

PHYSICAL & CHEMICAL CHARACTERIZATION OF LUNAR REGOLITH: CONSIDERATIONS FOR SIMULANTS



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Moon is an Enormous Earth-Orbiting Space Station

**A Natural Satellite outside of Earth's Gravity Well,
With Raw Materials that can be put to Practical Use
As Humanity Expands Outward into the Solar System.**

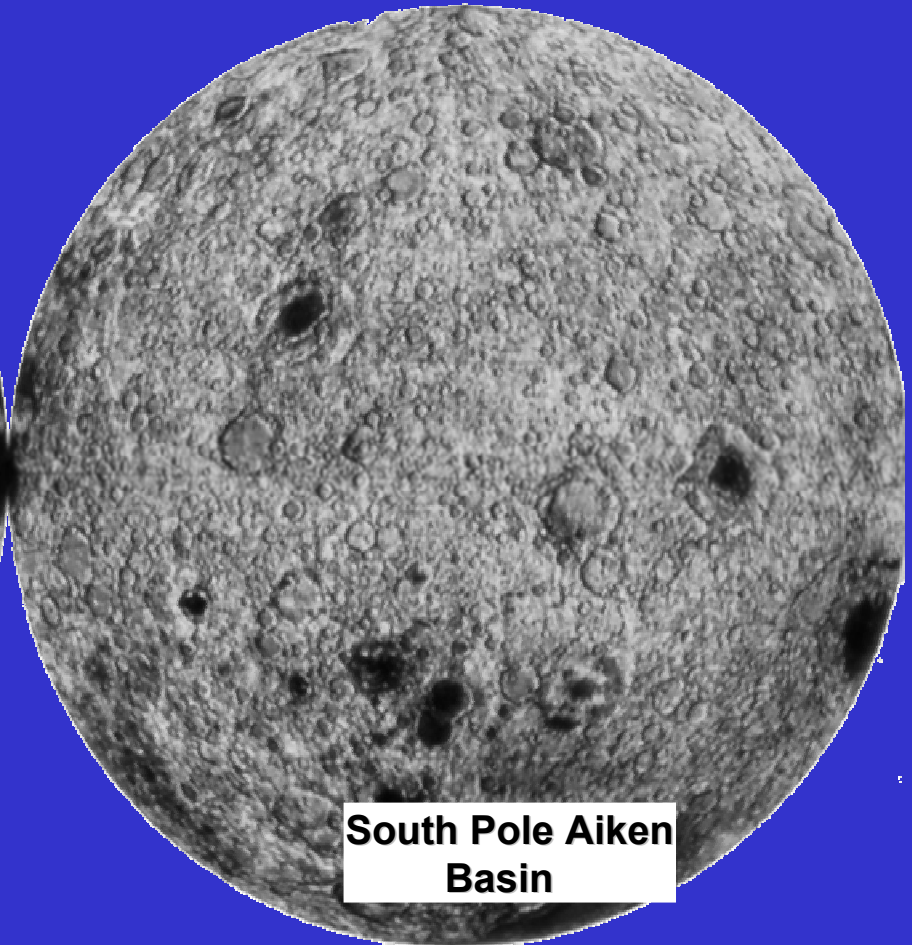
OUTLINE

- ◆ CHARACTERIZATION OF LUNAR SOIL
- ◆ UNIQUE PROPERTIES OF LUNAR SOIL
- ◆ ISRU CONSIDERATIONS
- ◆ INTRODUCTION TO SOIL SIMULANTS

