

# Recent Progress in MEMS Deformable Mirrors

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Mirror Technology/SBIR/STTR Workshop  
Northrop-Grumman Aerospace Systems Presentation Center  
Redondo Beach, California, United States  
16 November 2017





# Outline

- BMC DM Technology
- NASA funded mirror technology programs
- Space astronomy operations
- Ground astronomy operation

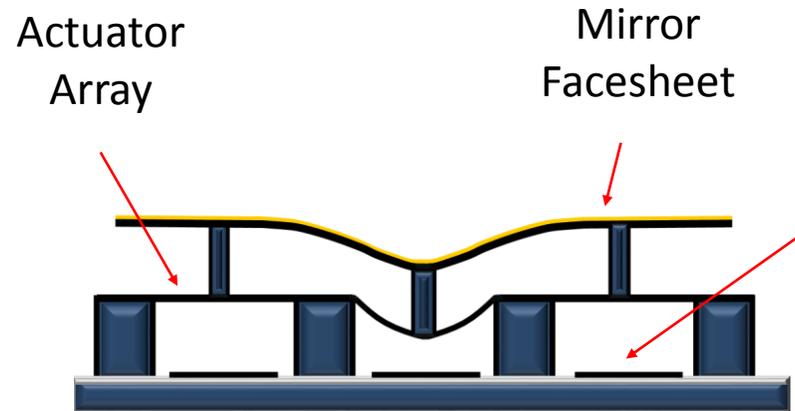


# Outline

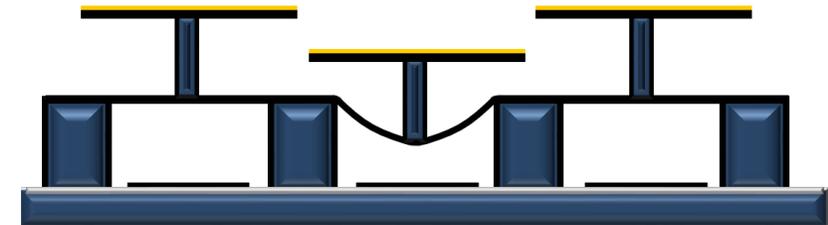
- **BMC DM Technology**
- NASA funded mirror technology programs
- Space astronomy operations
- Ground astronomy operation



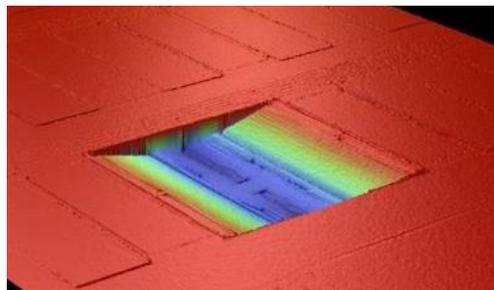
# MEMS DM Architecture



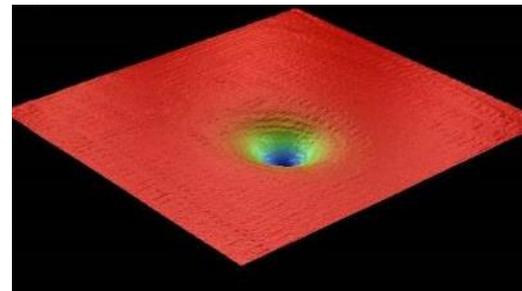
**Continuous mirror  
(smooth phase control)**



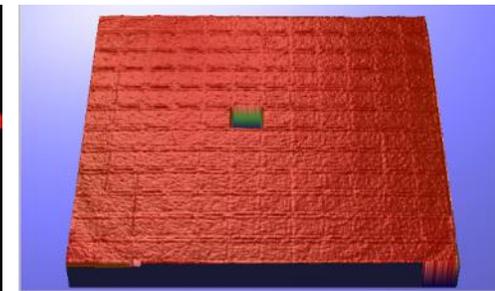
**Segmented mirror  
(uncoupled control)**



**Deflected Actuator**



**Deformed Mirror  
Membrane**



**Deformed  
Segmented Mirror**

# BMC Mirror Family



## Small Cartesian Arrays

- Square arrays from 32 to 140 actuators
- Strokes: 1.5 $\mu$ m, 3.5 $\mu$ m or 5.5 $\mu$ m

## Medium Cartesian Arrays

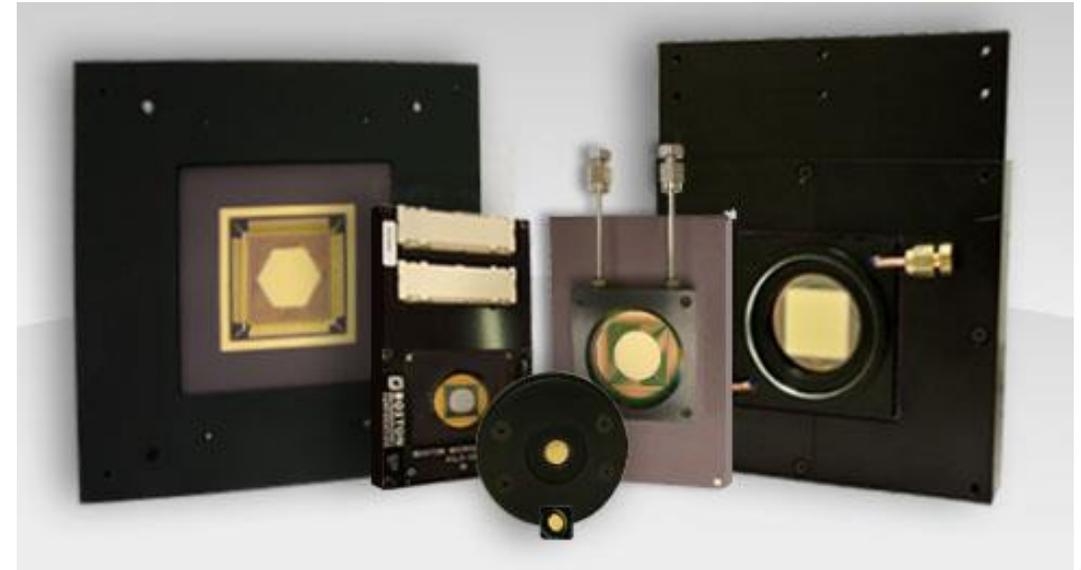
- Square and circular arrays from 492 to **1020**
- 1.5 $\mu$ m & 3.5 $\mu$ m stroke

## Large Cartesian Arrays

- Square and circular arrays from **2040** to 4092
- 1.5 $\mu$ m and 3.5 $\mu$ m stroke

## Hex Tip-Tilt-Piston

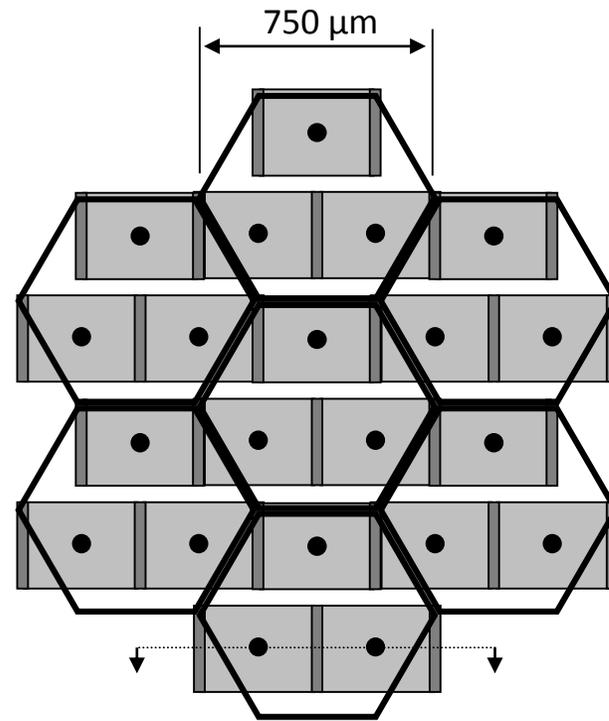
- 37, **331- and 1021-Segment Devices**



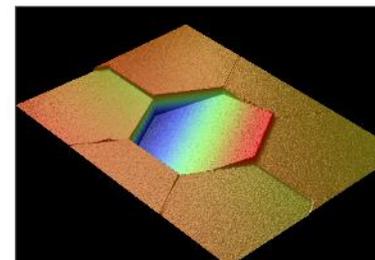
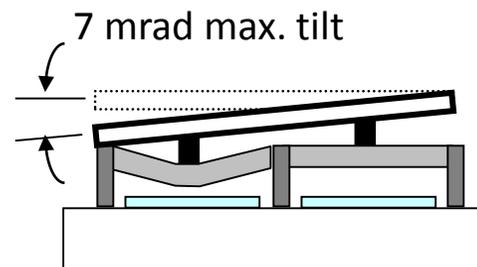
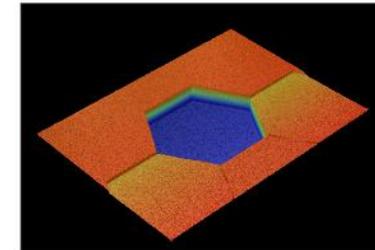
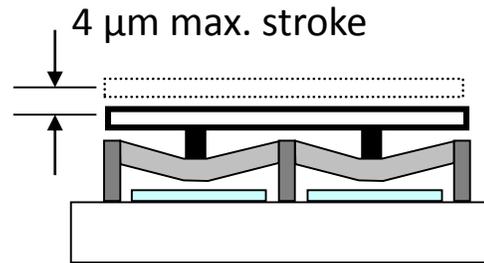
Developed through NASA funding



# Hex-Close Pack Deformable Mirror Tip-Tilt-Piston

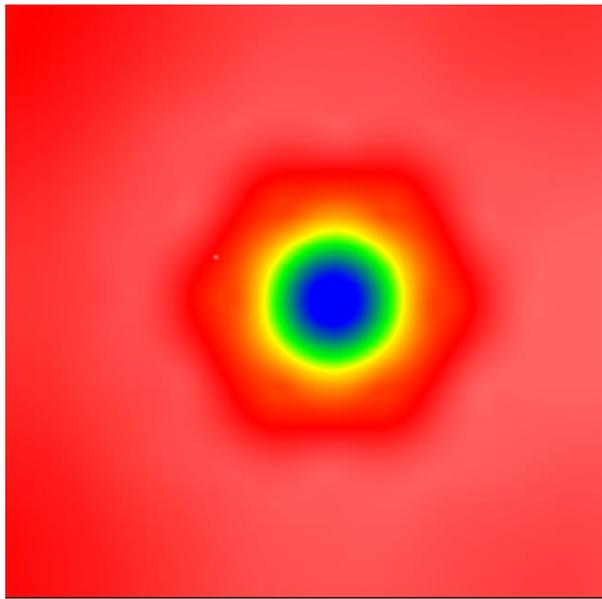


- Up to 3063 actuators
- Independent hexagonal segments
  - 3 actuators per segment
- 4  $\mu\text{m}$  max. stroke
- 7 mrad max. tilt angle

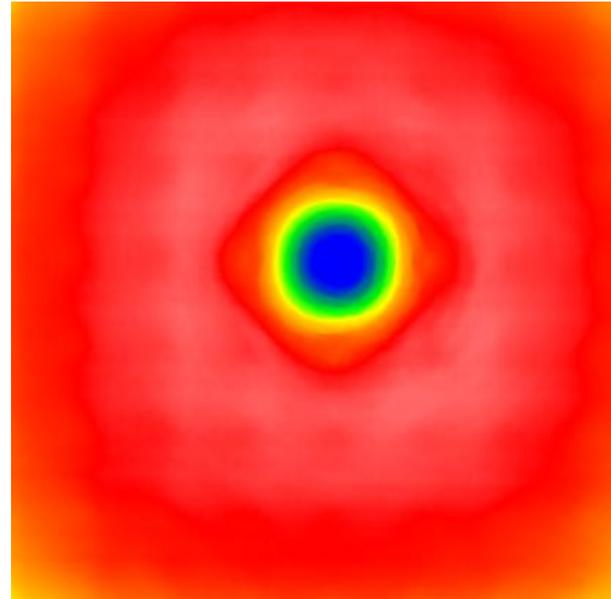




# Continuous Hex-Close Pack DM

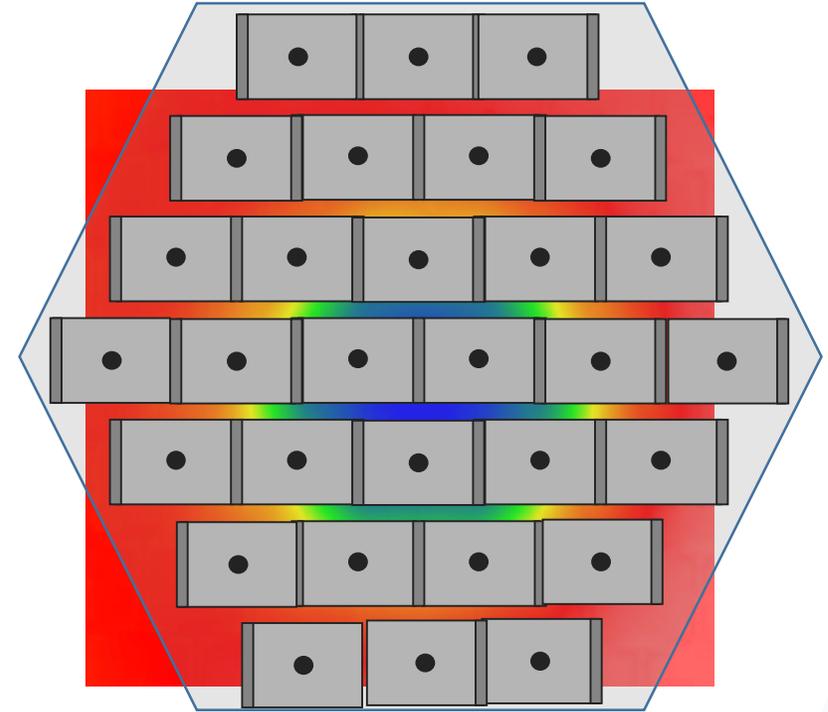


Hex



Cartesian

Same actuator design, 3.5 $\mu$ m stroke device



7 actuator "circle"



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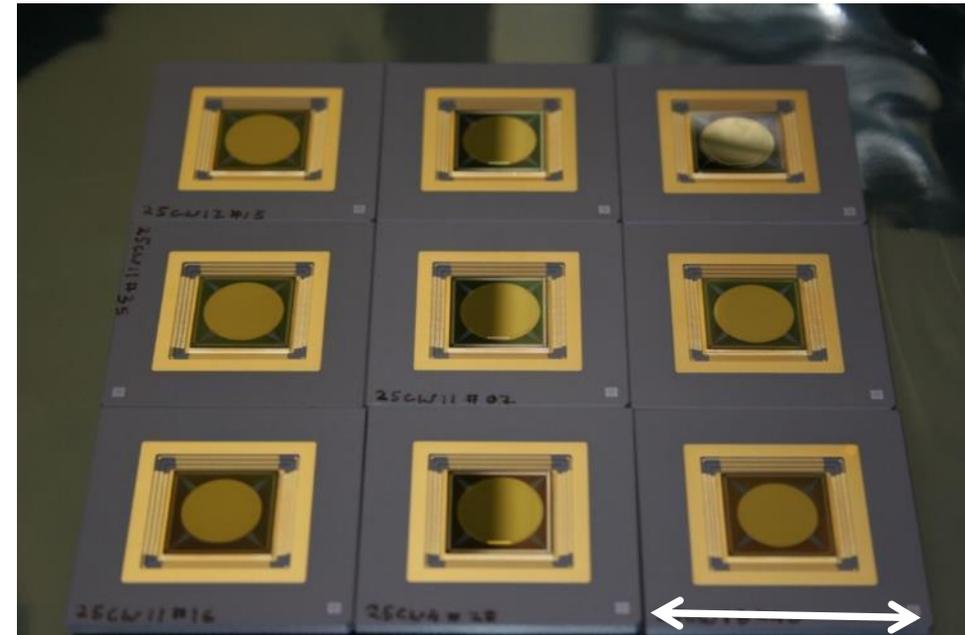


# *MEMS Deformable Mirror Technology Development for Space-Based Exoplanet Detection*

Contract#: NNH12CQ27CSAT/TDEM

Objective: Demonstrate survivability of the BMC MEMS Deformable Mirror after exposure to dynamic mechanical environments close to those expected in space based coronagraph launch.

