



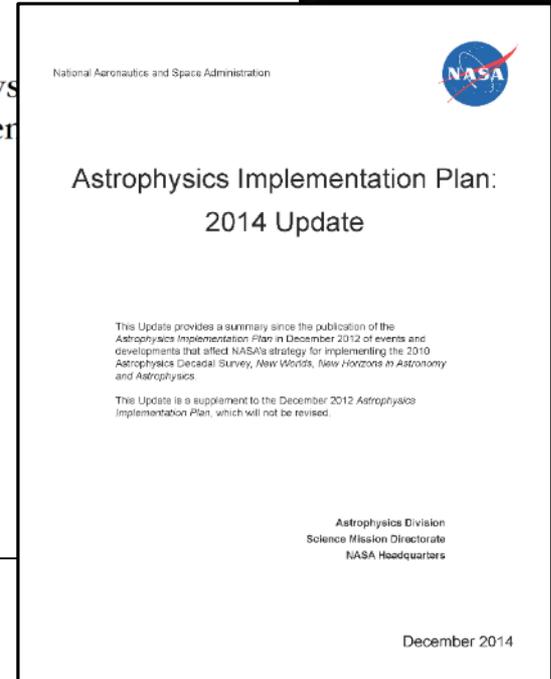
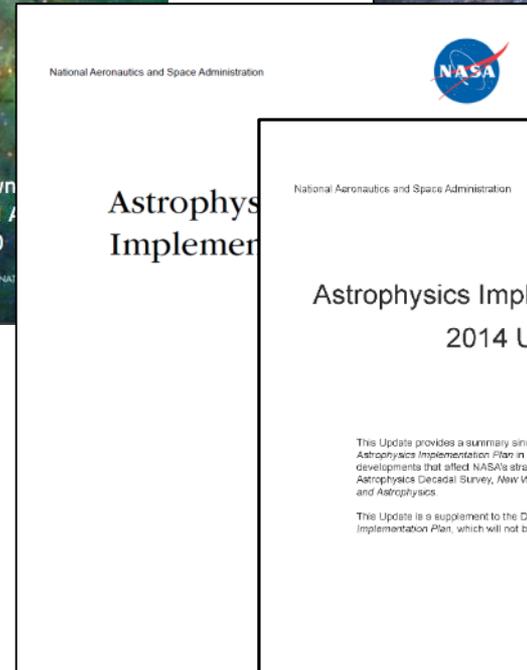
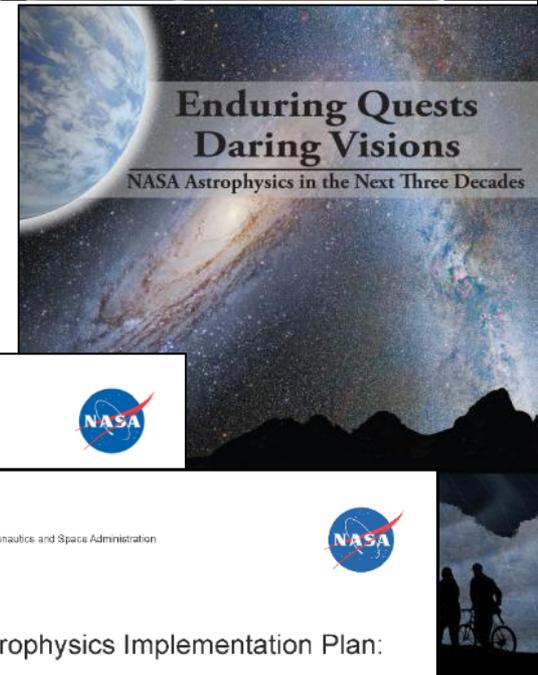
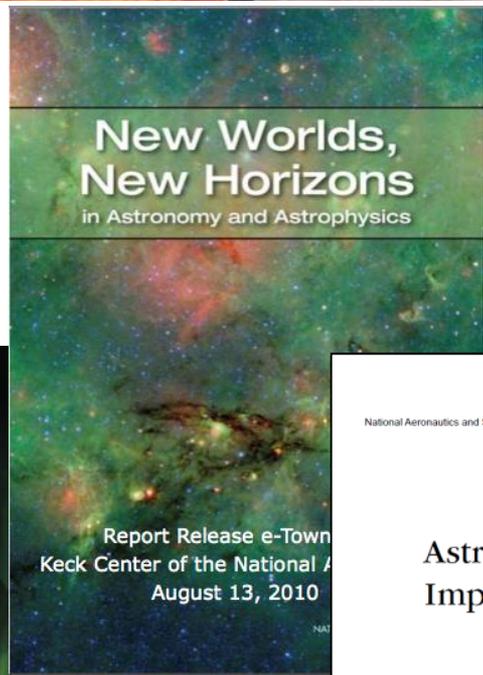
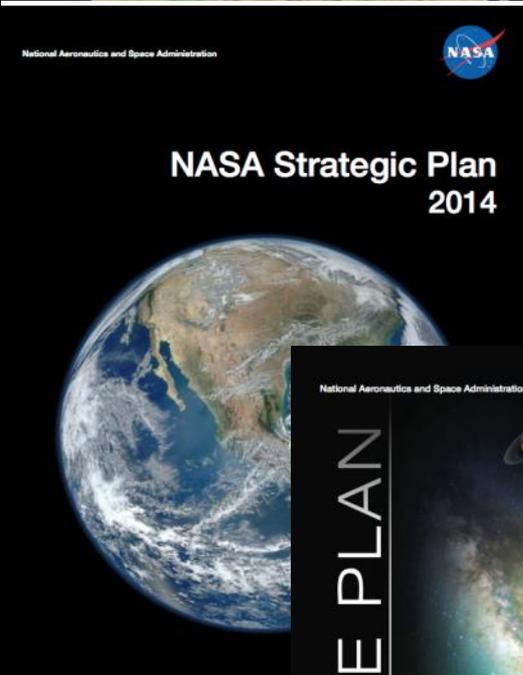
NASA Astrophysics Technology Needs