**16% Maria and 84% Highlands
Contrasting Regions of the Moon**

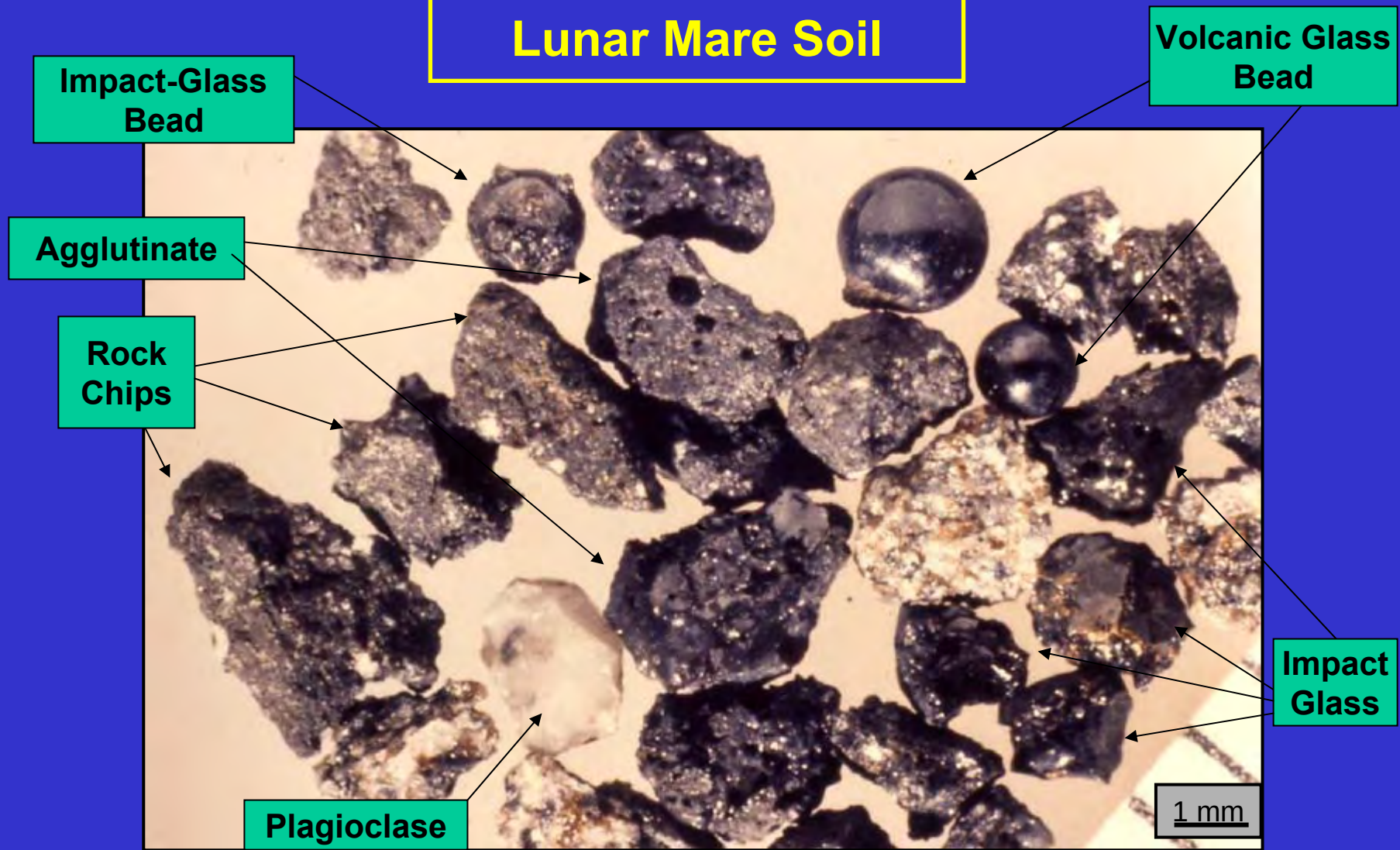


South Pole Amundsen Site



**South Pole Aiken
Basin**

Lunar Mare Soil

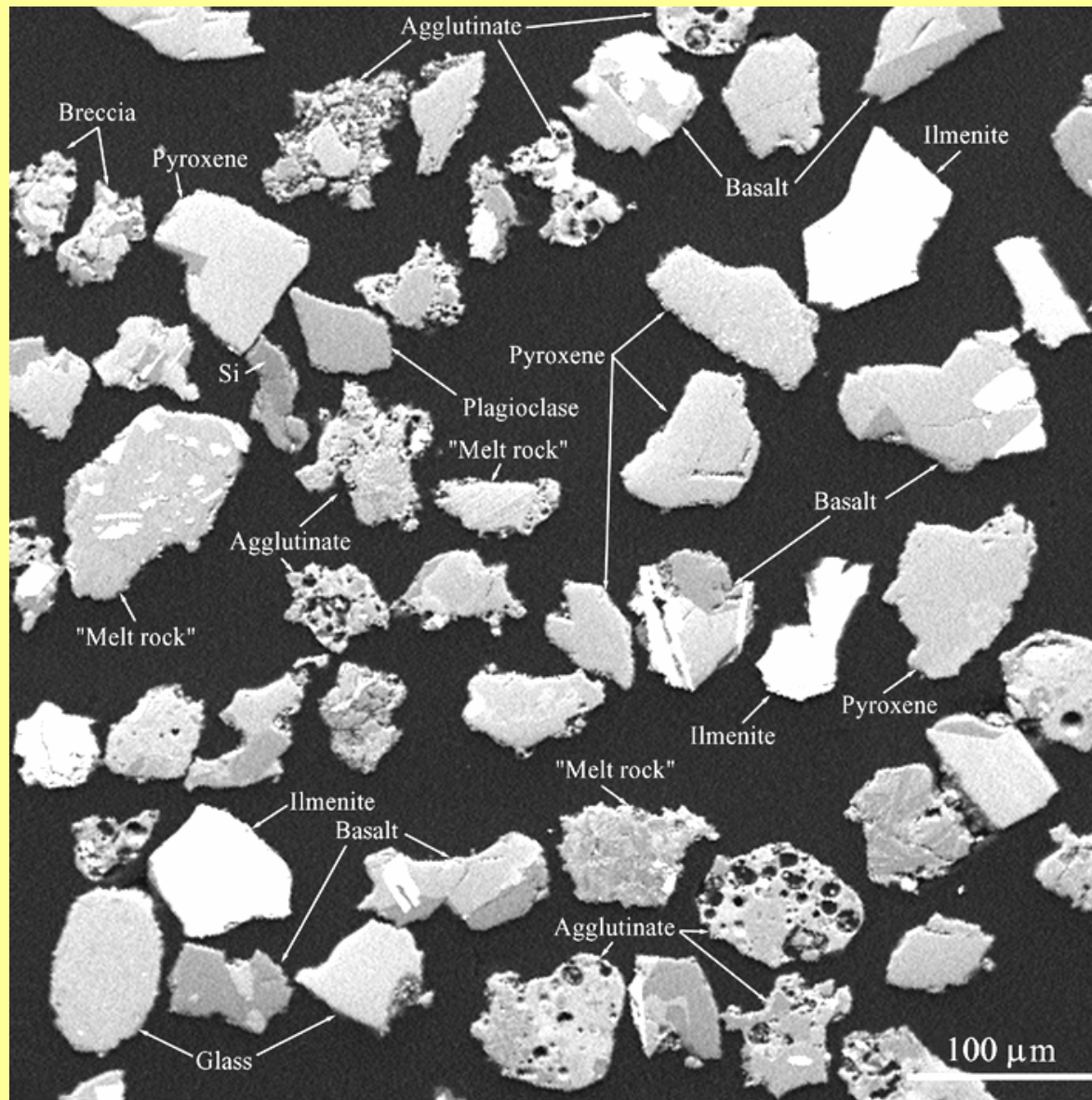


Regolith: broken up rock material; Soil: <1 cm portion of the Regolith
Dust: < 50 μm portion of the Soil

Major Components of Lunar Soils:

- Rock Fragments**
- Mineral Fragments**
- Glassy Particles**
 - Agglutinates**
 - Impact-Glass Droplets**
 - Volcanic-Glass Beads**