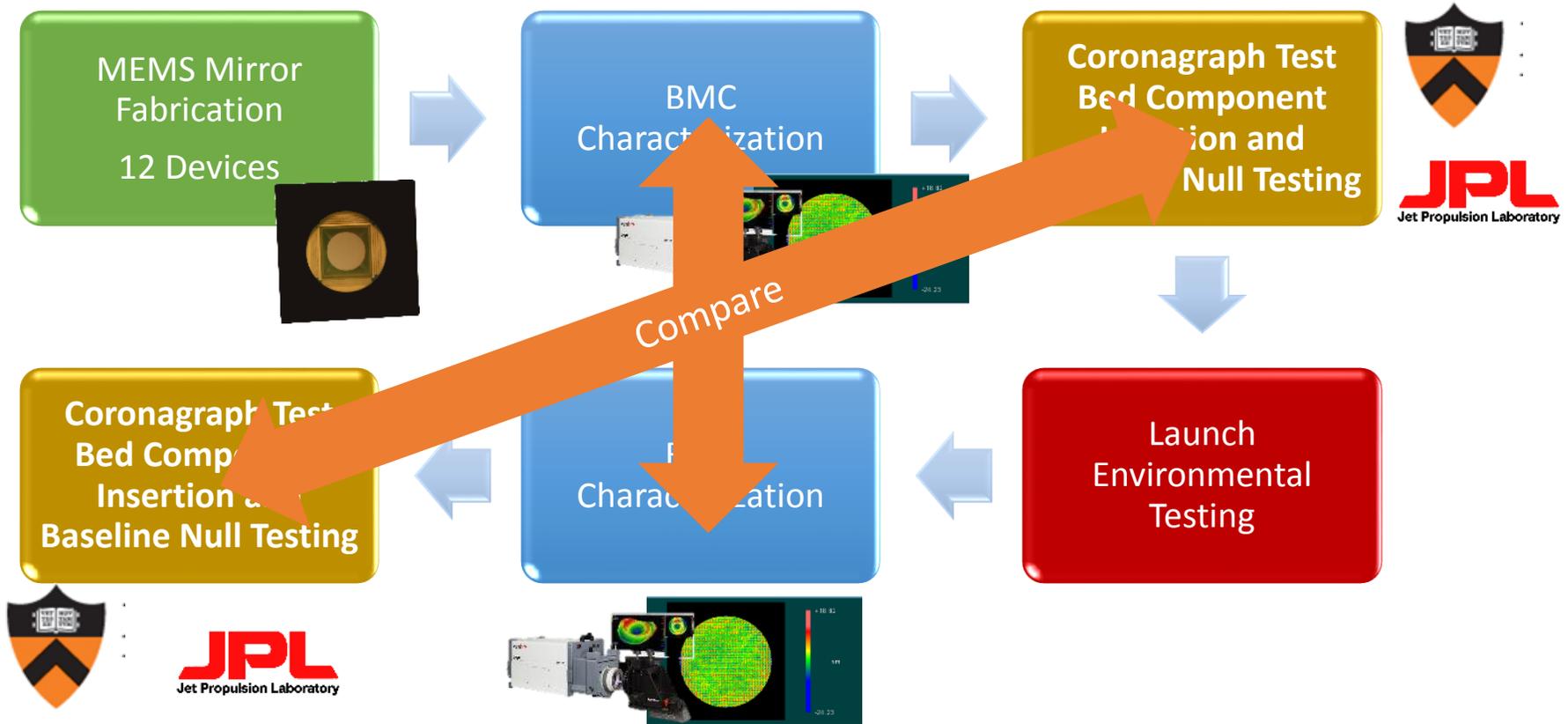
9 Mirrors ready for testing



5cm



# Project Flow



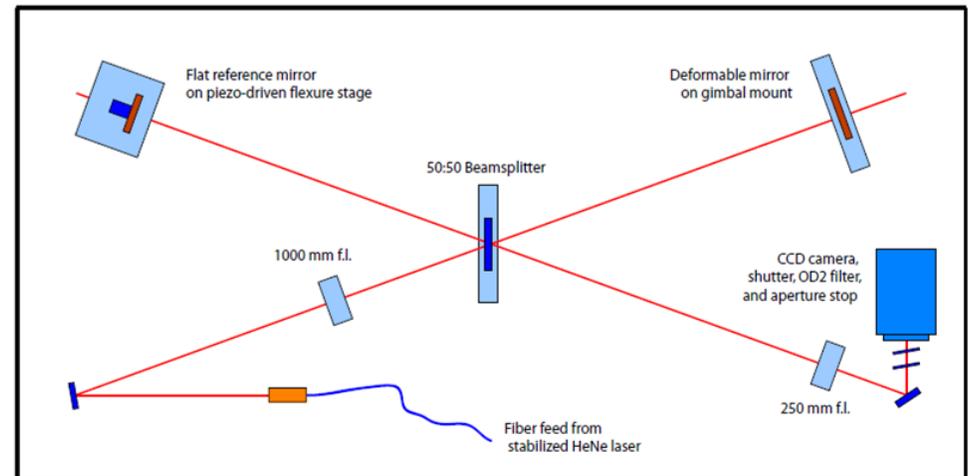


# Vacuum Surface Gauge (VSG) Measurements

Two 952 actuator MEMS DMs (tested separately)

- Surface figure of DM at zero bias
- Surface figure of DM for flat surface
- Actuator gains for all 952 actuators for small up/down pokes about the flat surface condition
- Drift in surface for “flat” condition for 48 hour period
- Repeatability from “flat” and BMC/JPL solution for 10 repeats

Work performed by: Frank Greer, Cory Hill, Brian Gordon, John Trauger



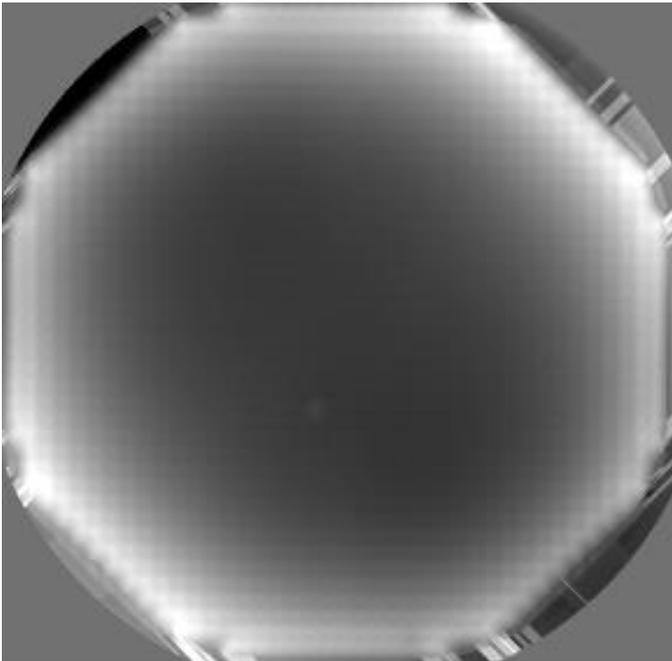
VSG is mounted on a 36 x 72 inch optical table. End-points of axes are the threaded holes 4.5 x 6.5 inches from table corners. Beam height = 4.405 inches.

- VSG is a Michelson interferometer mounted in a vibration isolated vacuum chamber
- Light source is 632.8 nm frequency HeNe laser
- Reference mirror is mounted on a piezo-driven flexure translation stage
- Deformable mirror under test is on a gimbal mount with a temperature controlled stage



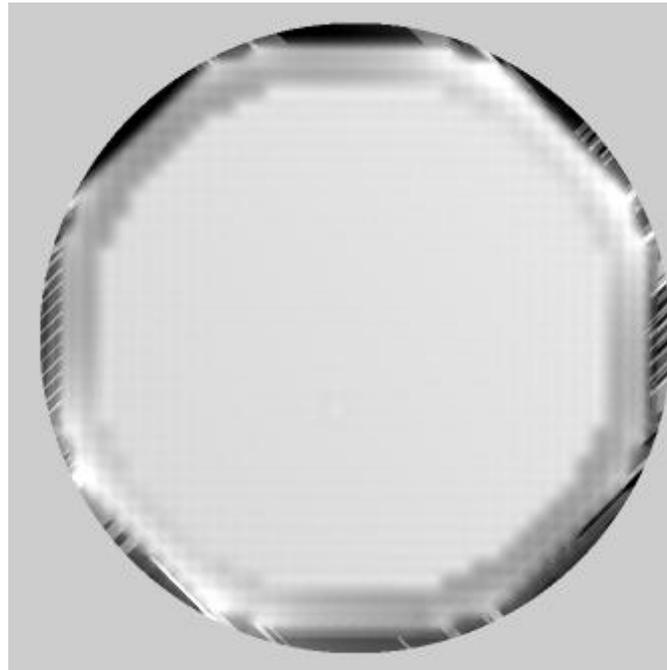
# Flattened DM

Unpowered mirror



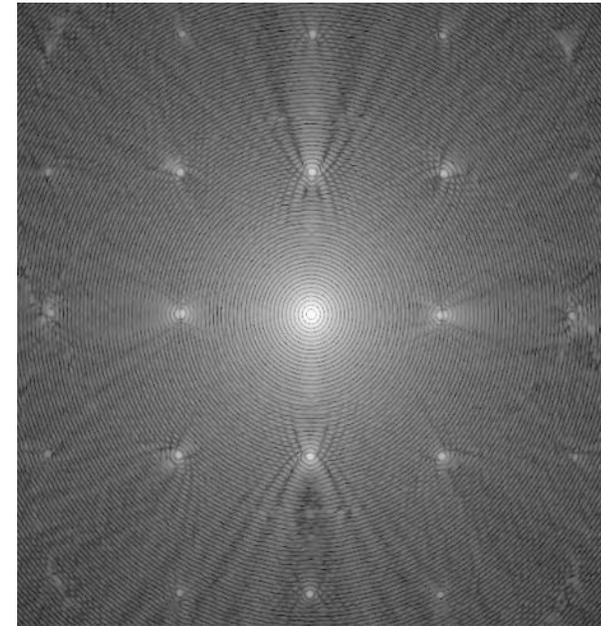
PV Focus 1428 nm, PV 45°  
astig 360 nm,  
PV 90° astig 3 nm, RMS in  
higher 110 nm

Flattened DM



6.6nm PV focus,  
2.9nm PV 45 deg astig,  
0.3nm PV 90 deg astig,  
7.6nm RMS higher  
order terms

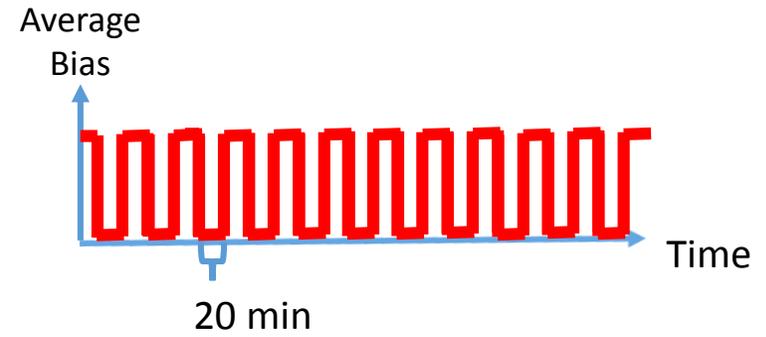
Point Spread Function Flattened DM



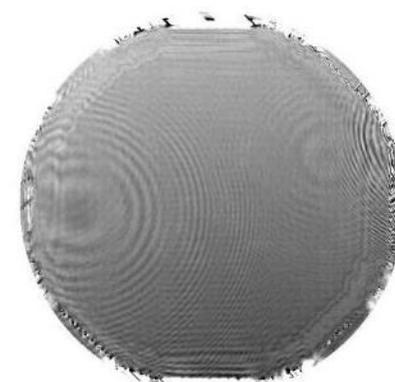
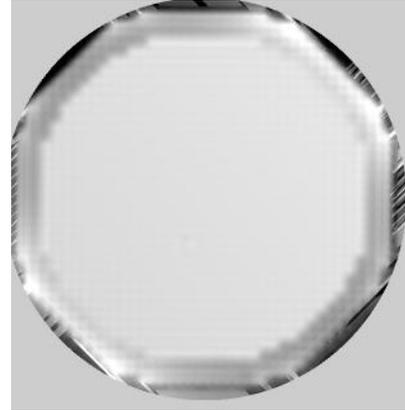
PSF with best flat BMC  
RMS= 7.6 nm RMS)  
Strehl at  $\lambda = 633 \text{ nm} : 0.96$