Astrophysics

Paul Hertz
Director, Astrophysics Division
Science Mission Directorate
[@PHertzNASA](https://twitter.com/PHertzNASA)



Astrophysics Driving Documents

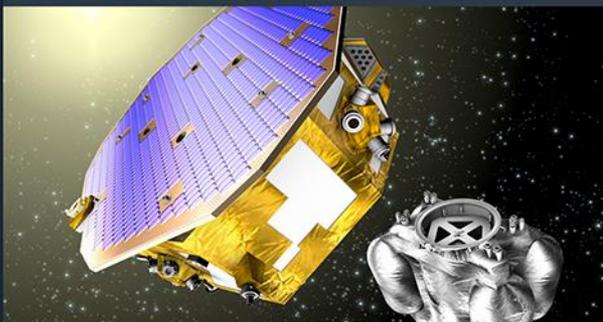


<http://science.nasa.gov/astrophysics/documents>

Astrophysics Missions in Development

LISA Pathfinder ^{12/2015}

ESA-led Mission



NASA supplied the ST7/Disturbance Reduction System (DRS)

ASTRO-H ^{11/2015}

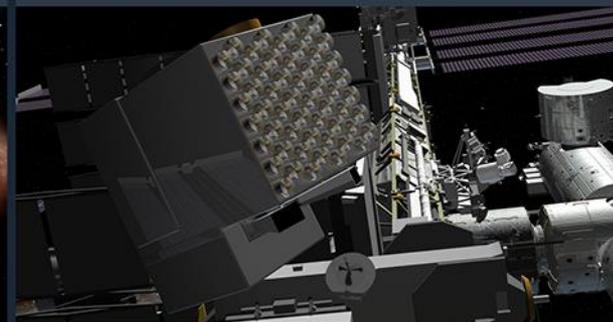
JAXA-led Mission



NASA supplied the Soft X-ray Spectrometer (SXS) instrument

NICER ^{8/2016}

NASA Mission



Neutron Star Interior Composition Explorer

TESS ^{8/2017}

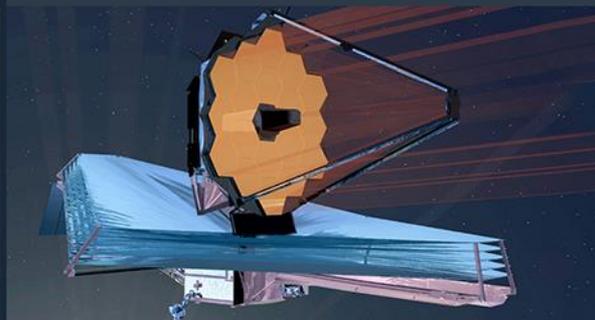
NASA Mission



Transiting Exoplanet Survey Satellite

JWST ^{10/2018}

NASA Mission



James Webb Space Telescope

Euclid ²⁰²⁰

ESA-led Mission



NASA is supplying the NISP Sensor Chip System (SCS)