Lunar Mare Soil

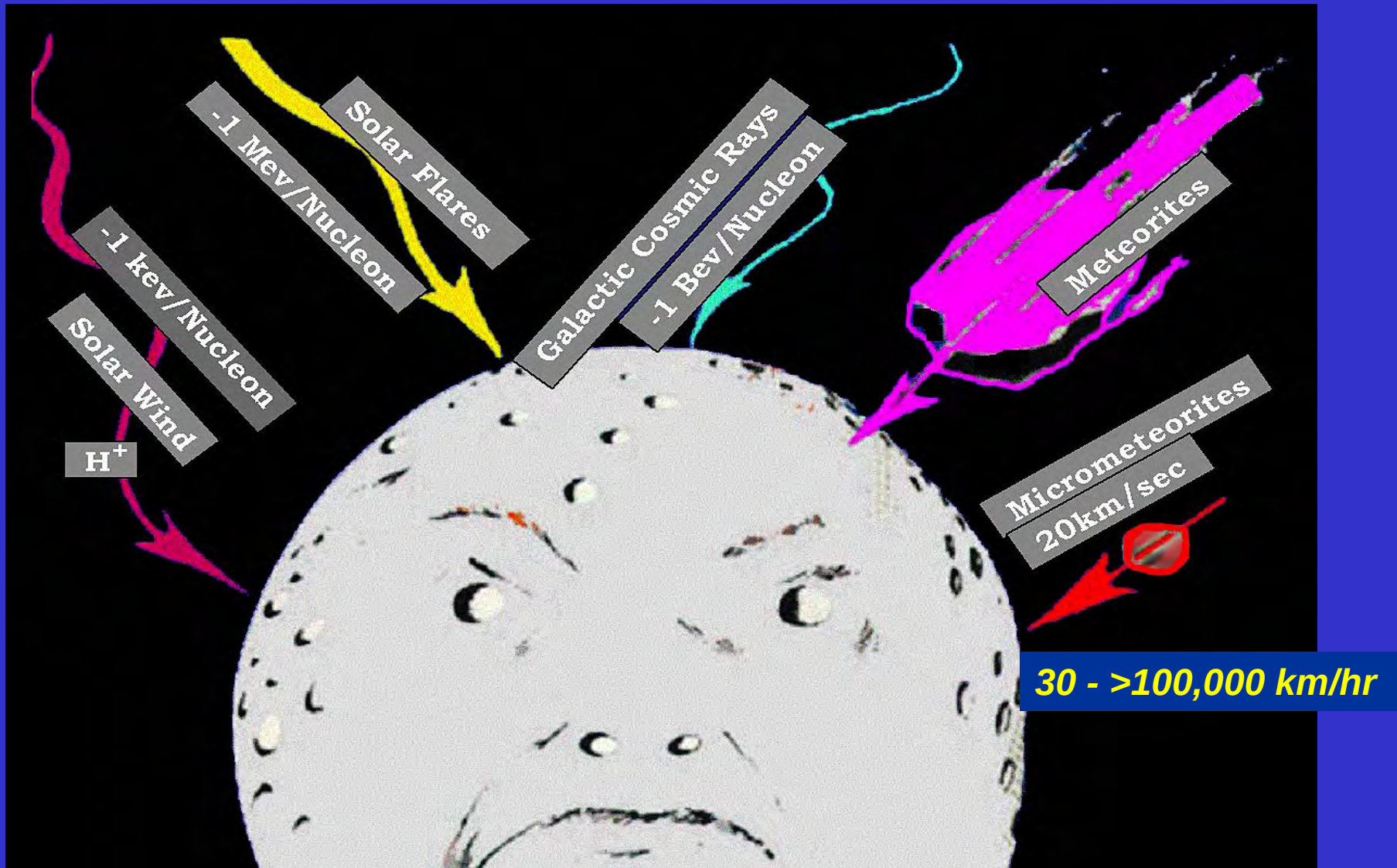


Space Weathering on the Moon

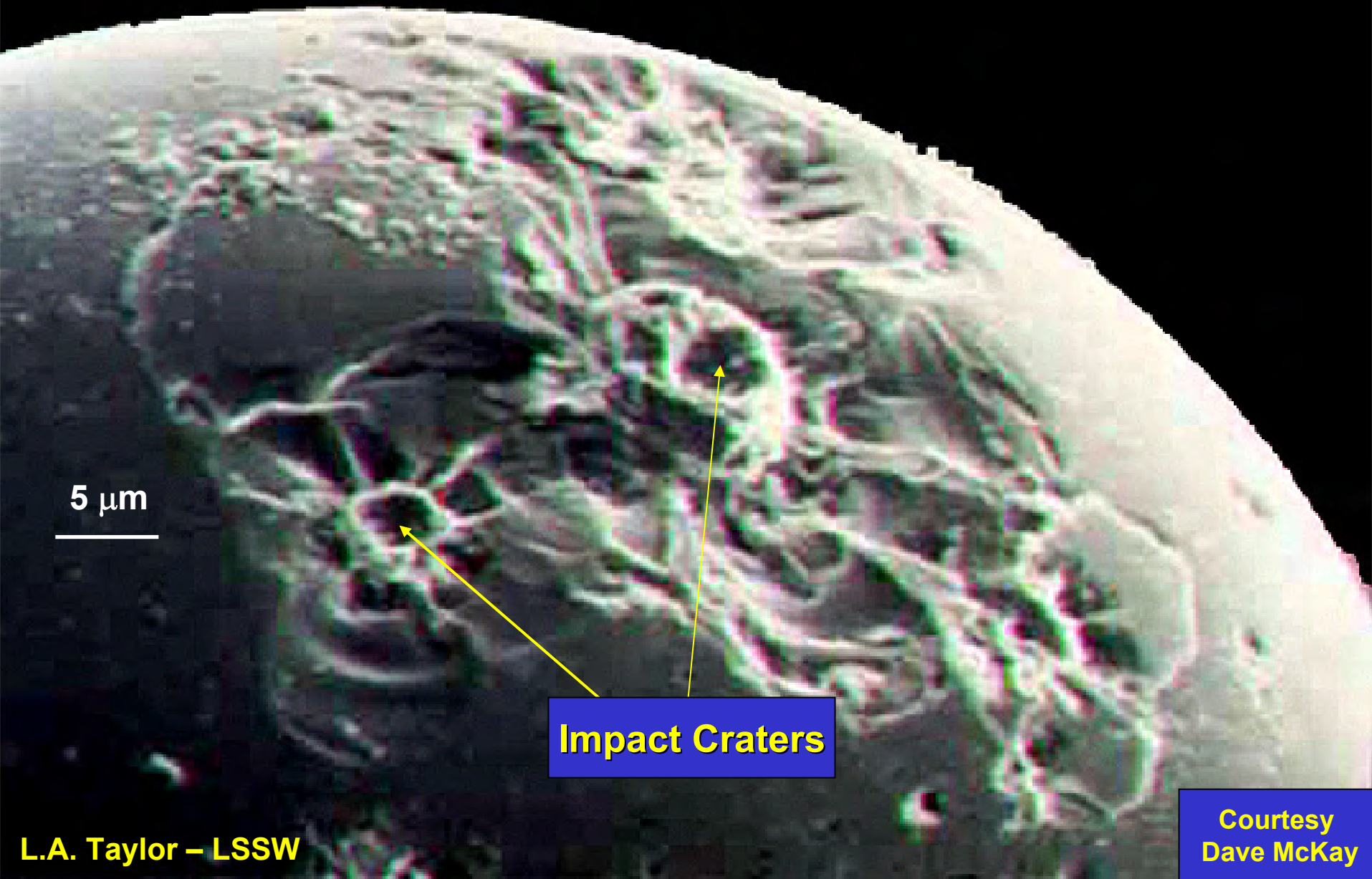
Cumulative Effects due to :