# Repeatability Test Results

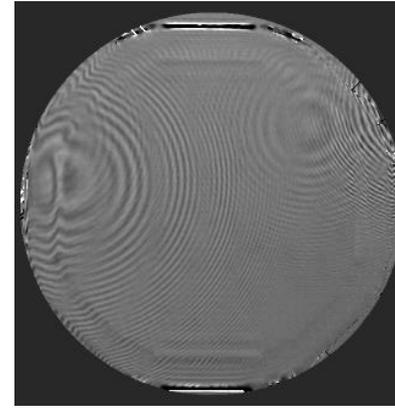
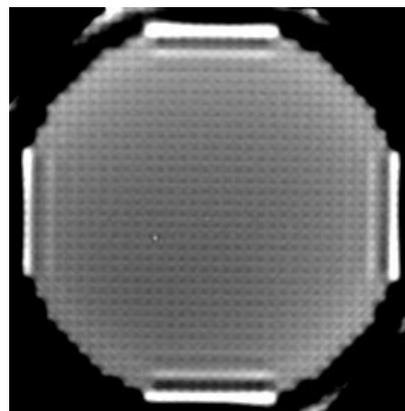
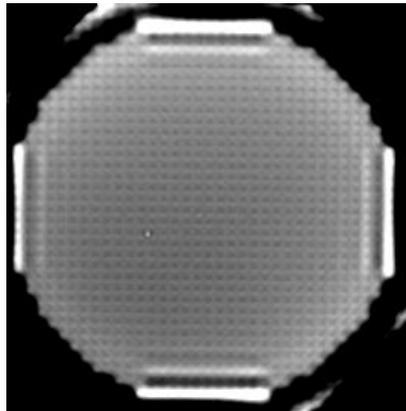


DM1



1.4nm RMS difference

DM2



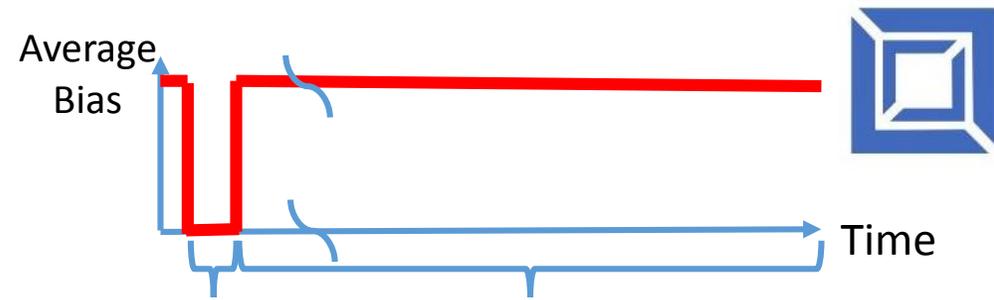
1.4nm RMS difference

Initial "Flat" solution

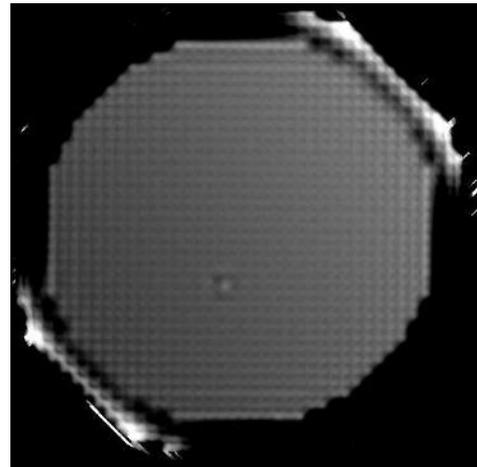
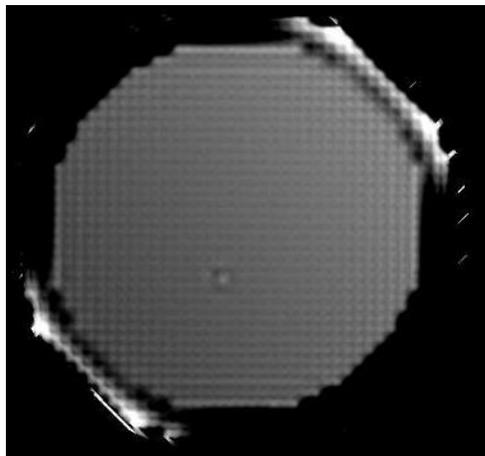
"Flat" solution 10 iterations  
of zero volts to "flat"

No obvious differences between maps.

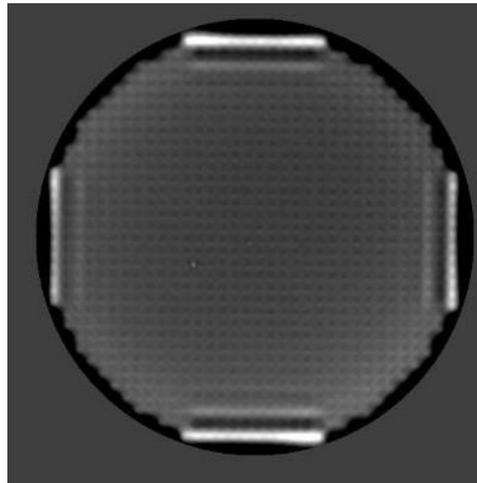
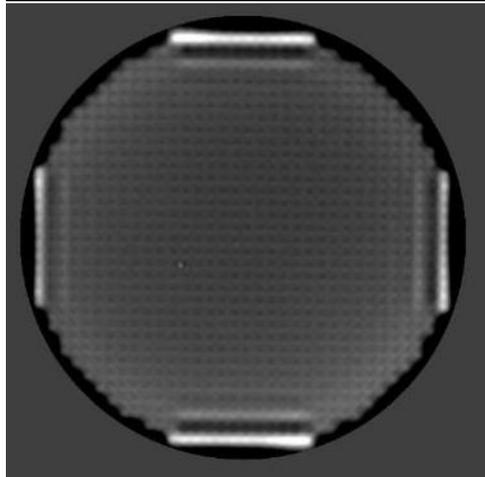
# Comparison of flats from settling test



DM1



DM2



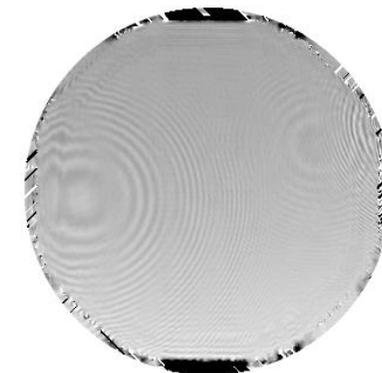
**First map**

**Last map**

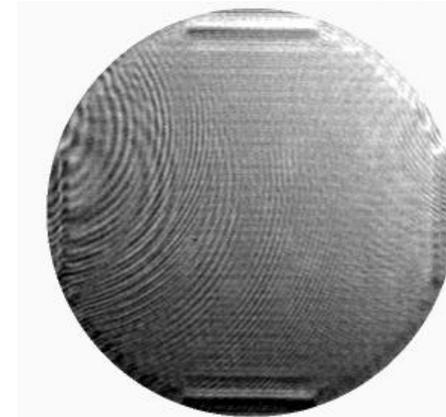
No significant changes observed in the settling time test (48hrs)

20 min

48 hours



2.1nm RMS difference



0.9nm RMS difference



# High Contrast Imaging Laboratory(HCIL)

Kasdin Lab, Princeton University

Focal Plane Wavefront Correction (FPWC) for Exoplanet Coronagraph Imaging

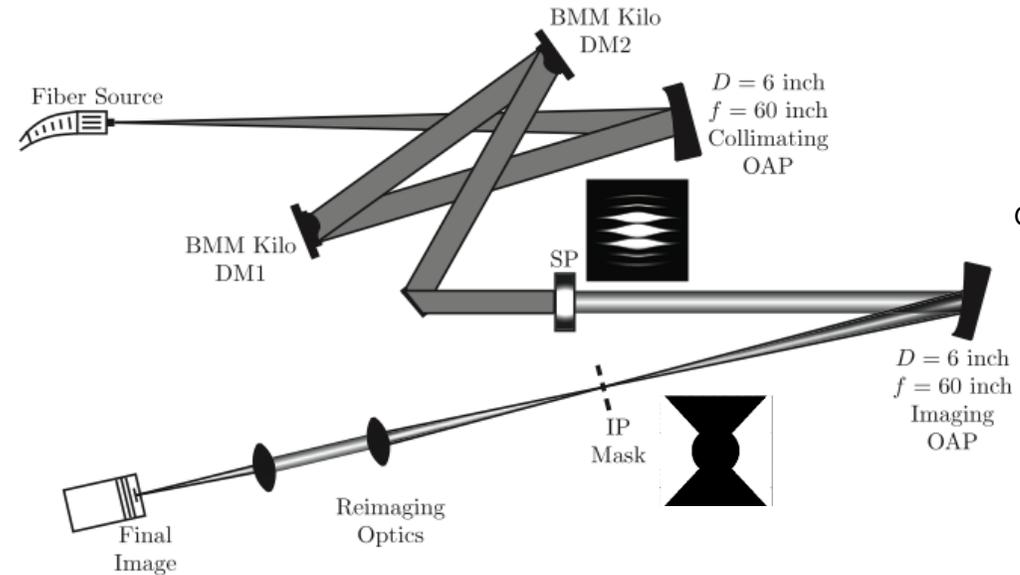
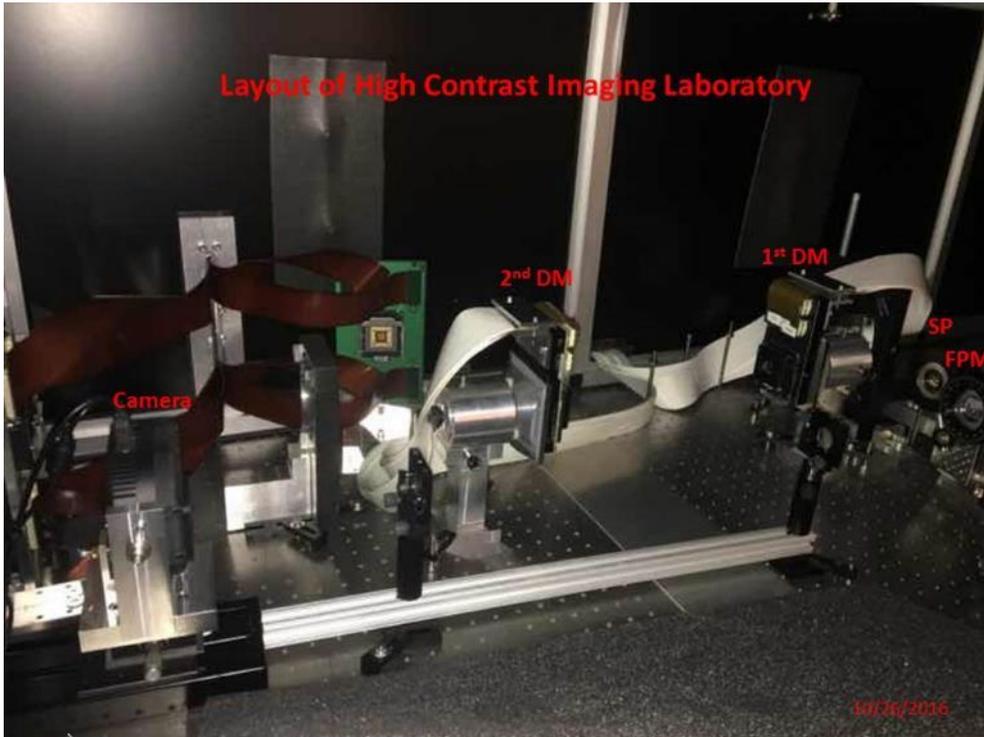


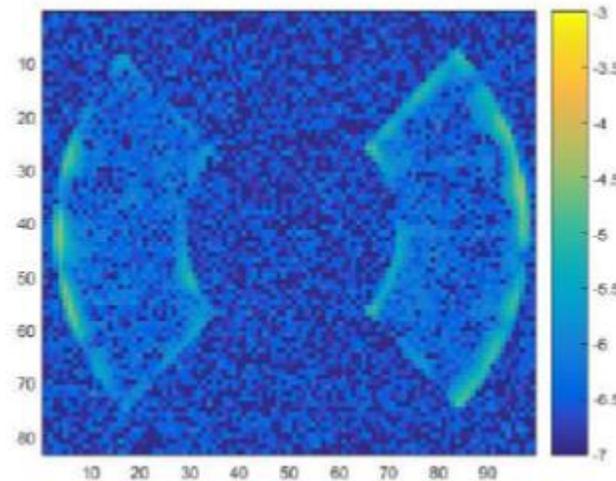
Image Credit:  
Groff & Kasdin 2013

- **Shaped pupil coronagraph** technique is used to achieve high contrast for exoplanet direct imaging.
- 2 BMC **deformable mirrors** are included to compensate optical aberrations in the system.

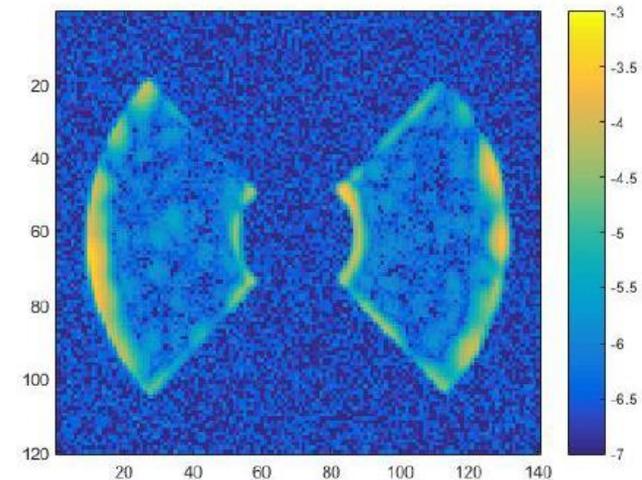


# Lab Results

- Batch process estimator with two pairs of probes
- Stroke minimization controller
- Two BMC DMs with 952 actuators on each
- Achieved  $2 \times 10^{-7}$  contrast within 6-11  $\lambda/D$  and  $9 \times 10^{-7}$  contrast 5-14  $\lambda/D$