Astrophysics Missions in Pre-Formulation



SMEX / MO – 2019/2020

IXPE, PRAXyS, SPHEREx
GUSTO, LiteBIRD

MIDEX / MO – 2022/2023

WFIRST-AFTA – NLT 2026

Athena – 2028

All launch dates notional

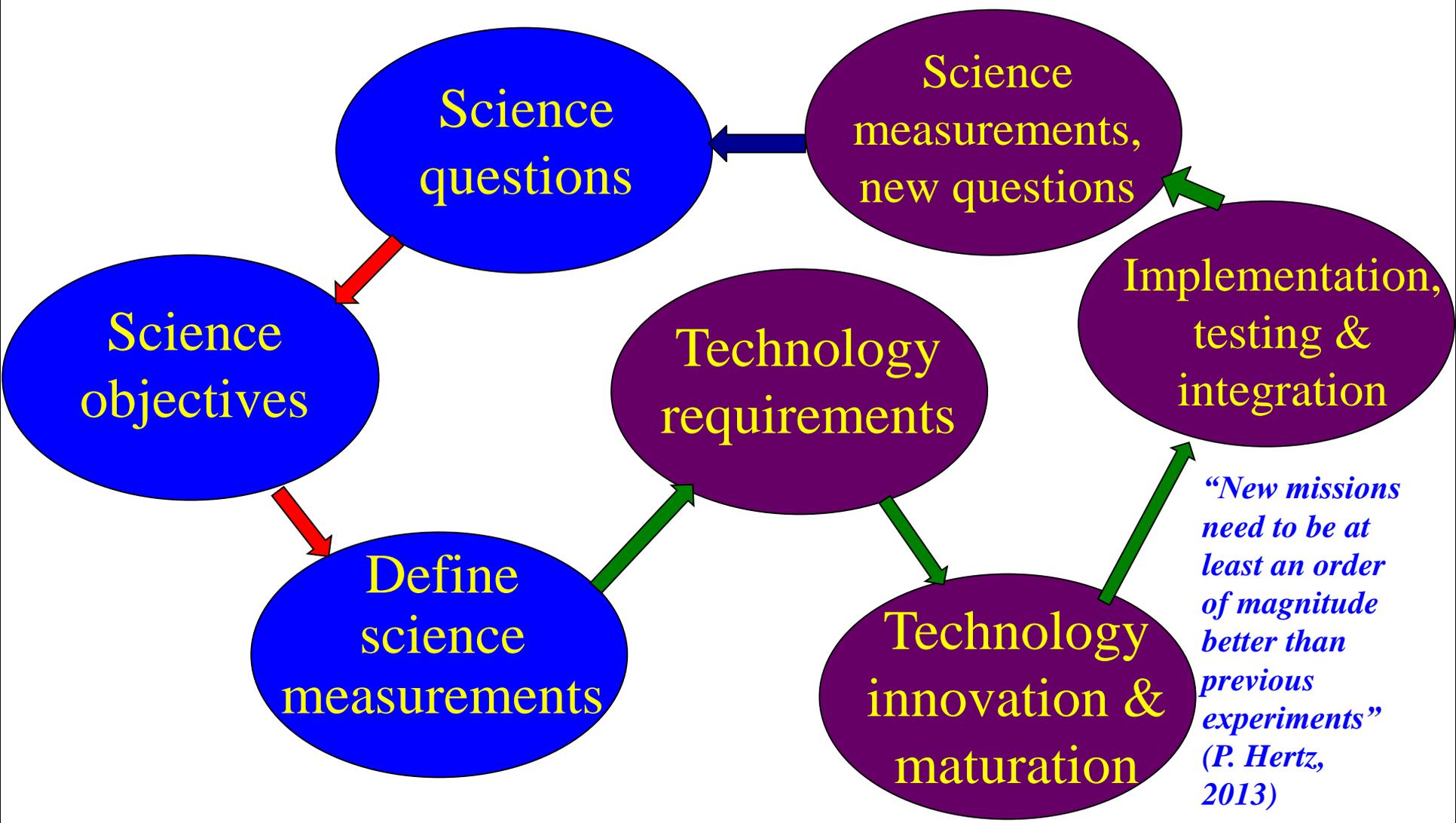


Why Advanced Mirror Technology?

- Astrophysics is a photon starved discipline, demanding high performance from all systems and subsystems utilized for on-sky observations and detections.
- Most of the low-hanging fruit science goals, including synoptic astronomical observations have been explored and exploited.
- New science efforts and techniques such as
 - persistent time domain observations,
 - high contrast imaging,
 - diffraction limited mapping,
 - high spatial and spectral resolution data, and
 - wide field surveyswill open up new discovery spaces with important breakthroughs.
- New wavelength regimes and major gains in sensitivity also lead to significant new discoveries.