- **Deep Vacuum** : $>10^{-12}$ torr
- **Micrometeorite Impact** : Velocities to $>10^5$ km/hr
- **Radiation**: Galactic / Cosmic / Solar-ray Particles
- **Temperatures** : Equator: $-20\text{ }^{\circ}\text{C} \pm 120\text{ }^{\circ}\text{C}$
South Pole: Light $-50\text{ }^{\circ}\text{C} \pm 20\text{ }^{\circ}\text{C}$
Dark $-220\text{ }^{\circ}\text{C} \pm 20\text{ }^{\circ}\text{C}$

Our Unhappy Moon



Micrometeorite Impacts on Lunar a Glass Bead



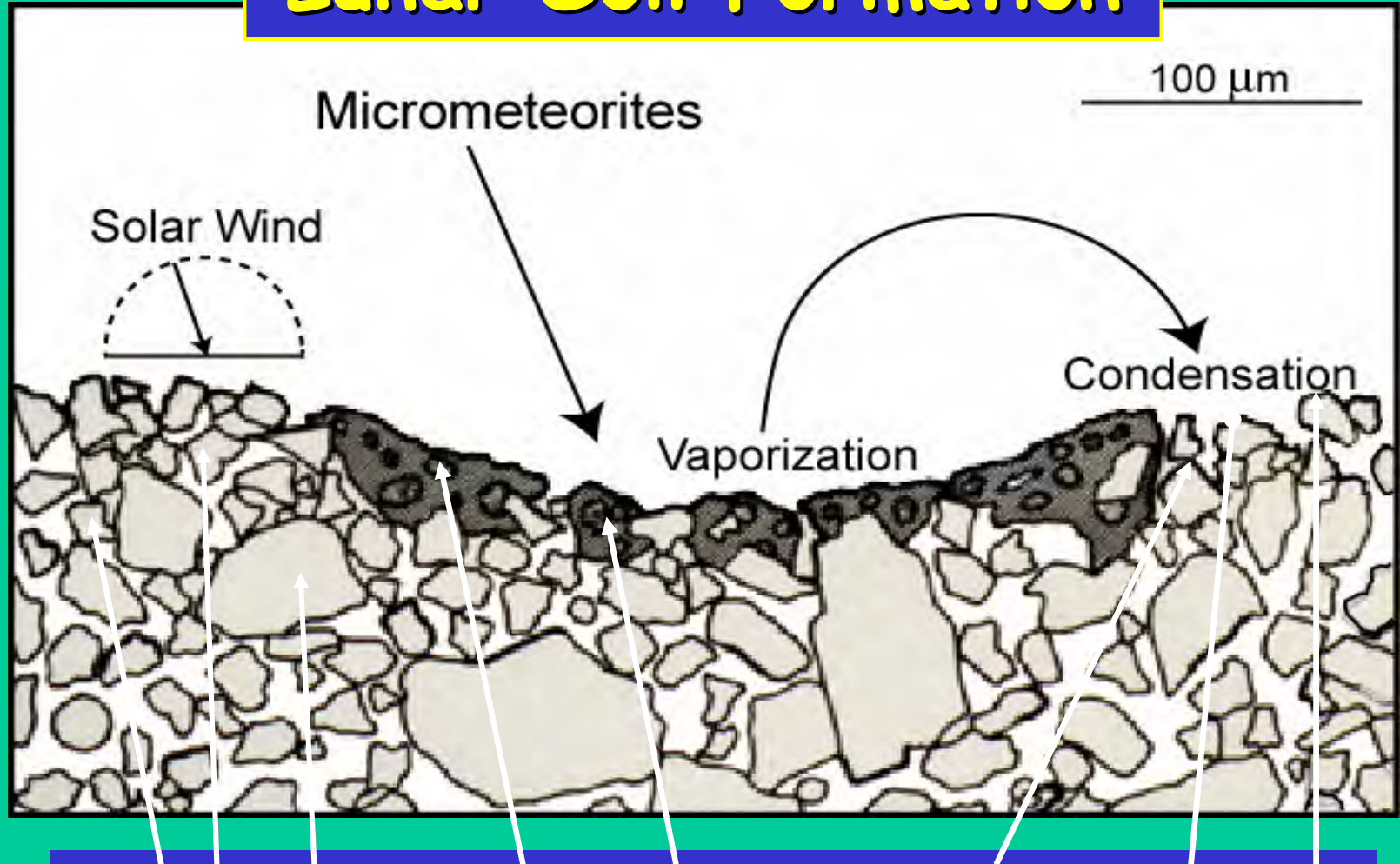
5 μm

Impact Craters

Courtesy
Dave McKay

L.A. Taylor – LSSW

Lunar Soil Formation



Comminution, Agglutination, & Vapor Deposition

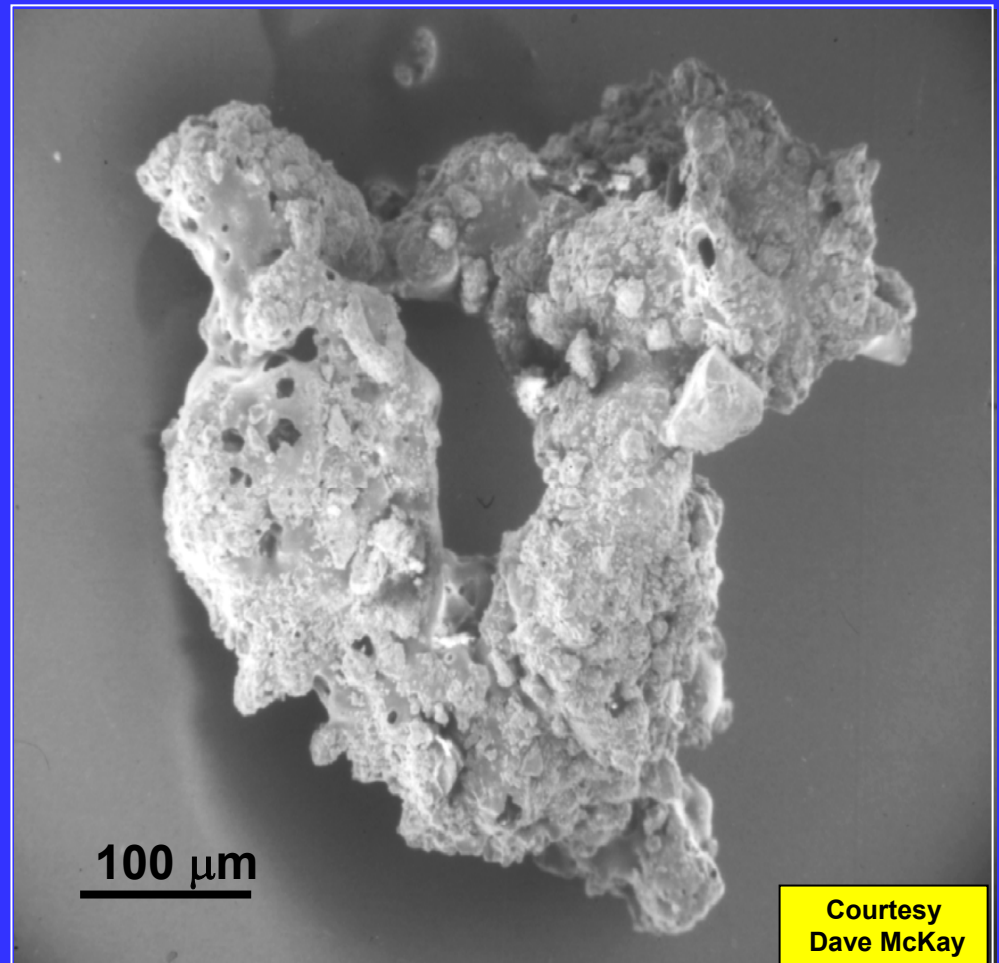
**The major Weathering and Erosional agent on the Moon
is Meteorite and Micrometeorite Impact.**