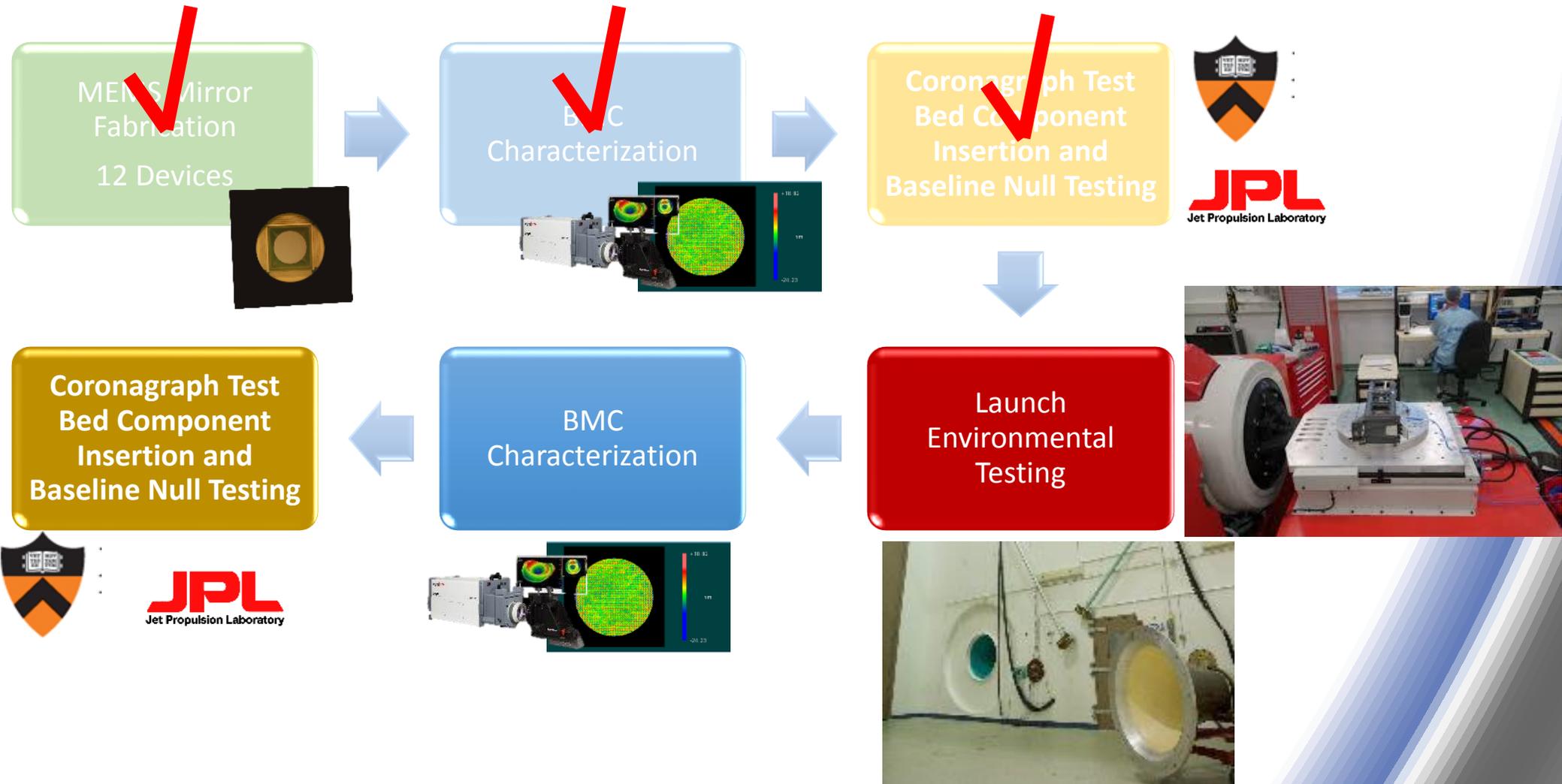
6-11  $\lambda/D$



5-14  $\lambda/D$



# Project Flow

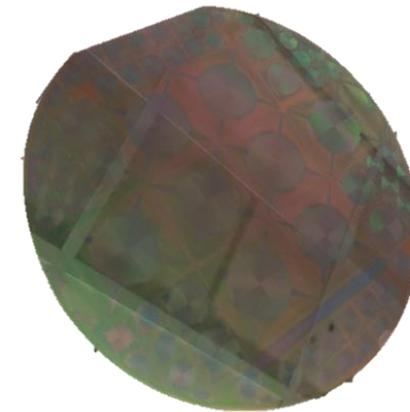
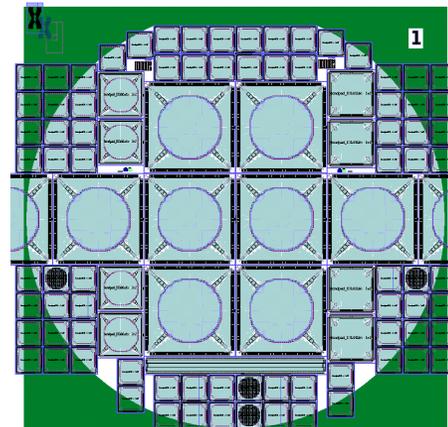
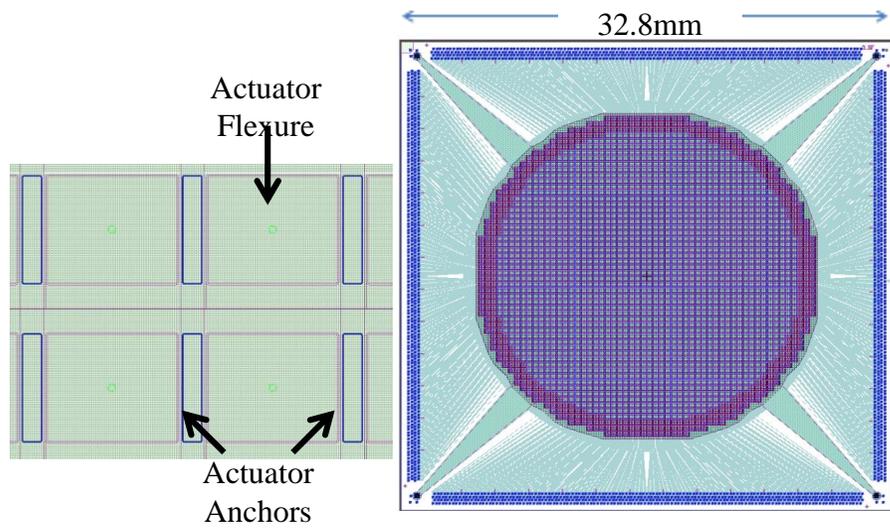




# Improved Yield, Performance and Reliability of High-Actuator-Count Deformable Mirrors

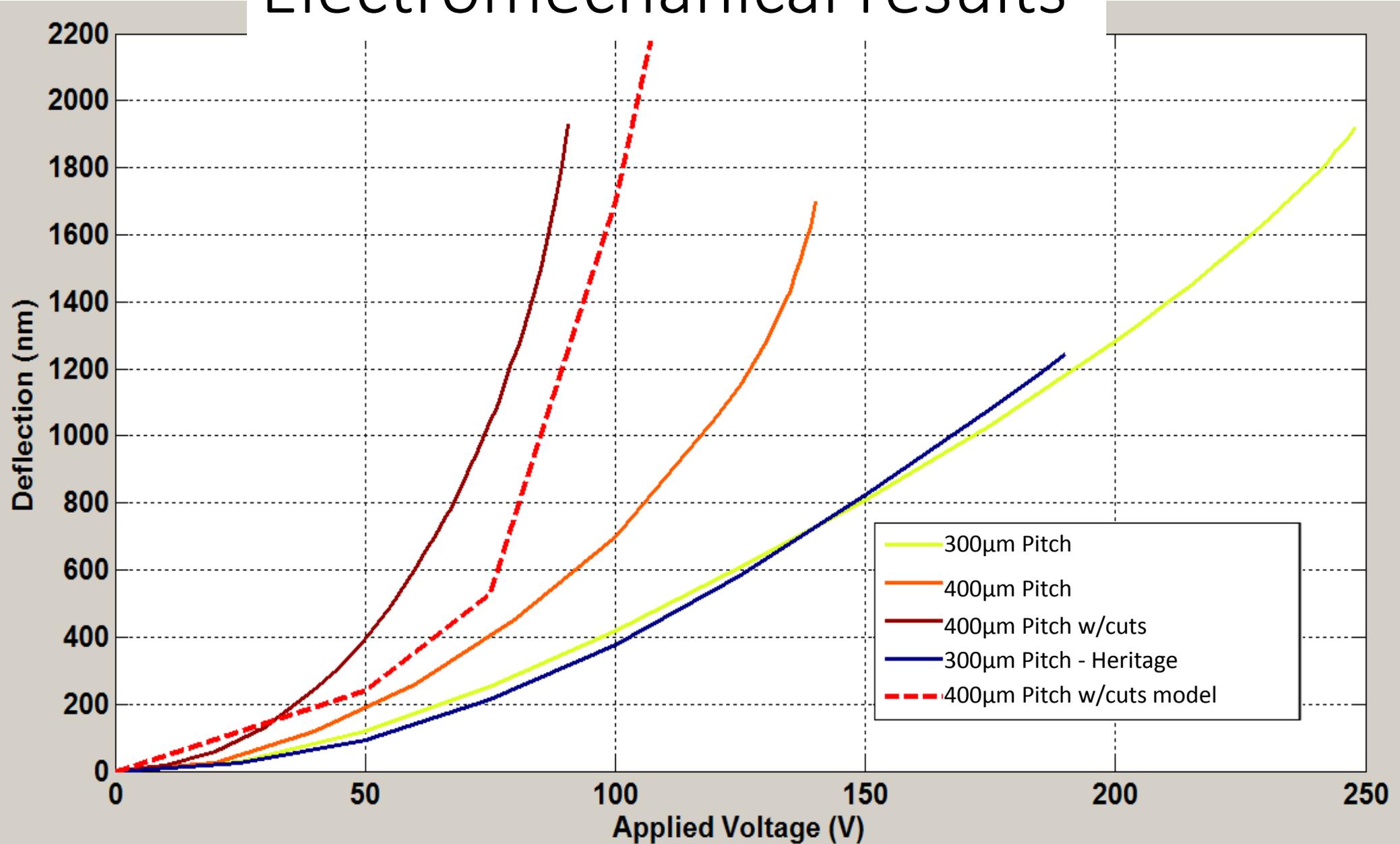
Contract Number: NNX16CP14C Phase II SBIR

Mirror architecture	2040 actuators
Active Aperture Diameter	19.6mm
# Actuators across active diameter	50
Actuator Pitch	400 $\mu$ m
Actuator Stroke	1.5 $\mu$ m
Operating Voltage	0-100V
Mirror Surface Figure	<5nm RMS



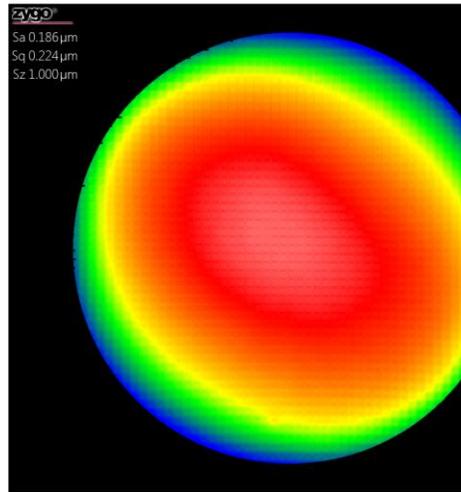
Design of actuators for lower voltage operation

# Electromechanical results





# Initial Die Inspection



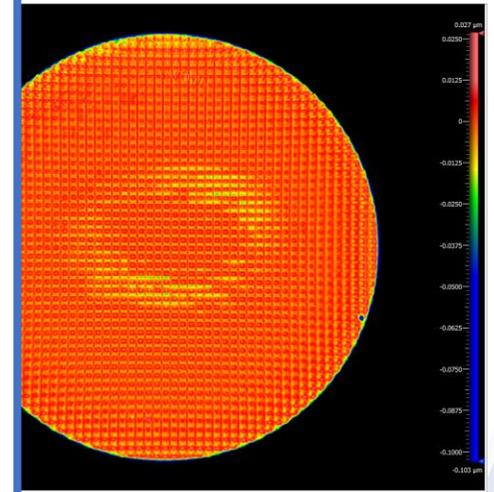
← 20mm →

Unpowered  
P-V 900nm  
RMS 200nm

Currently coating  
and packaging die.  
Project completion  
April 2018.

Delivery – 2k DM.

RMS 84nm

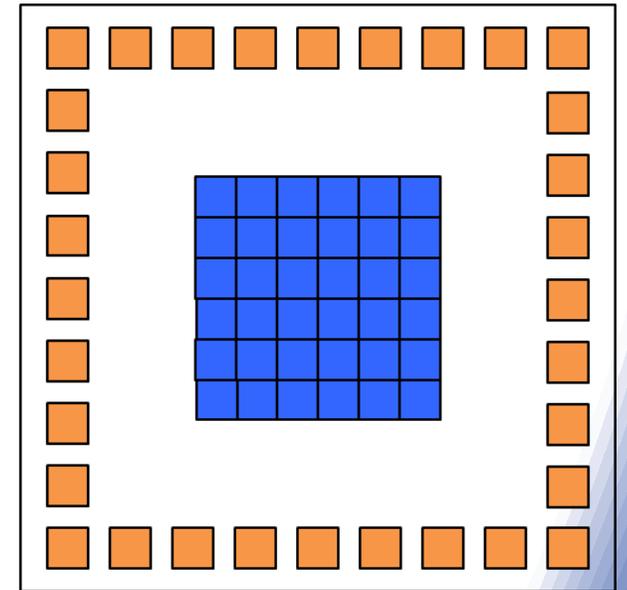
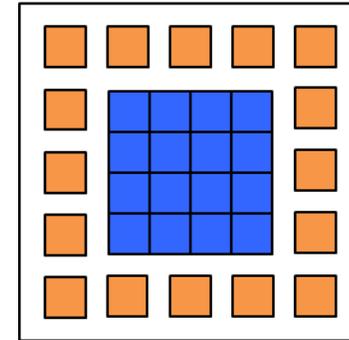


Pass Filter ( $\lambda = 2.5/\text{mm}$ )  
33nm  
7nm



# Need for Even Higher Actuator Count DM (10k +)

- For many next generation instruments(ground and space based), more actuators are needed
- Limited by electrical interconnects
  - Wirebond for each actuator
  - Span of active optical surface scales with  $N$
  - Span of the chip scales with  $N^2$
  - Limits number of die on a wafer
  - Increases the likely hood of a single point defect causing short/failure