Science Flow and Mapping



*“New missions need to be at least an order of magnitude better than previous experiments”
(P. Hertz, 2013)*



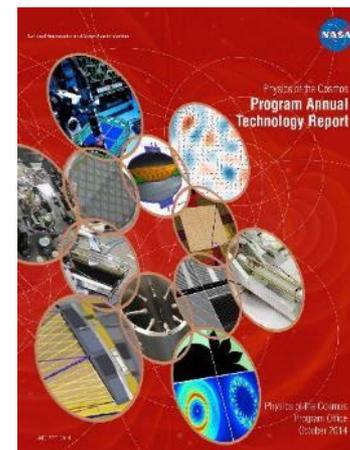
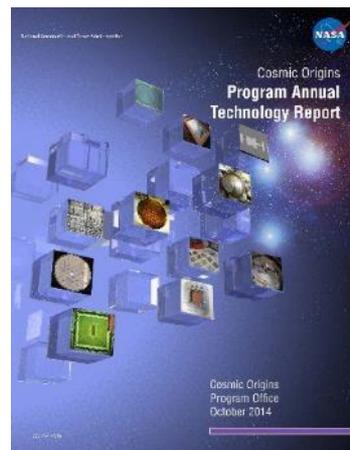
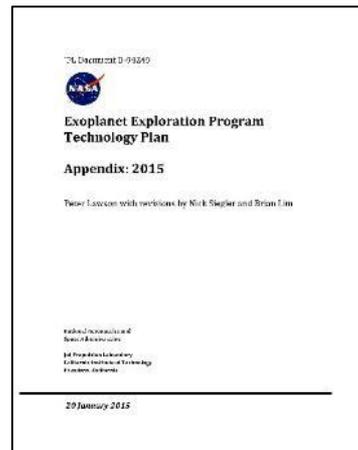
Technology Gaps and Requirements

- Technology gaps are identified and prioritized in the Program Annual Technology Reports (PATRs).
 - PATRs are developed with considerable community input including an open call for identification of technology gaps and use of community based Program Analysis Groups and Technology Assessment Committees to prioritize technology gaps.
 - Gap lists serve to identify where technology development is needed.