The only Weathering and Erosional agent on the Moon is Meteorite and Micrometeorite Impact.

BASIC PROCESSES IN LUNAR SOIL FORMATION

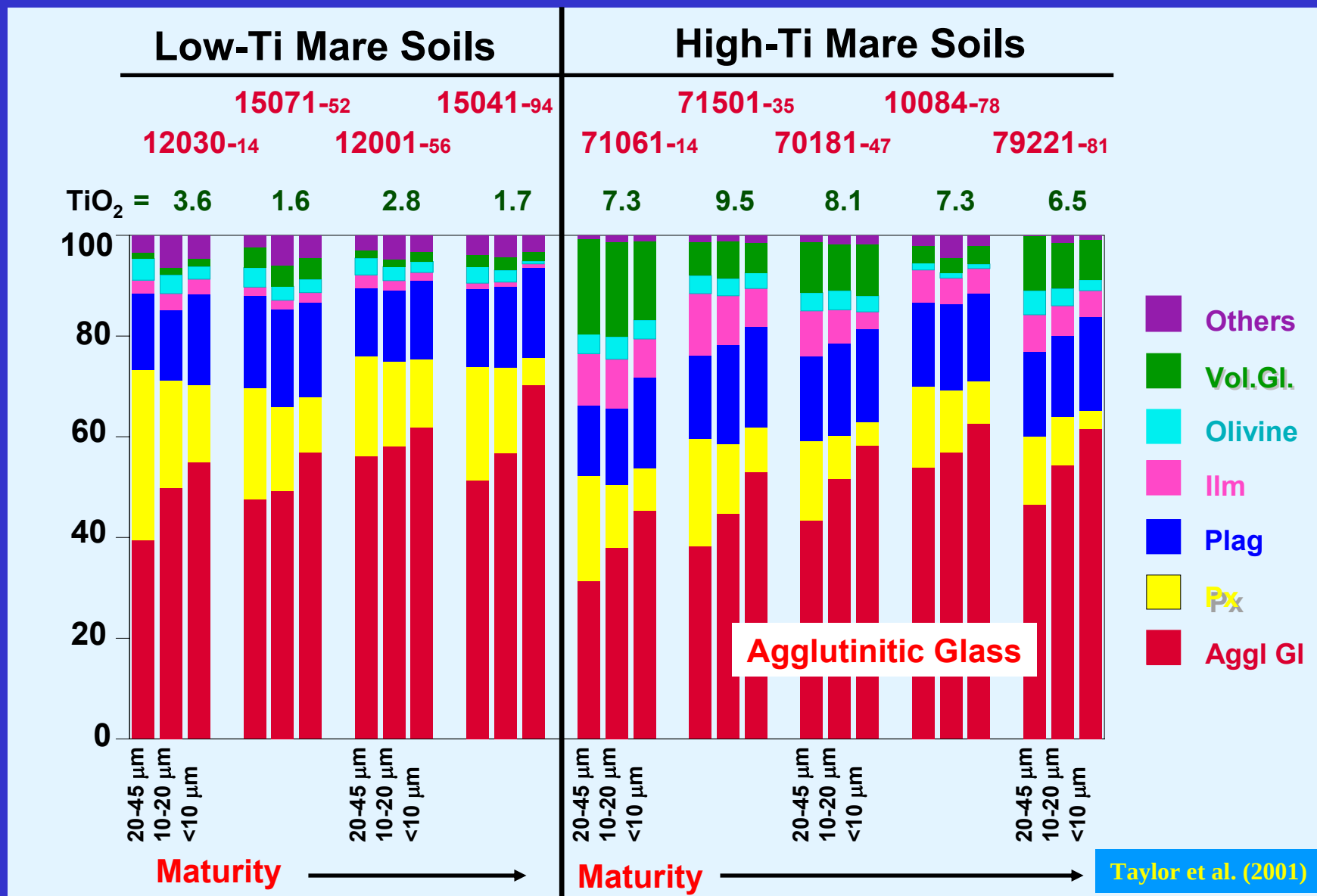
- **COMMINUTION:** breaking of rocks, minerals, and glasses into smaller particles;
- **AGGLUTINATION:** welding of rock, mineral, and glass fragments together by micrometeorite-produced, impact-generated melt (quenched to glass);
- **IMPACT-MELT VAPORIZATION AND DEPOSITION:**
Vaporization of volatile components in the micro-meteorite-produced, impact-generated melt.
- **SOLAR-WIND SPALLATION AND PARTICLE IMPLANTATION:**
Erosion and vaporization caused by sputtering from impacting high-energy particles;