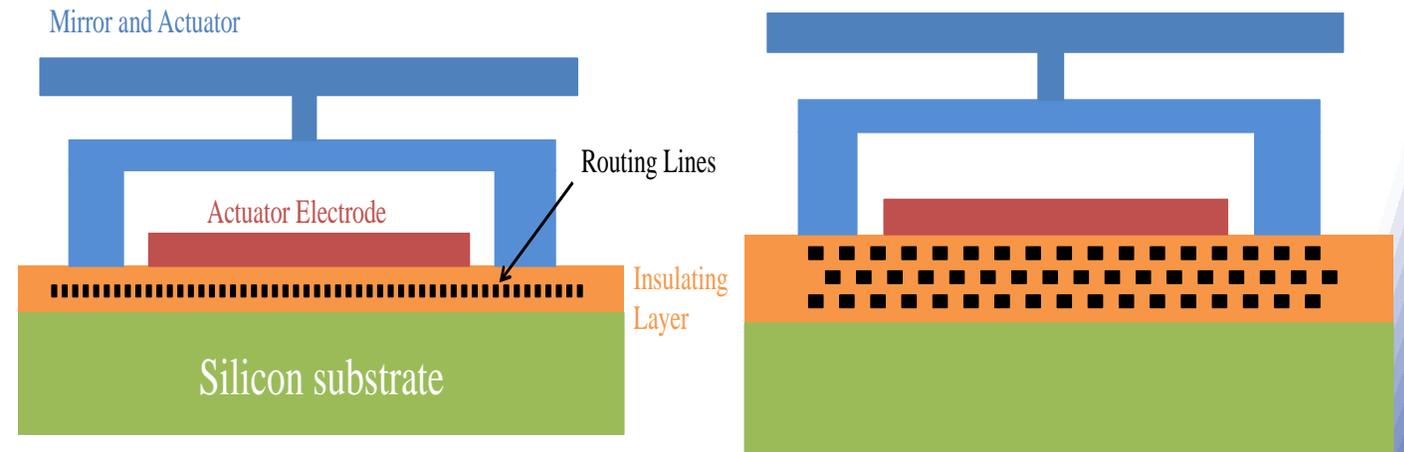
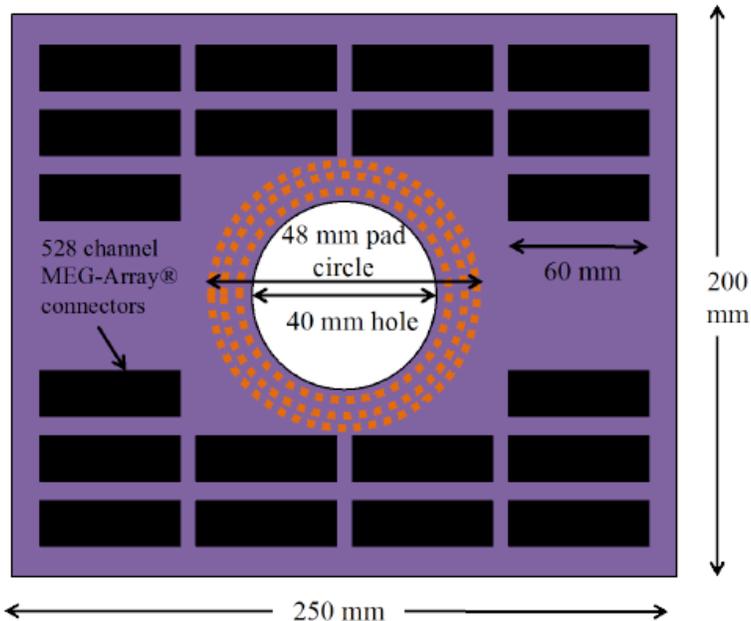
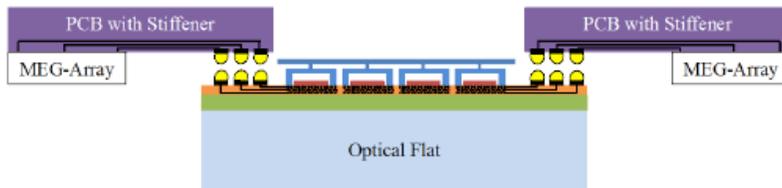


By adding 2 more actuators across the aperture, the die size increased by  $\sim 3x$

# Technology Development for High-Actuator-Count MEMS DM Systems



- NASA Contract #NNX17CP76P Phase I SBIR



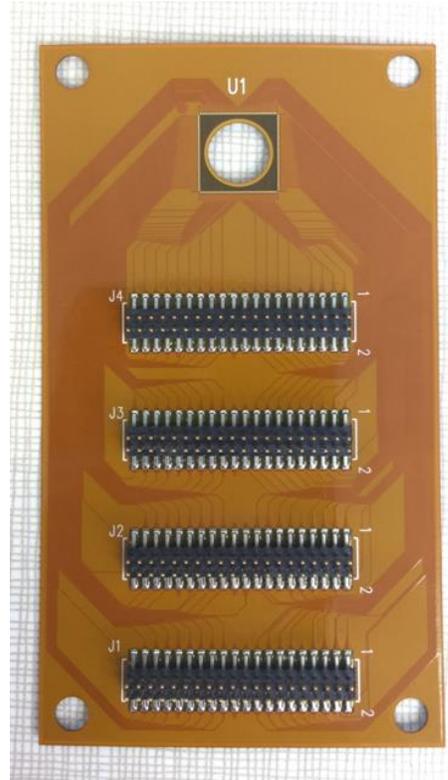
## Goals:

- Demonstrate flip-chip bonding approach on smaller actuator count devices
- Layout 100 Actuator-Across –Aperture DM with multiple routing lines

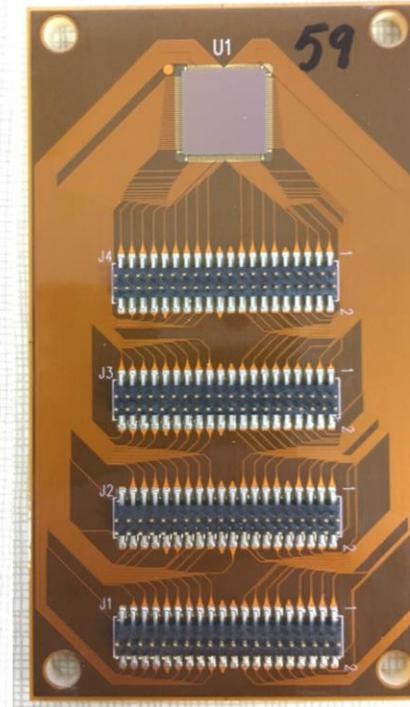


# Demonstration of Bonding

- Flex circuit PCB for high density bond pads
- Used heritage 140- actuator DM
- Central hole provides optical access for active aperture



Backside  
No Die attached



Backside  
Die attached



Frontside  
Die attached

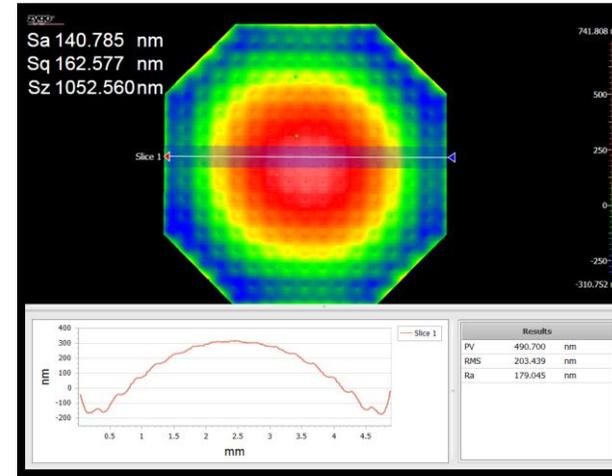
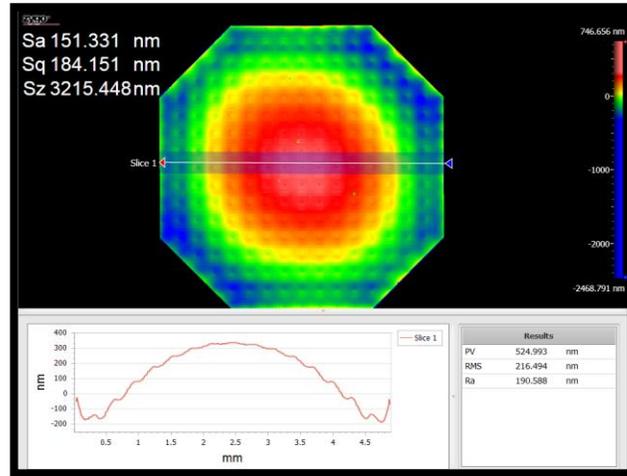
# Compliant Stencil Epoxy Bond Surface figure



Before Bond

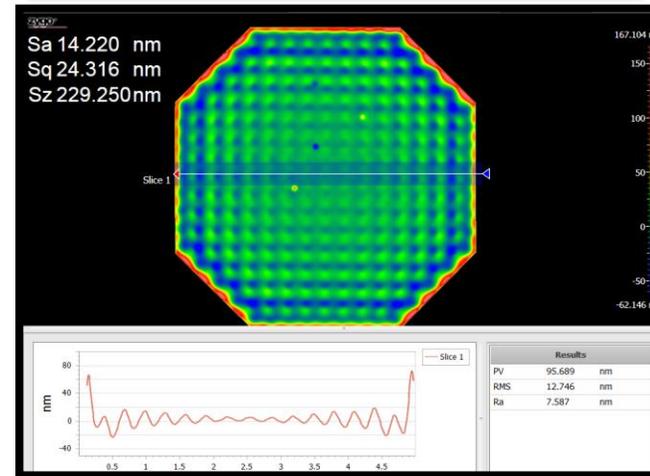
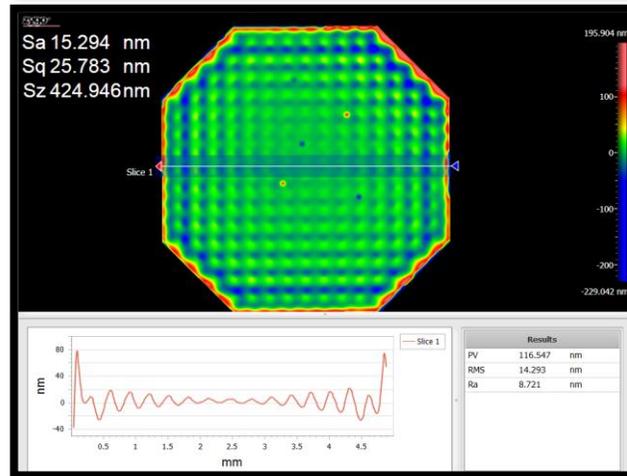
After Bond

Unpowered



Vertical scale  
-200 to 400nm

Unpowered, filtered



-40 to 80nm





Bondpad

Area

## Phase II Plan

Fabricate and Package an 8k DM

- New manufacturing processes
- Bonding on large scale
- Design a large interface board
- December 2017



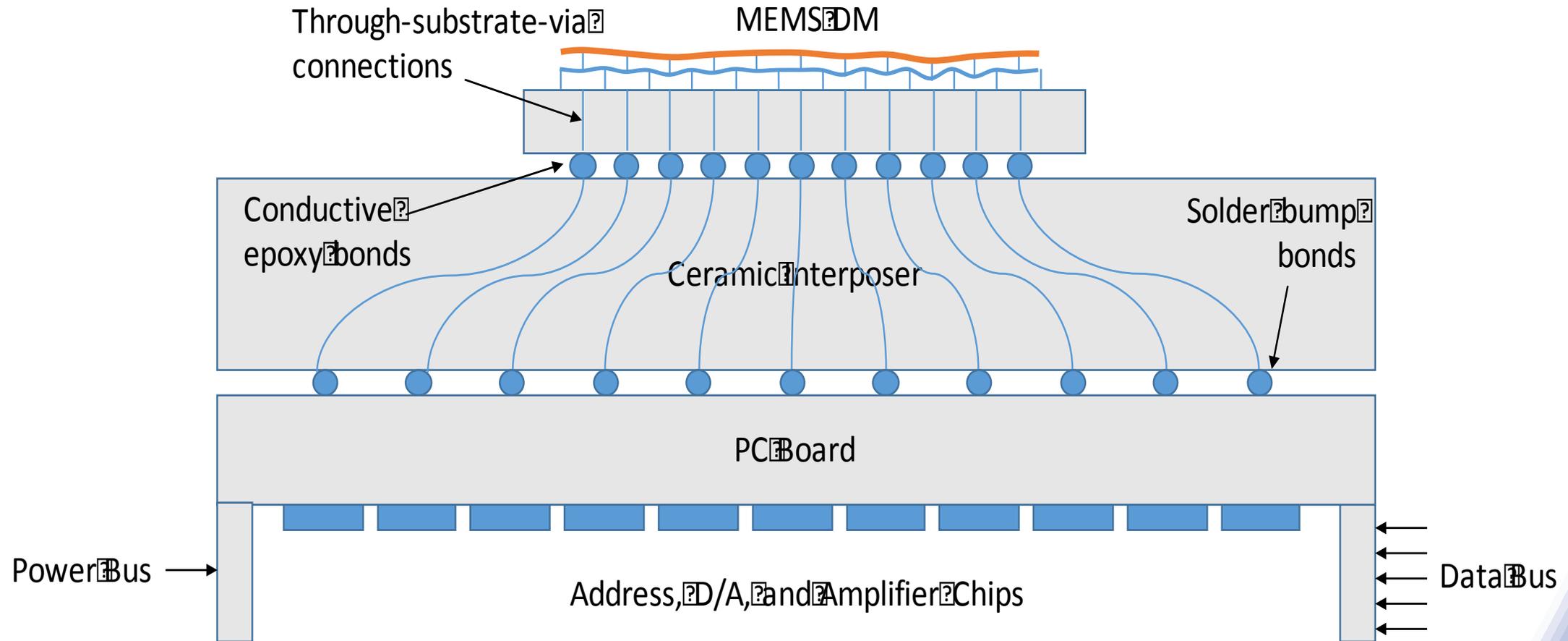
# *Compact, scalable deformable mirror systems for space-based imaging of exo-earths*

NASA Contract: NNX17AI66G APRA Program

- Goals:
  - Demonstrate a **new architecture and integration approach** for compact, robust large-format deformable mirror systems, and
  - Show **a feasible path for scaling up** that demonstration platform and manufacturing integration approach to larger formats with up to 10,000 actuators