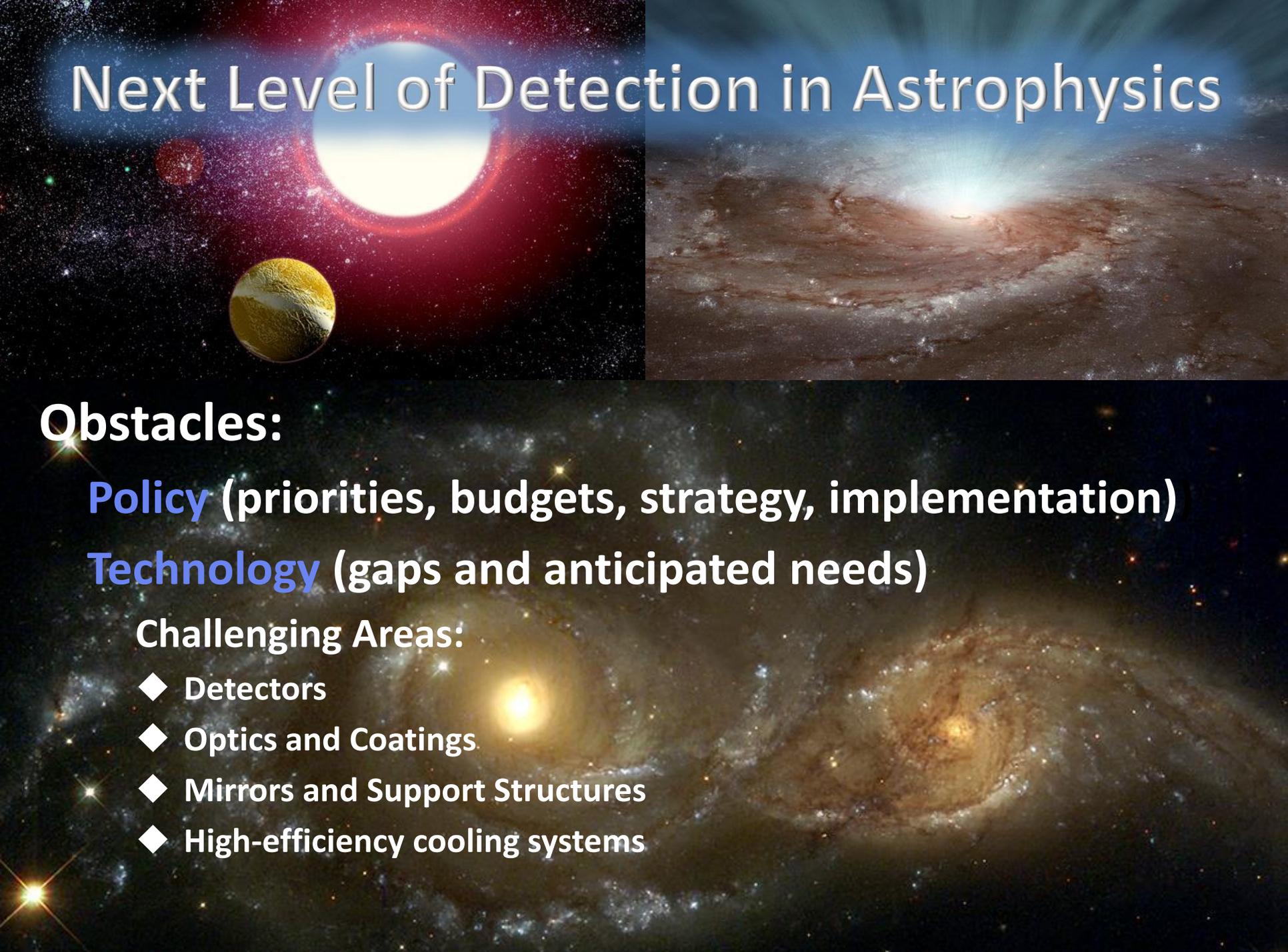
<http://cor.gsfc.nasa.gov/technology/>

<http://exep.jpl.nasa.gov/technology/>

<http://pcos.gsfc.nasa.gov/technology/>



Next Level of Detection in Astrophysics

The background of the slide is a composite of astronomical images. The top left shows a bright yellow star with a prominent red ring, a planet with a yellow and white band, and a dark starry field. The top right shows a galaxy with a bright central core. The bottom right shows a large, detailed image of a galaxy with a bright central core and spiral arms.

Obstacles:

Policy (priorities, budgets, strategy, implementation)

Technology (gaps and anticipated needs)

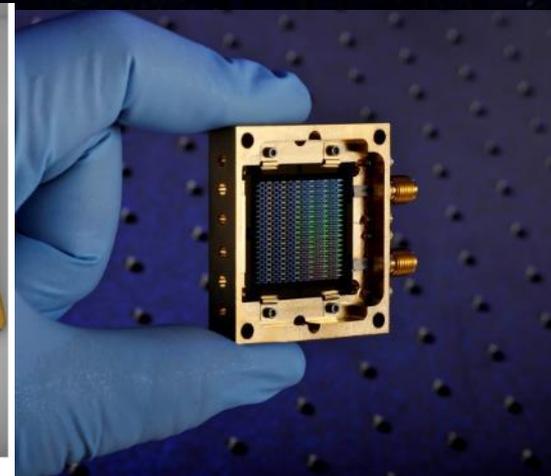
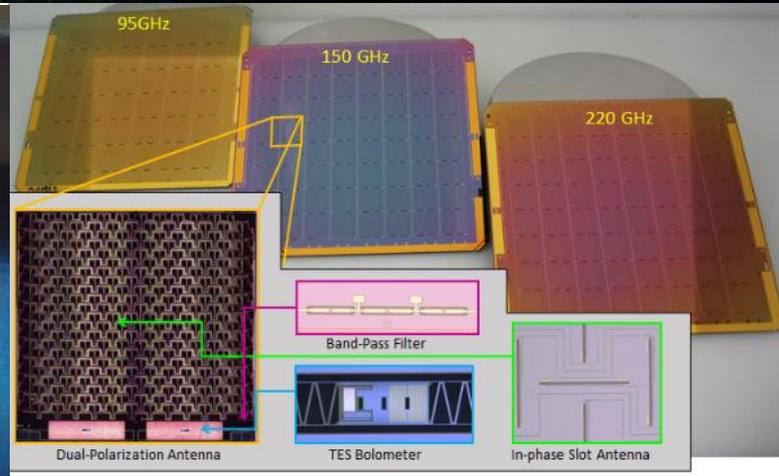
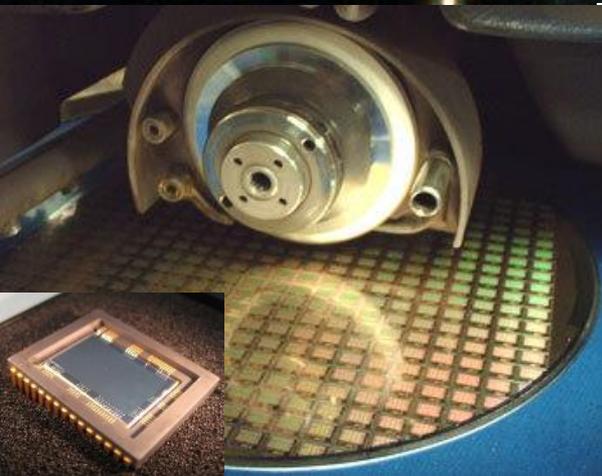
Challenging Areas:

- ◆ Detectors
- ◆ Optics and Coatings
- ◆ Mirrors and Support Structures
- ◆ High-efficiency cooling systems

Detectors (across wavelengths)