Mare-Soil Agglutinate

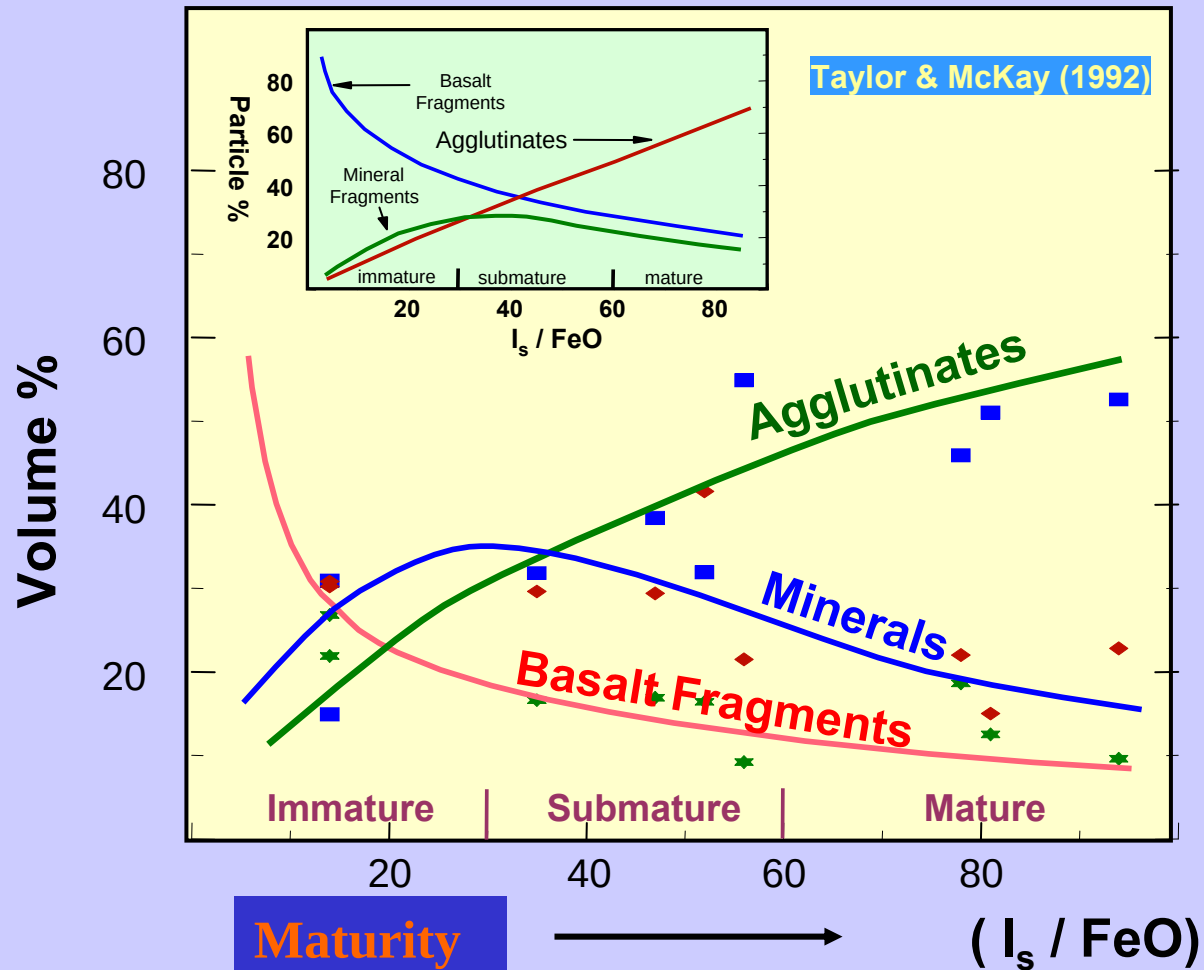


**Pieces of minerals, rocklets, and glass
cemented together by shock-melt glass**

Cumulative Modal Abundances of Mare Soils



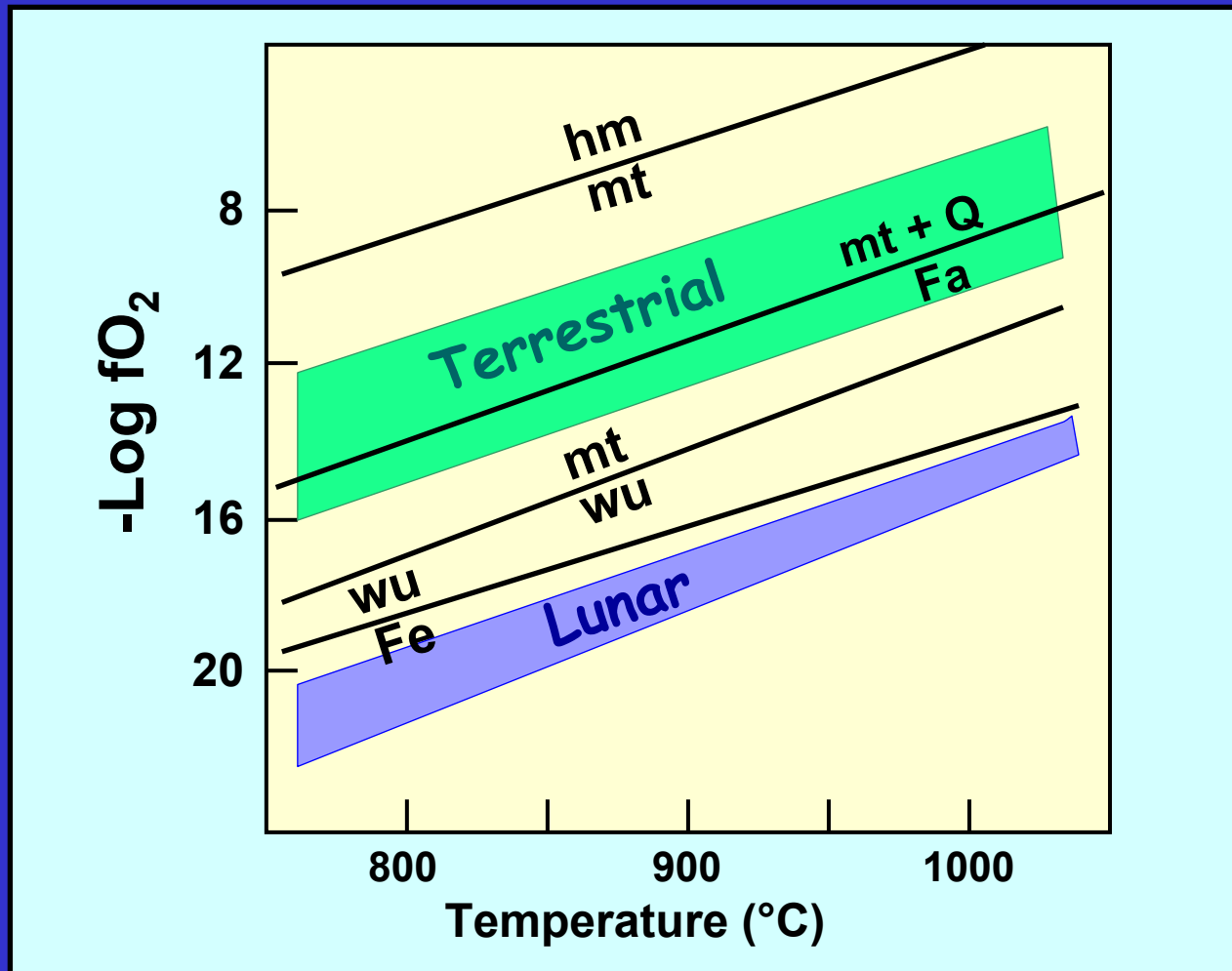