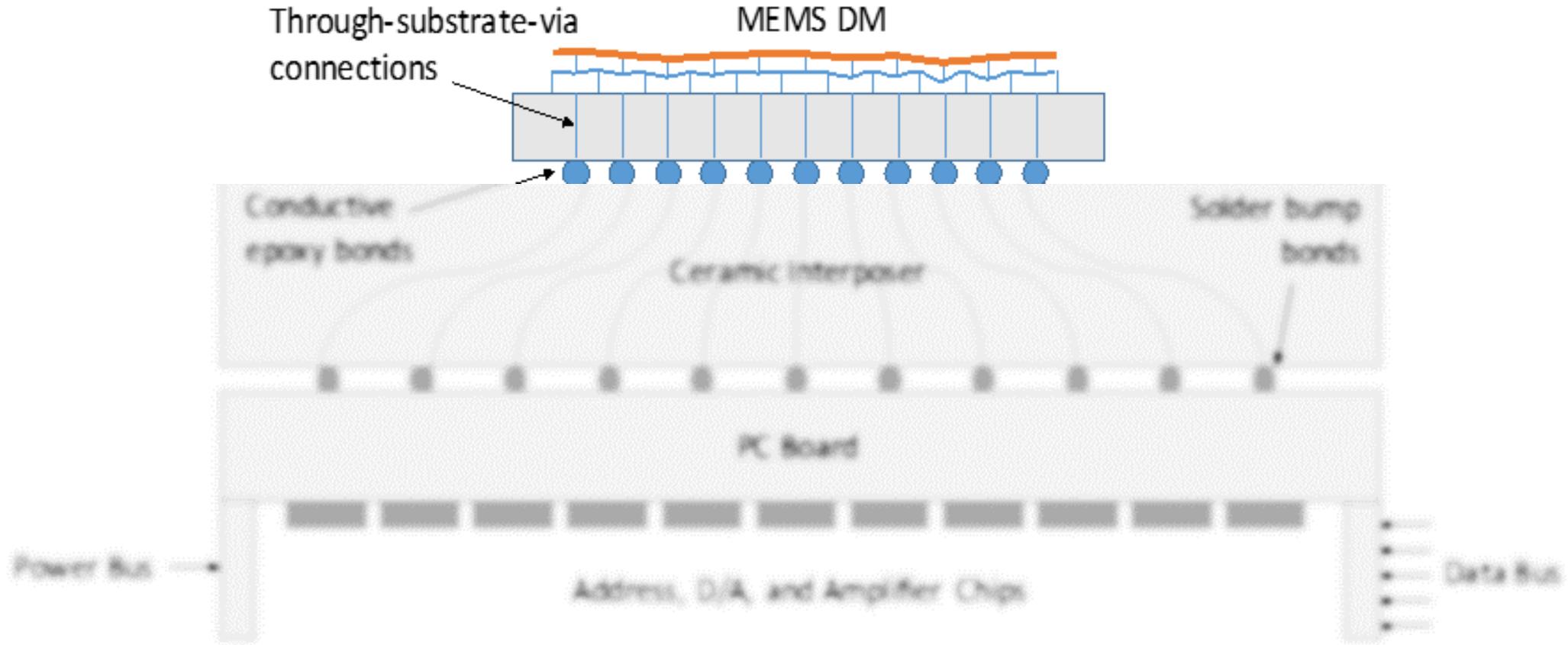


# Proposed Architecture





# Proposed Architecture





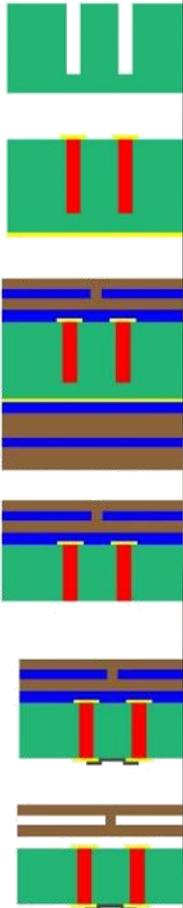
# Through-Wafer Via DM Prototype

(manufacturing) relies  
on design, but  
uses through-wafer-via

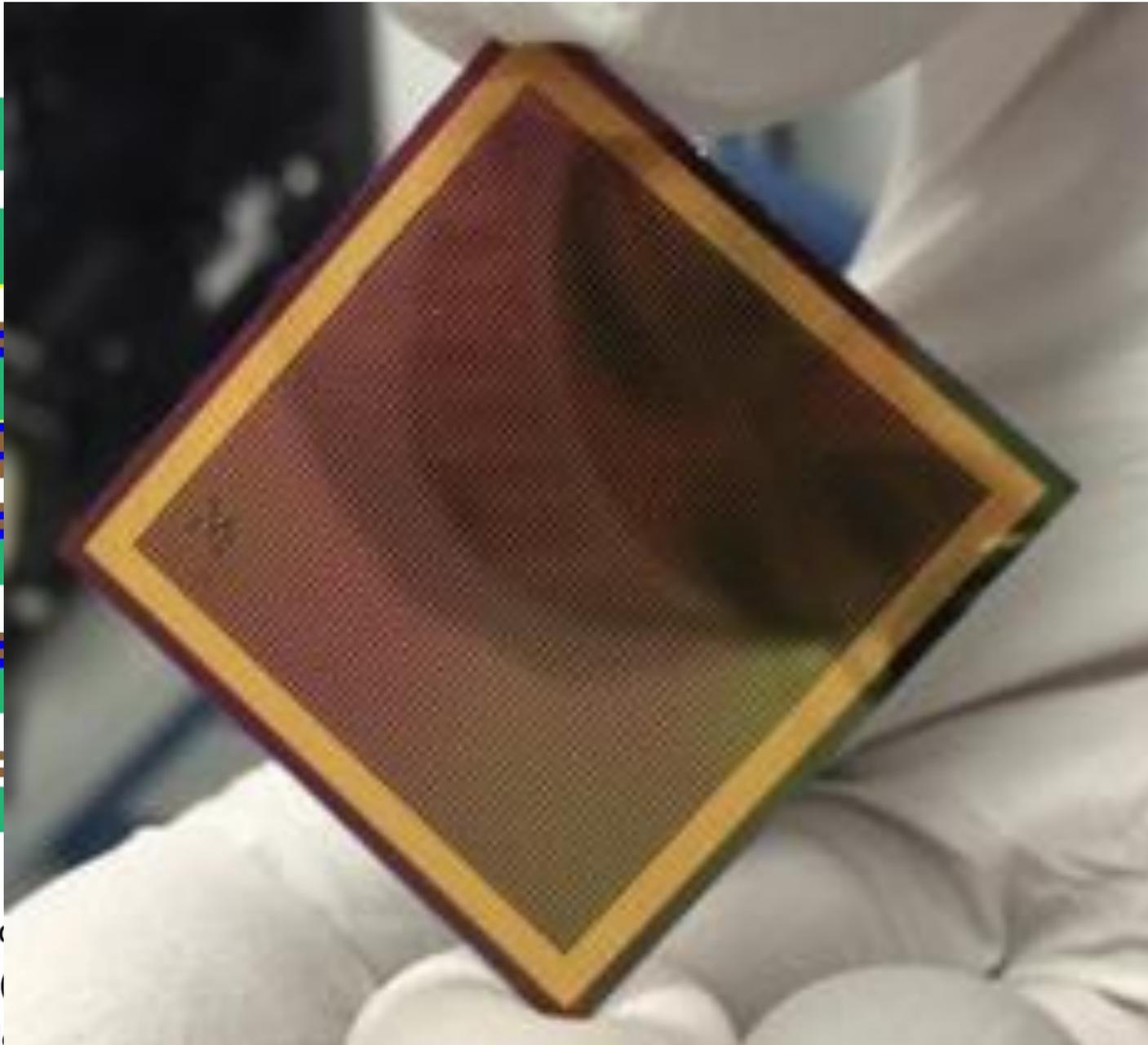
onally low defect level  
and reliability

