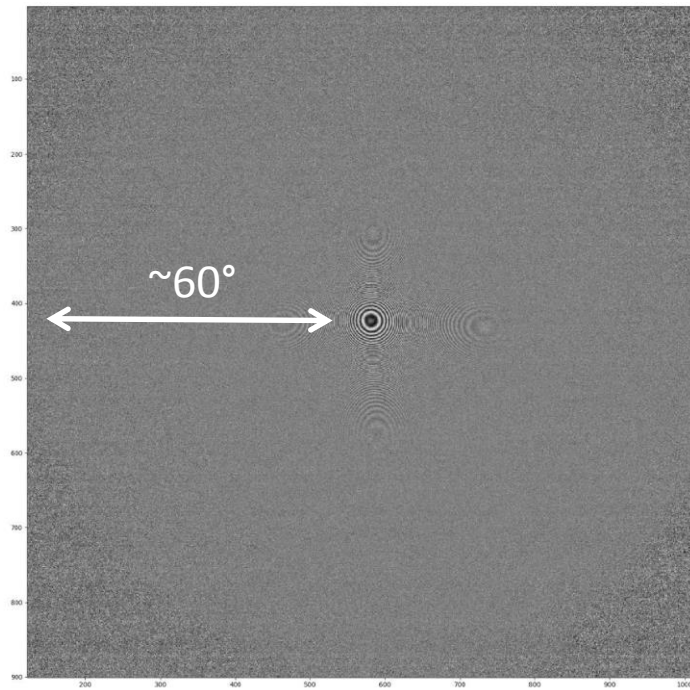
Breadboard SCI Source



Alpha Interferometer Probe Tip

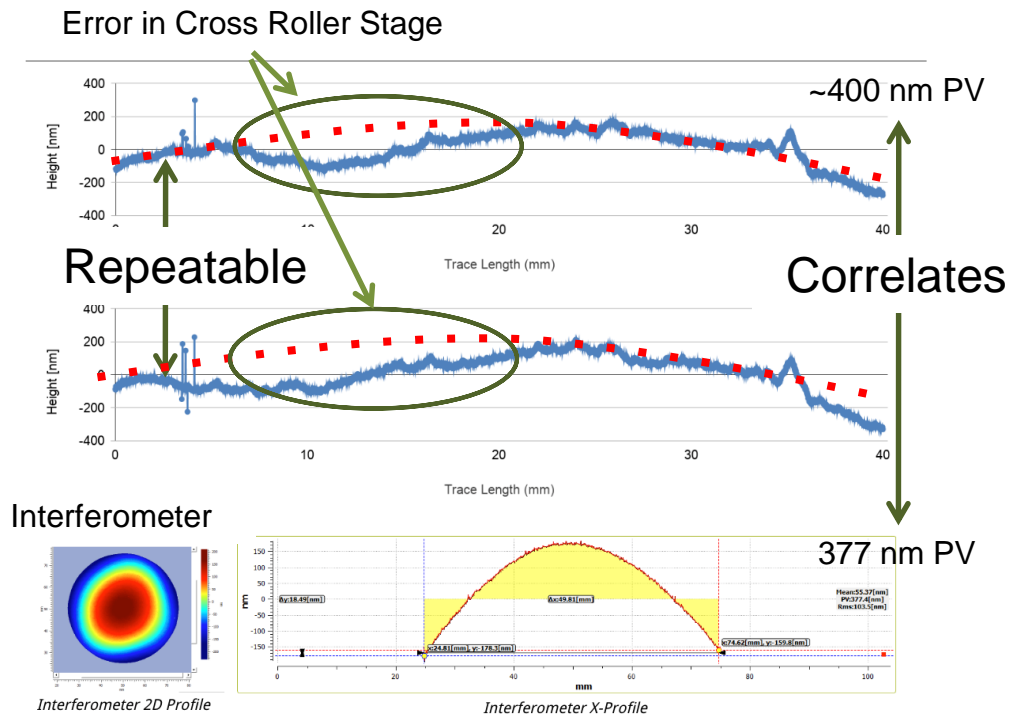
Probe: Repeatable and Correlates to Interferometer

Probe Detecting Flat Surface



400 Hz Data Rate Possible with Present Detectors

Measured Flat Surface 1D Profile Compares with Interferometer (Tilt/Power Removed)