Mare Soil Maturation



**What are the differences between
Lunar Soils and the Rocks from
which they were derived ??**

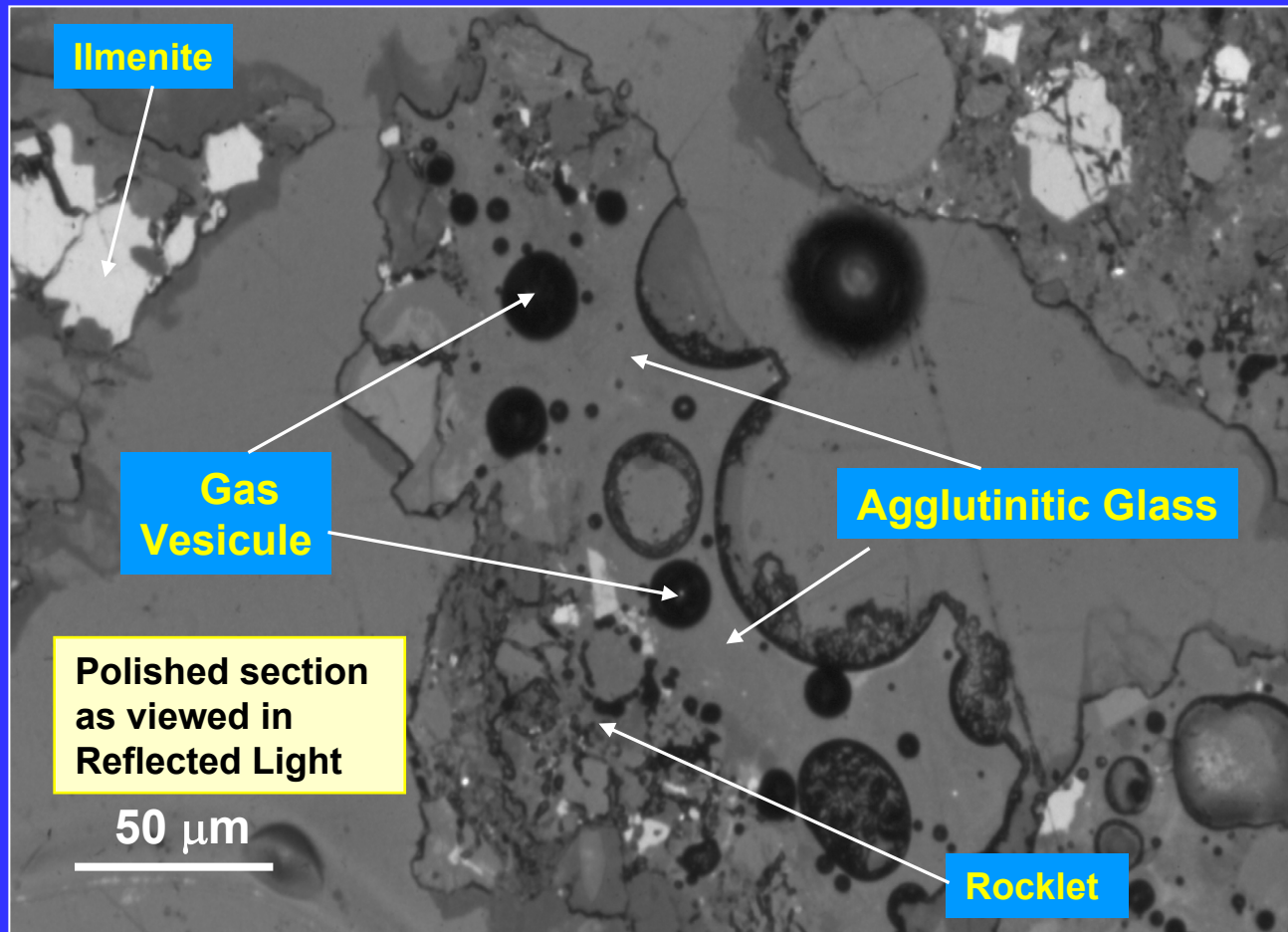
Major Difference = ~ 10 X more native Fe⁰ in the soil