Increase efficiency, SNR, resolution, and speed

- ◆ Increase QE (>80-90%)
- ◆ Large format and high pixel count
- ◆ Radiation tolerant
- ◆ Photon-counting
- ◆ Low-power and fast readout
- ◆ Low read-noise
- ◆ Low dark current



Optics and Coatings

Improve system throughput, image quality, and information collected

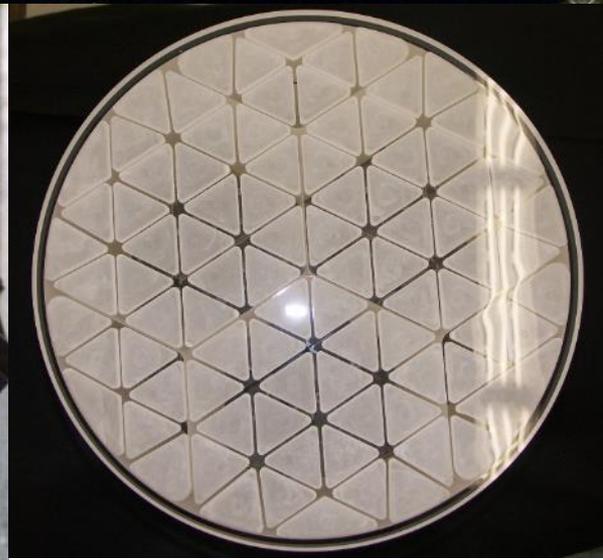
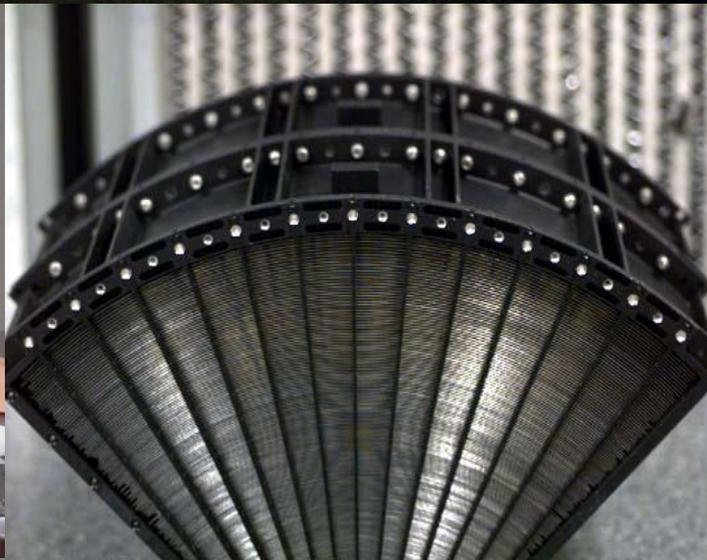
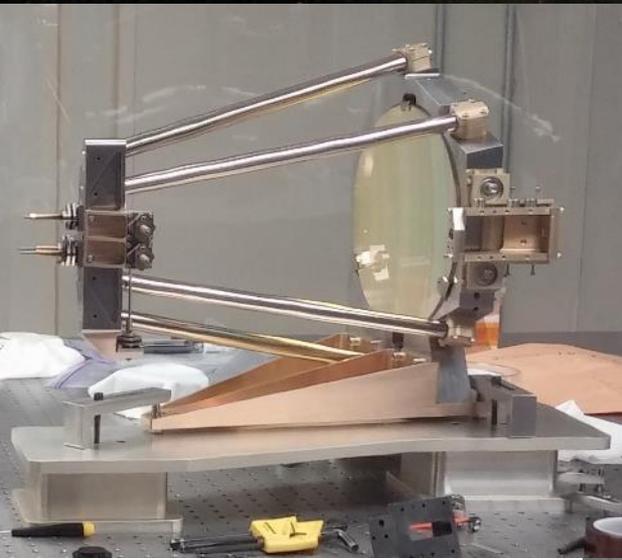
- ◆ High contrast imaging (10^{-10})
- ◆ Wavefront control
- ◆ High spectral and angular resolution X-ray optics
- ◆ X-ray polarimeters
- ◆ X-ray grating arrays
- ◆ Multi-object devices (digital micro-mirror and micro-shutters)
- ◆ Coatings (reflective/UV-Vis, antireflective/far-IR, and low-stress/X-ray optics)
- ◆ Dichroic filters
- ◆ Interferometers