to packaging of TWV

020 actuator, and  
ed and tested

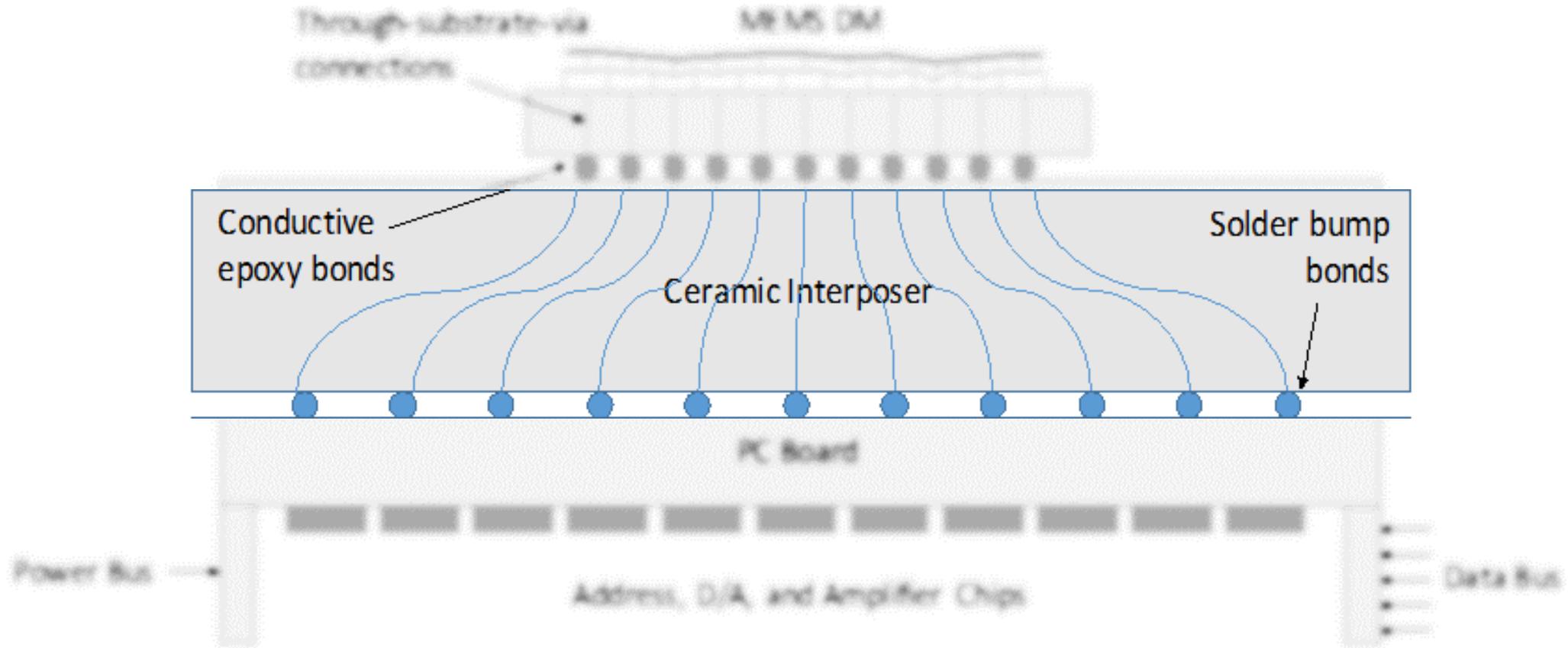


-  Highly Doped Silicon
-  Via Isolation (SiO<sub>2</sub>)
-  Silicon Nitride



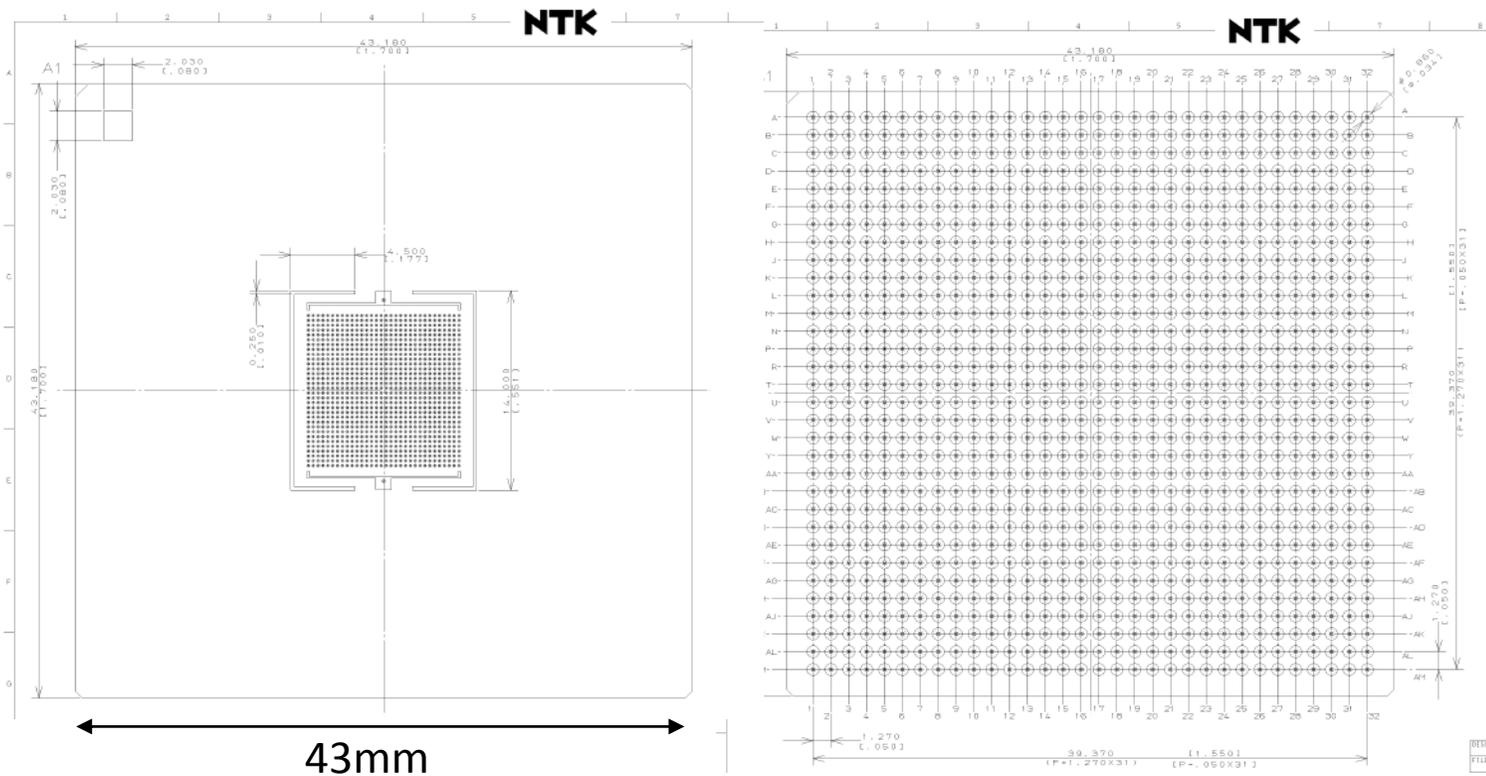


# Proposed Architecture





# Ceramic Interposer



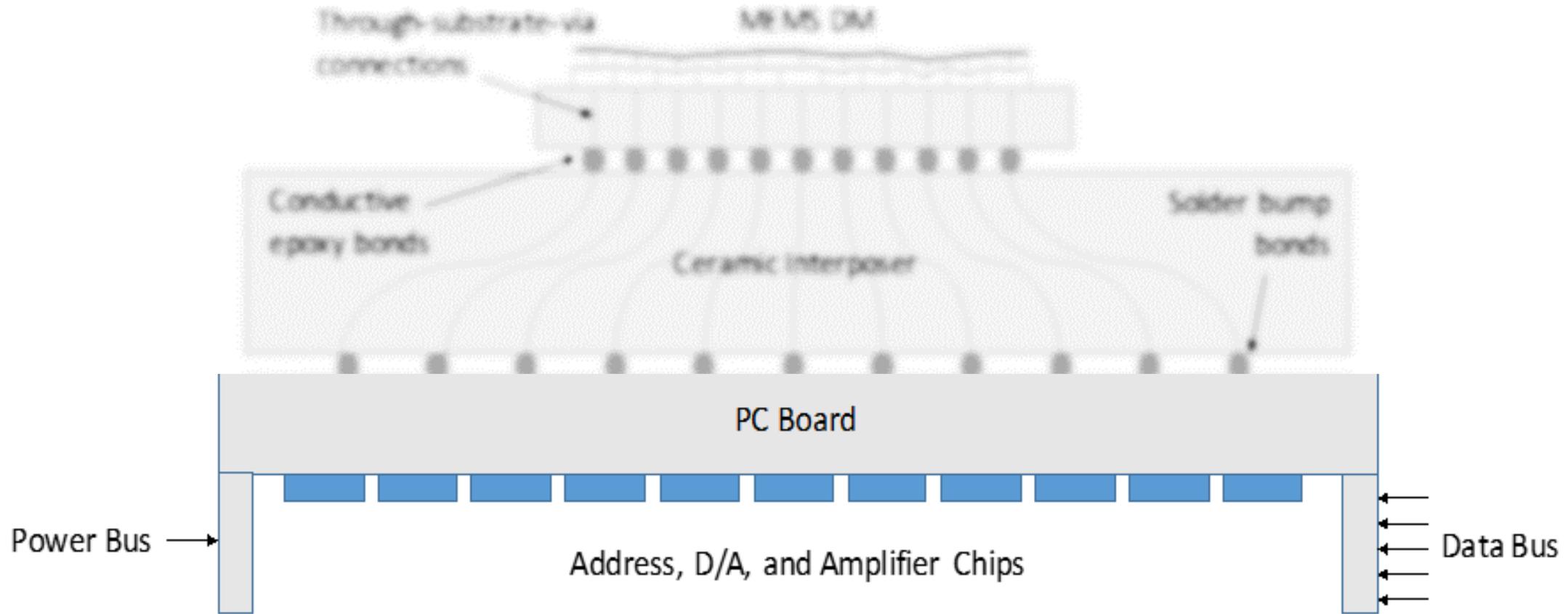
43mm

Currently being fabricated

- Material: aluminum oxide with gold traces
- Top Surface: 1024 LGA Pads electroless gold, (32x32 on 340um pitch)
- Bottom Surface: 1024 BGA pads Electroless gold (32x32 on 1mm pitch)

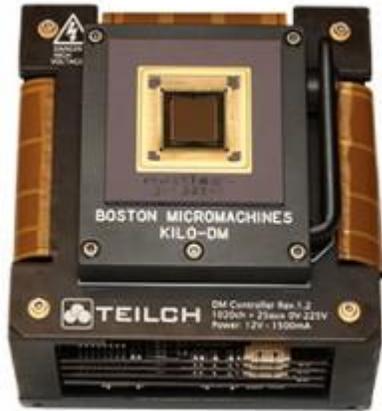


# Proposed Architecture (Concept)

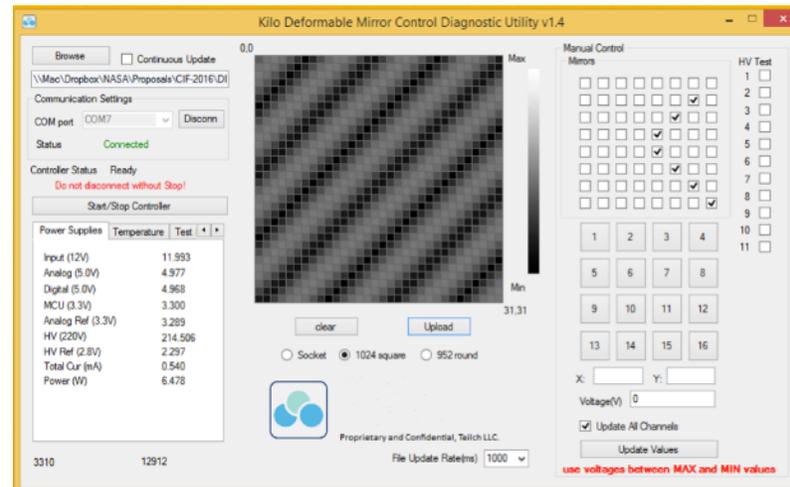




# Electronics Design



- The controller has a volume of 90mm (w) x 90mm (l) x 54.6mm (h), w/o mirror and socket.
- It only requires a 12V power supply and consumes 6W.
- USB interface for data
- 0-215V, 16 bits
- Scalable technology for greater channel count





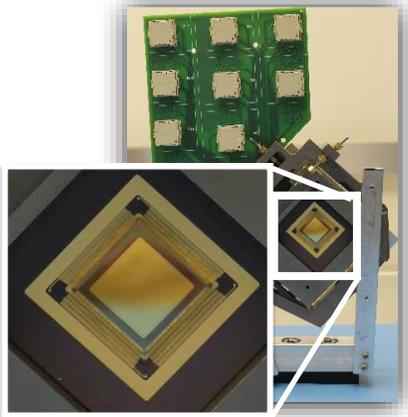
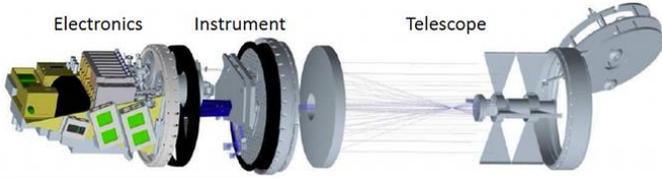
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- Ground astronomy operation



# PICTURE-B PROGRAM

PI: Supriya Chakrabarti, UMASS Lowell



Launched 2015

DOUGLAS 2016, COURTESY UML

# CUBESAT: DEFORMABLE MIRROR DEMONSTRATION MISSION (DEMI)

PI: John Merk, Aurora Flight Systems, Keri Kahoy, MIT



# PICTURE-C PROGRAM

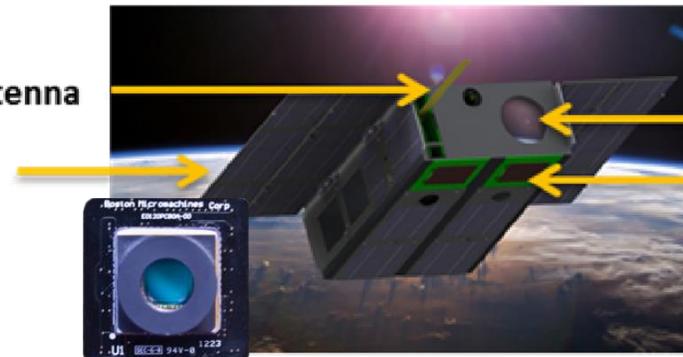
PI: Supriya Chakrabarti, UMASS Lowell



DM Delivered November 2017  
First Flight 2018

Monopole Antenna

Solar panels



Payload Aperture

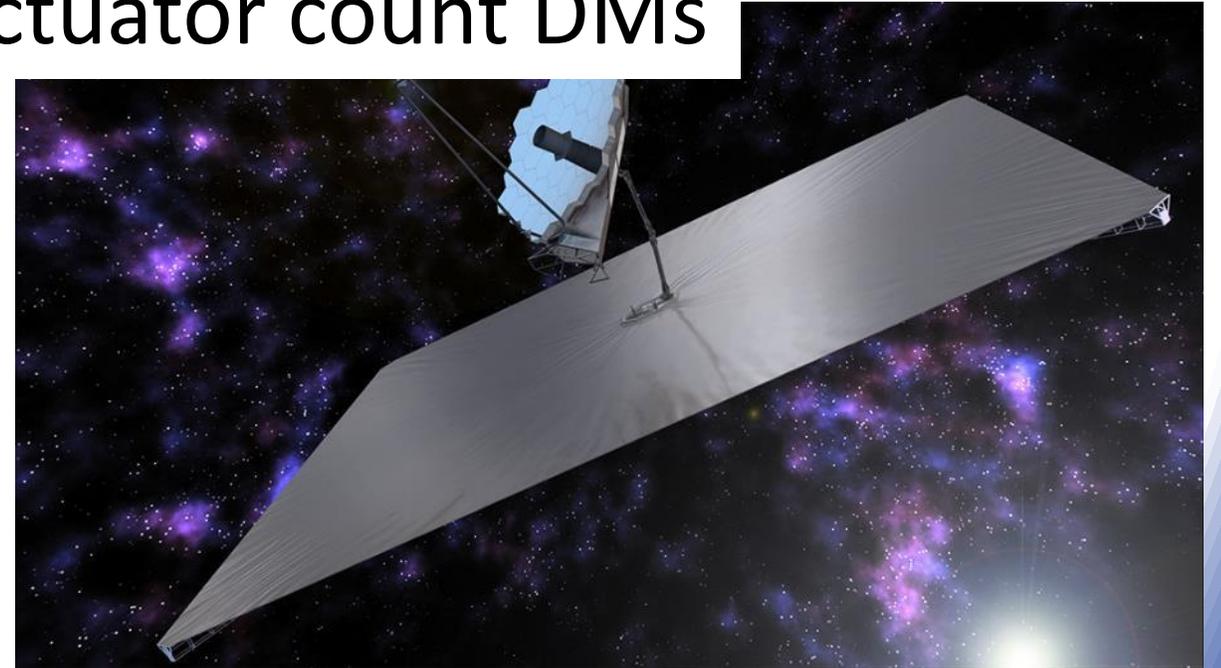
Patch Antennas

Proposed mission configuration



# Future Space Missions

Need for higher actuator count DMs



Habitable Exoplanet Imaging Mission (HabEx)    Large UV/Optical/IR Surveyor (LUVOIR)



# Outline

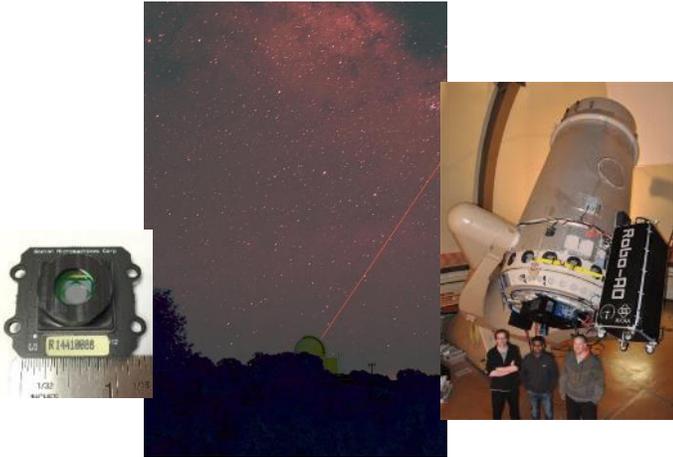
- BMC DM Technology
- NASA funded mirror technology programs
- Space astronomy operations
- **Ground astronomy operation**

# On-Sky Instruments using BMC Mirrors



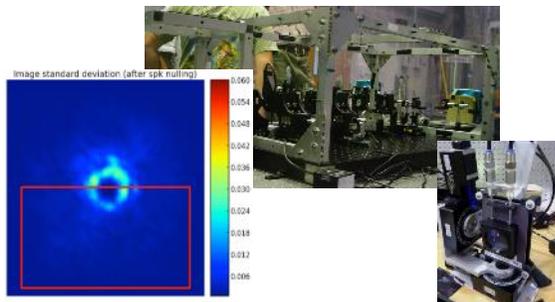
## ROBO-AO

- Multi-DM Installed Palomar 2011/ Moved to Kitt Peak 2015



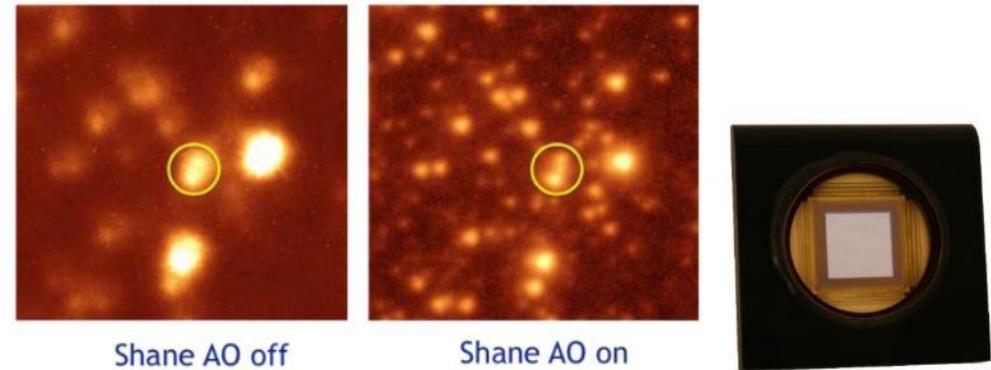
## SCEXAO, Subaru telescope

- 2040 installed 2013



## Shane-AO, Lick Observatory

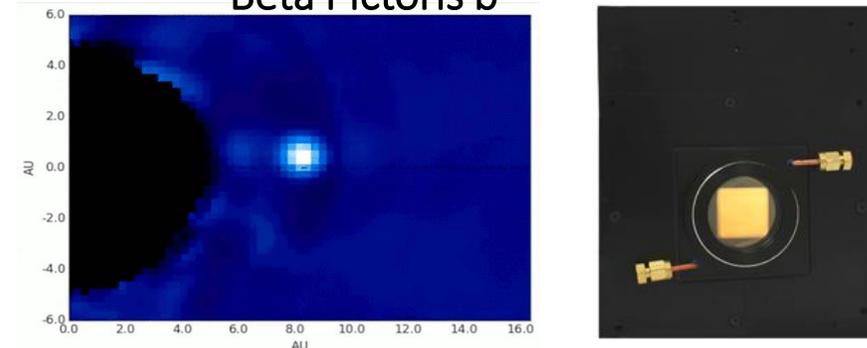
- Kilo-DM installed 2013
- Visible Light Laser Guidestar Experiments

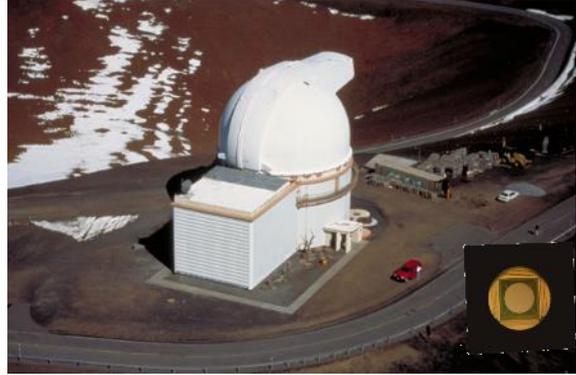


Portion of the M92 globular cluster taken in H band.