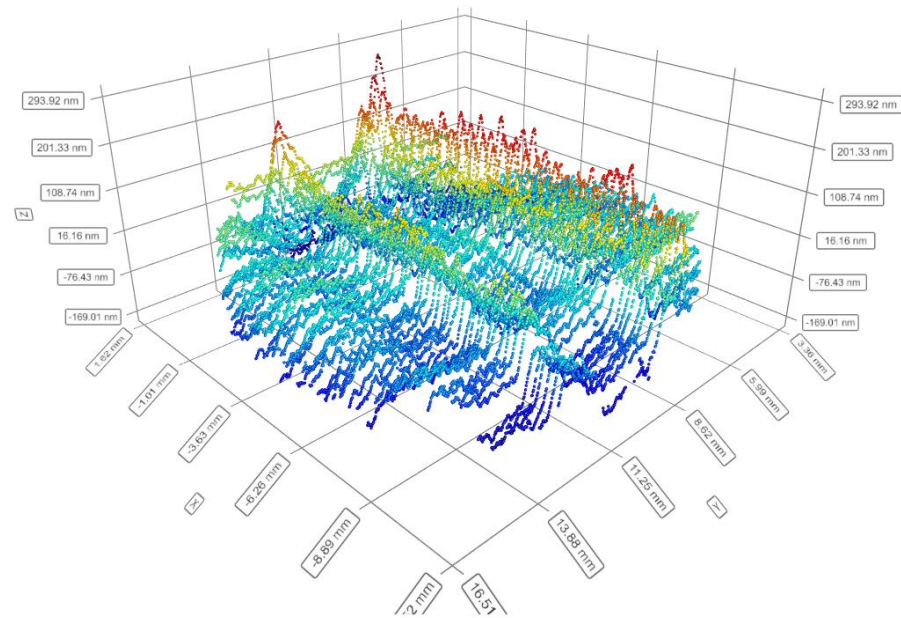
Programming a 2D surface scan now

Next Steps: NASA SBIR

- Construct & measure with α -Probe



Low accuracy XYZ
Metrology Frame

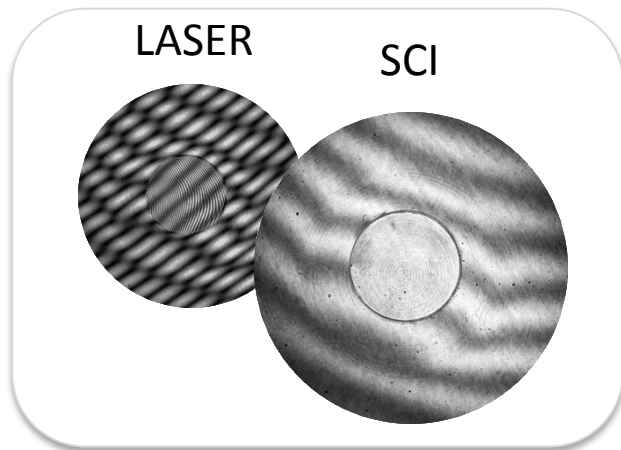


2D Profile Probe Data

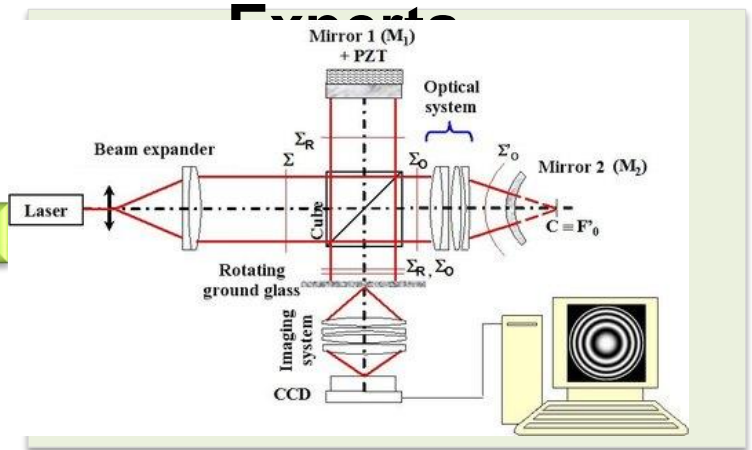
- **Standard Fizeau interferometers + SCI**
 - 500 mm Range
 - Eliminates back surface reflections
 - Isolates surfaces 50 μm
 - Electronically controlled
 - Position
 - Coherence
 - Phase
 - Applications:
 - Etalons, Windows, Prisms down to 50 μm thin
- **Interest in the Probe system is growing**

Interferometers to Advance Your Optical Manufacturing

Technology Creation



Interferometry

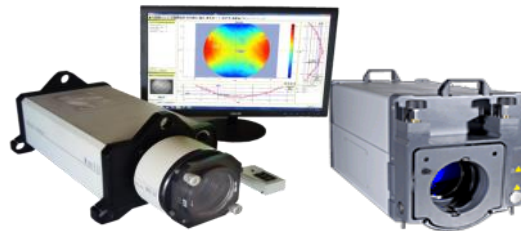
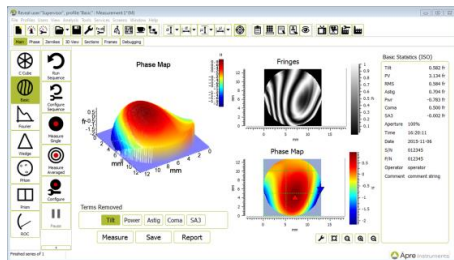


Software

Upgrade/Restore

Performance

Custom





Thank you

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