Mirrors and Structures

Improve stability, performance, and efficiency of light collection

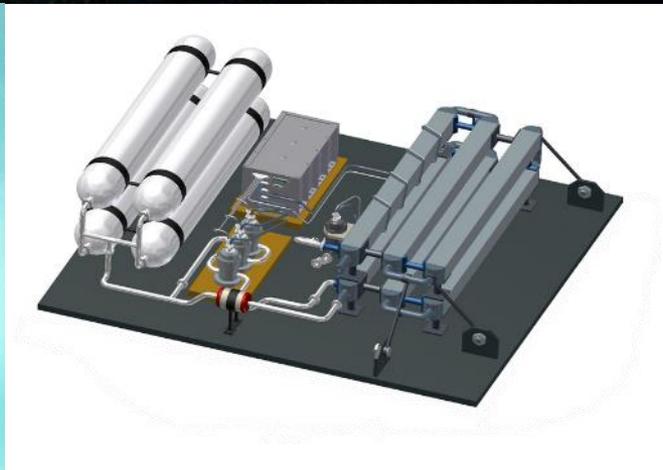
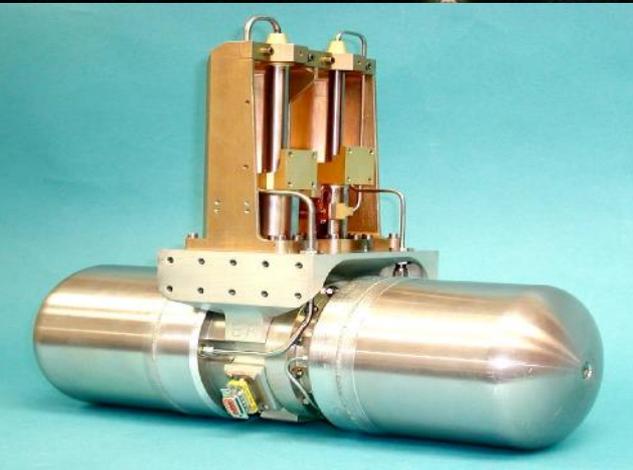
- ◆ Advanced X-ray mirror technologies
- ◆ UVOIR mirror materials
- ◆ Ultra stability (sensing and control from micrometers, nanometers to picometers)
 - ✧ Nano composite materials (\sim zero CTE)
 - ✧ Actuators
 - ✧ Metrology (lasers and measuring techniques)



High-efficiency cooling systems

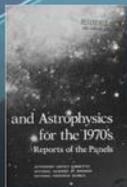
Improve efficiency and heat-lift at ultralow temperatures

- ◆ High-performance sub-Kelvin coolers
- ◆ Advanced cryo-coolers
- ◆ Solid-state coolers