## Gemini Planet Imager, Gemini South

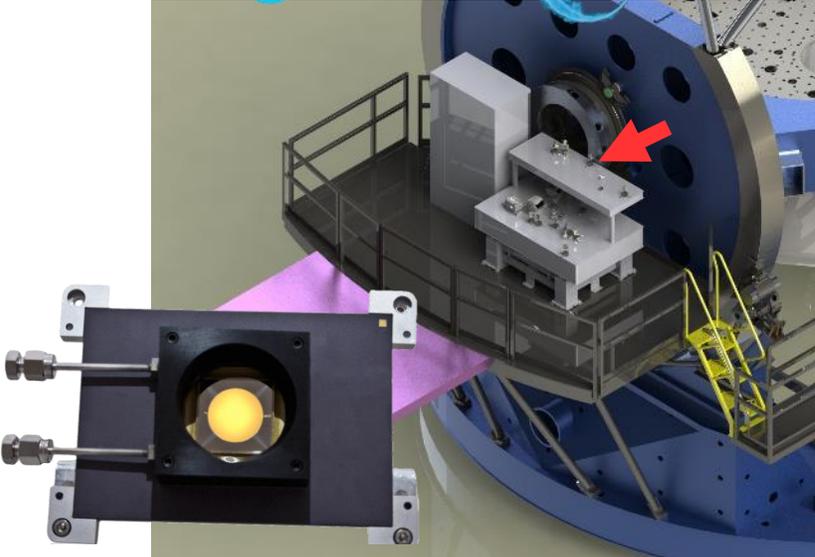
- 4092 installed 2013
- ### Beta Pictoris b



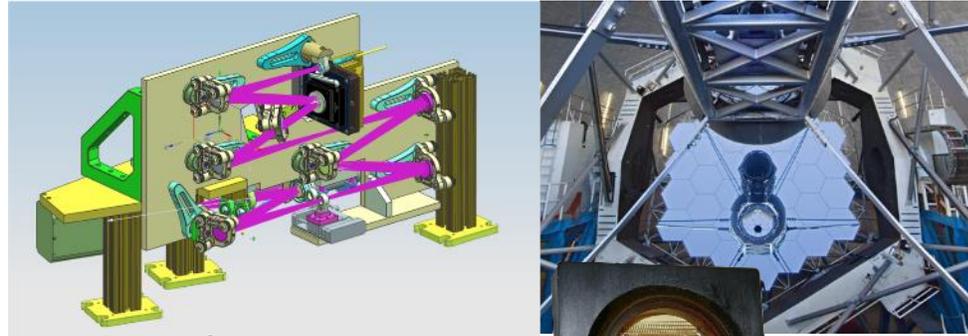


Rapid Transit Surveyor  
492 DM  
Successor of Robo-AO

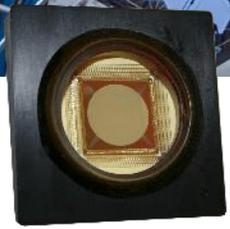
# Next Instruments



MagAO-X on the Clay Nasmyth platform.  
Arrow shows location of BMC 2K.



Keck Planet Imager and  
Characterizer (KPIC)



3.5 um Stroke 952

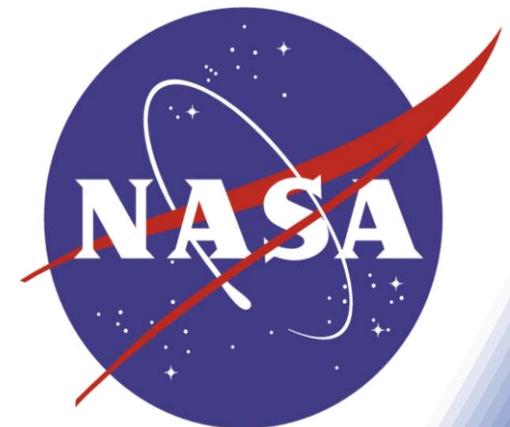
# Conclusion



- Results from our Phase I and II program show good promise for next generation MEMS DMs.
- Testing is ongoing with our TDEM program. Parts are ready for environmental testing.
- APRA program setting new path for high actuator count system

## Acknowledgements

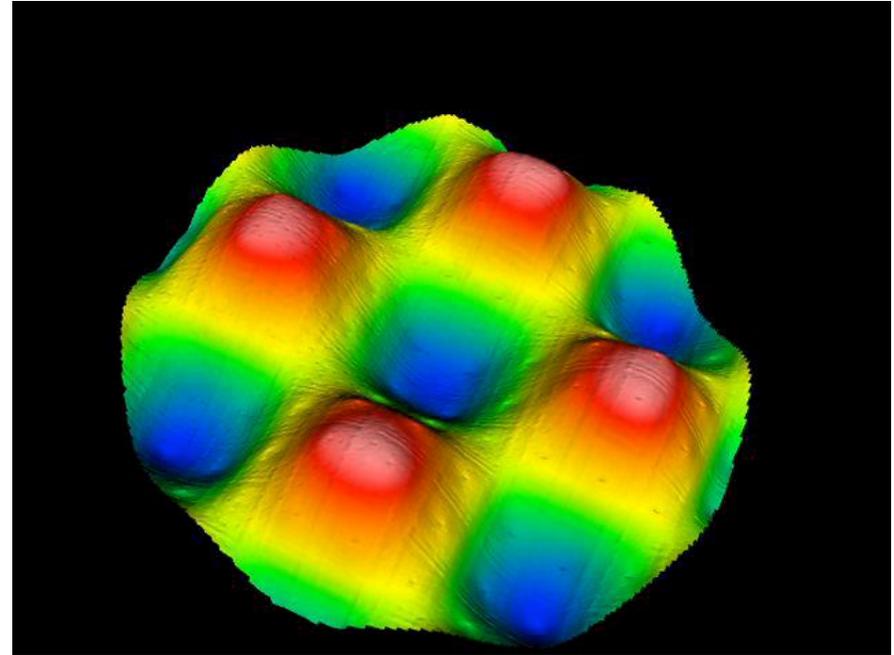
- Funding from NASA
  - Contract#: NNH12CQ27C SAT/TDEM
  - Contract #: NNX16CP14C NASA Phase II SBIR
  - Contract#: NNX17CP76P NASA Phase I SBIR
  - Contract#: NNX17AI66G/80NSSC18K0082 APRA





# Thank You

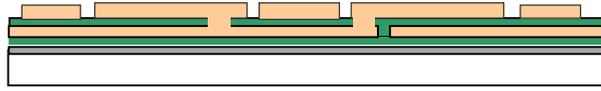
Questions?



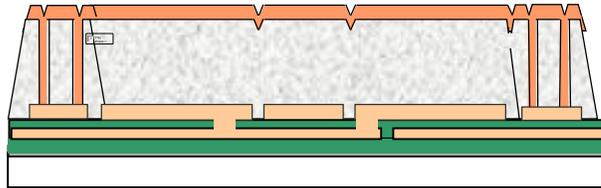
Paul Bierden, [pab@bostonmicromachines.com](mailto:pab@bostonmicromachines.com)

# MEMS DM Fabrication

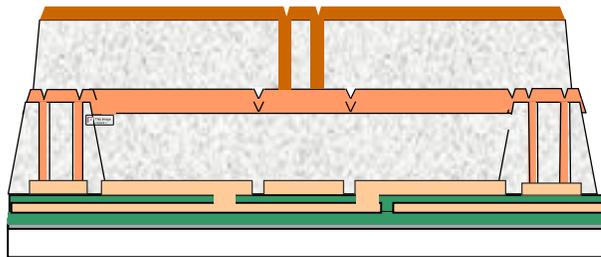
(deposit, pattern, etch, repeat)



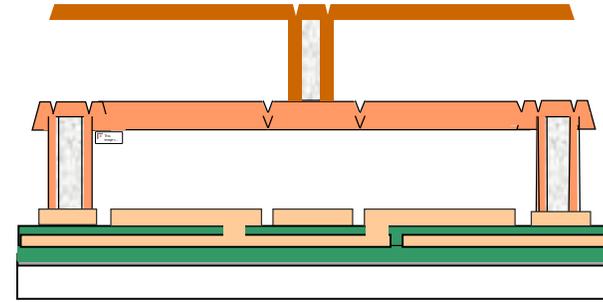
Electrodes & wire traces:  
polysilicon (conductor) & silicon nitride (insulator)



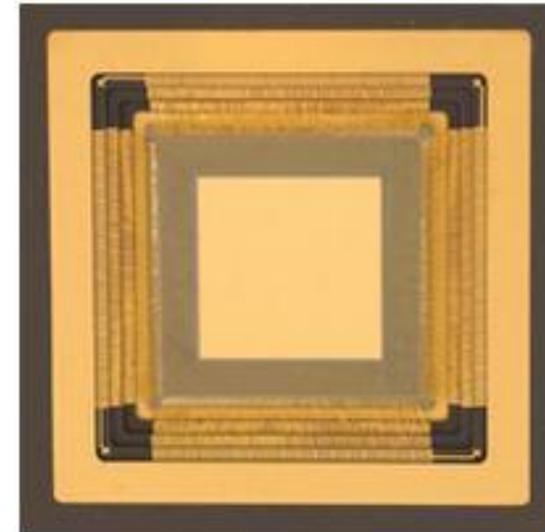
Actuator array:  
oxide (sacrificial spacer) and polysilicon (actuator structure)



Mirror membrane:  
oxide (spacer) and polysilicon (mirror)



MEMS DM:  
Etch away sacrificial oxides in HF, and  
deposit reflective coating

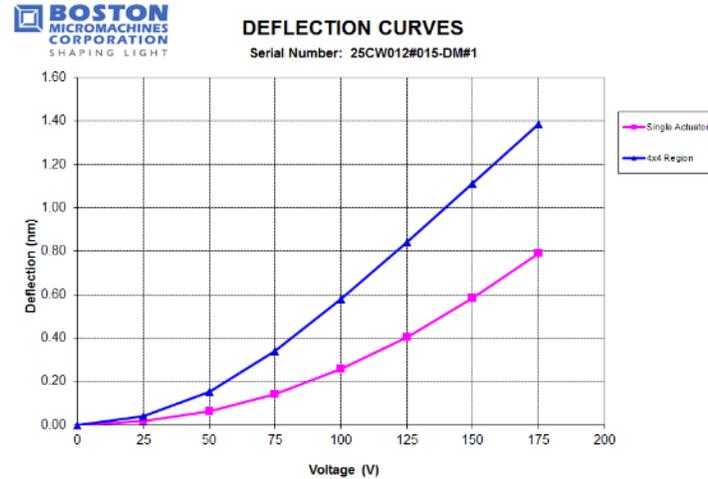


Attach die to a ceramic package and  
wirebond

# 12 DMs Fabricated and Characterized

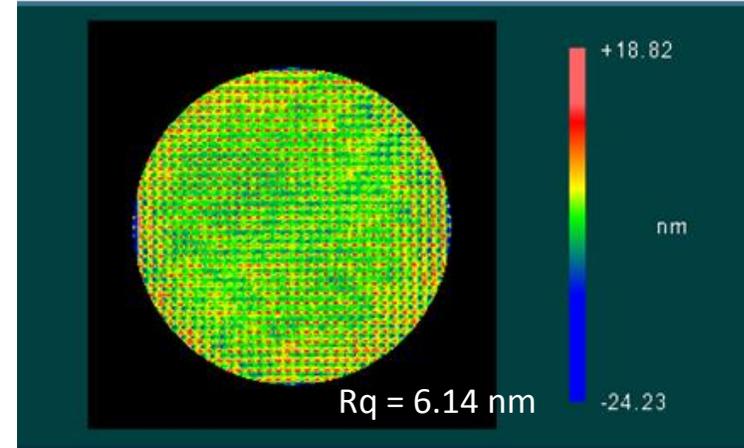


### Voltage v. Deflection

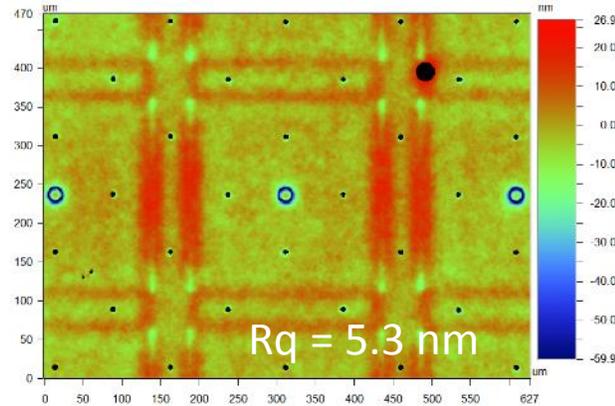


### Active Flattening of DM Surface

Continuous DM Surface Data  
Powered Flat Image Circular Aperture  
(Tilt Removed)

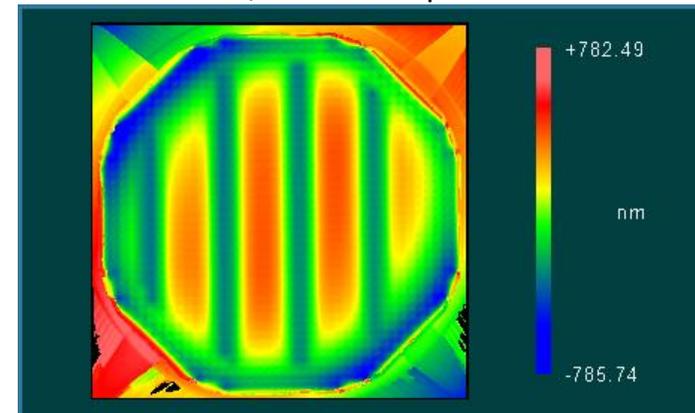


### Single Actuator Surface Figure



### Sinusoid Shape

4 Period, 400nm Amplitude

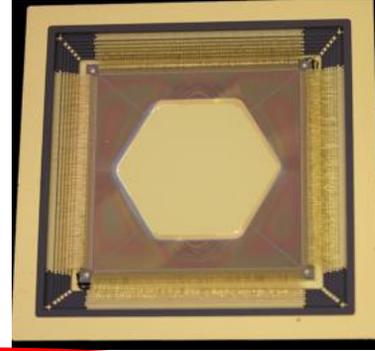


Delivered to JPL (2) and Princeton (2)

# High Spatial Frequency DM: ELTs and Space Astronomy

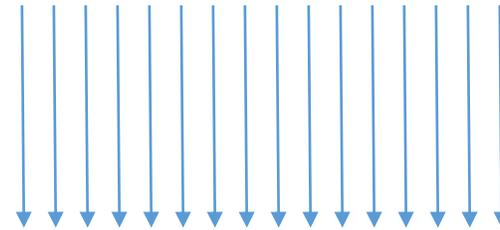


DM



Ground 3-4um stroke, 200 across aperture  
Space 1-2 um stroke, 100 across aperture

Interconnection



Ground and Space same

Electronics



Ground 10kHz, 12-14 bit  
Space 100 Hz, 16+ bit, SWAP important