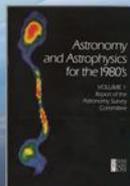


ASTROPHYSICS

Decadal Survey Missions



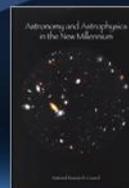
1972
Decadal
Survey
Hubble



1982
Decadal
Survey
Chandra



1991
Decadal
Survey
Spitzer, SOFIA

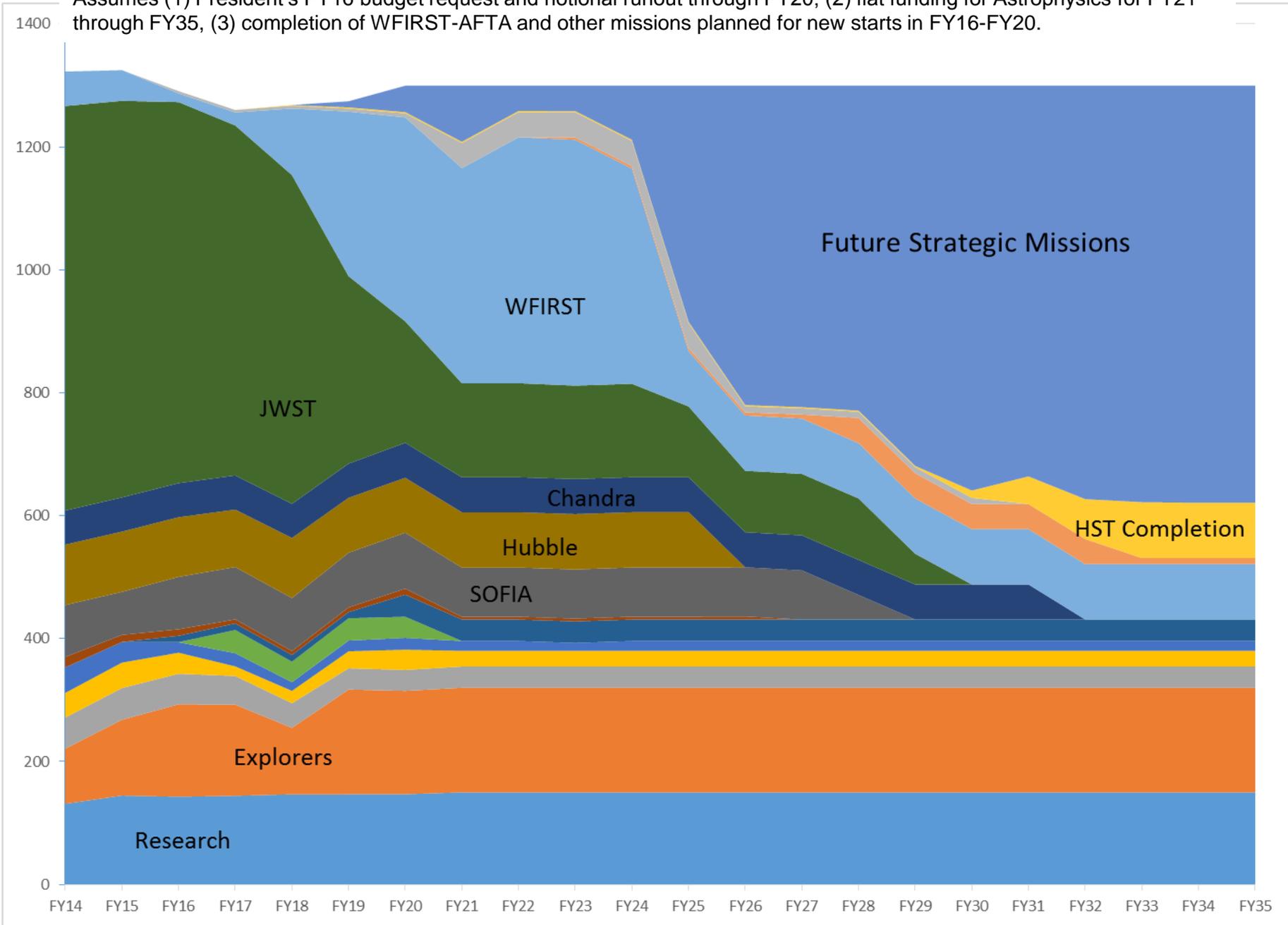


2001
Decadal
Survey
JWST



2010
Decadal
Survey
WFIRST

Assumes (1) President's FY16 budget request and notional runout through FY20, (2) flat funding for Astrophysics for FY21 through FY35, (3) completion of WFIRST-AFTA and other missions planned for new starts in FY16-FY20.





Preparing for the 2020 Decadal Survey Large Mission Concepts

The recommended short list (in alphabetical order):

- **FAR IR Surveyor** – The Astrophysics Visionary Roadmap identifies a Far IR Surveyor as contributing through improvements in sensitivity, spectroscopy, and angular resolution.
- **Habitable-Exoplanet Imaging Mission** – The 2010 Decadal Survey recommends that a habitable-exoplanet imaging mission be studied in time for consideration by the 2020 Decadal Survey.
- **UV/Optical/IR Surveyor** – The Astrophysics Visionary Roadmap identifies a UV/Optical/IR Surveyor as contributing through improvements in sensitivity, spectroscopy, high contrast imaging, astrometry, angular resolution and/or wavelength coverage. The 2010 Decadal Survey recommends that NASA prepare for a UV mission to be considered by the 2020 Decadal Survey.
- **X-ray Surveyor** – The Astrophysics Visionary Roadmap identifies an X-ray Surveyor as contributing through improvements in sensitivity, spectroscopy, and angular resolution.



Preparing for the 2020 Decadal Survey Large Mission Concepts

- NASA Plan for Community Input
 - 2015: PAGs gather community input on selecting concepts for study
 - 2016: Appoint STDT and Center study office, STDT assesses technology
 - 2017: Fund technology development through SAT, STDT develops DRM
 - 2018: STDT submits DRM for cost assessment
 - 2019: STDT issues report and provides input to Decadal Survey

NASA plans a call for nominations (including self-nominations) for membership on the Science and Technology Definition Teams (STDTs) in early 2016.

All are invited to respond to the call.

- Formulation
- Implementation
- Primary Ops
- Extended Ops



XMM-Newton (ESA)
12/10/1999



Swift
11/20/2004



Suzaku (JAXA)
7/10/2005



Fermi
6/11/2008



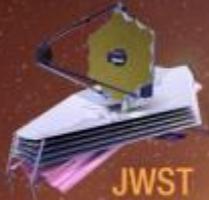
Euclid (ESA)
2020



Hubble
4/24/1990



Kepler
3/6/2009



JWST
2018



Spitzer
8/25/2003



Astro-H (JAXA)
2015

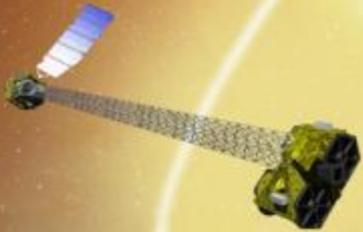


Chandra
7/23/1999



NICER (on ISS)
2016

NuSTAR
6/13/2012



TESS
2017



LISA Pathfinder (ESA)
2015



SOFIA
Full Ops 2014