Oxygen Partial Pressures



All Lunar Rocks have Elemental Fe^0 as a Stable Mineral

Mare-Soil Agglutinates

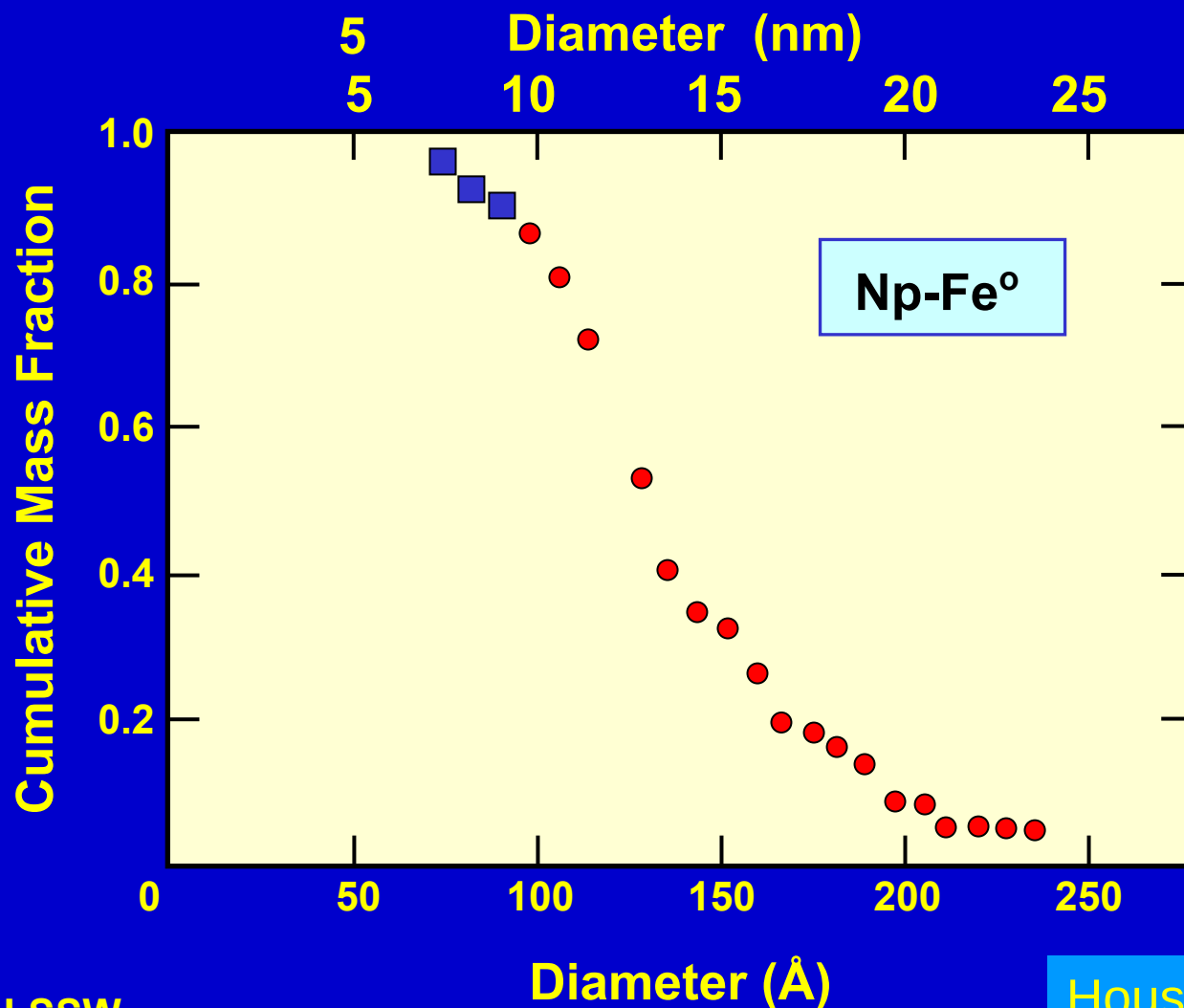


SEM BSE-Image of Mare Agglutinitic Glass



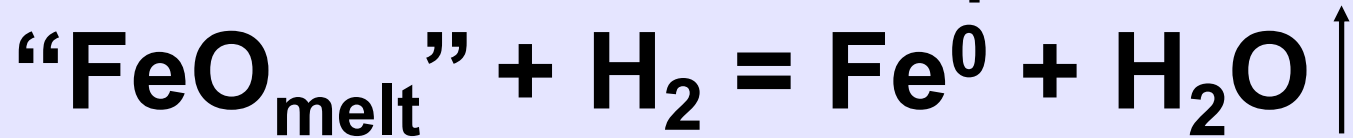
Courtesy – Dave McKay

TEM-measured Size Distribution of Fe Metal Spheres in Agglutinitic Glass of Apollo 11 Soil 10084



FORMATION OF NANOPHASE Fe^0 in AGGLUTINITIC GLASS:

Auto-Reduction Reaction in Impact-Soil Melt



Solar-Wind Implanted H^+ in Lunar Soil

Causes Reduction of Fe^{2+} to Fe^0

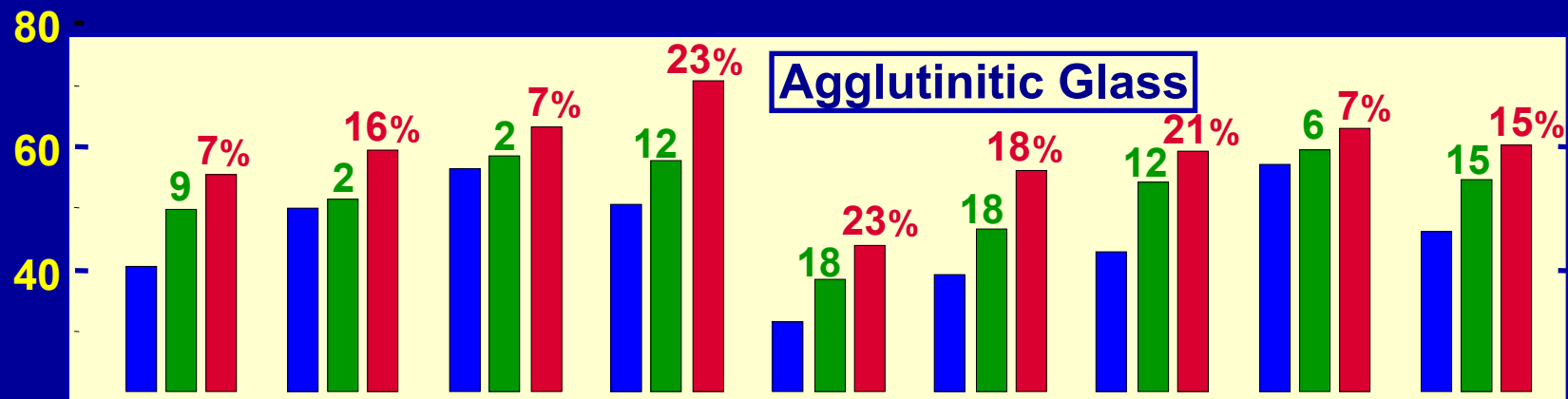
in Micrometeorite-Produced Impact Melt

FERROMAGNETIC RESONANCE (FMR):

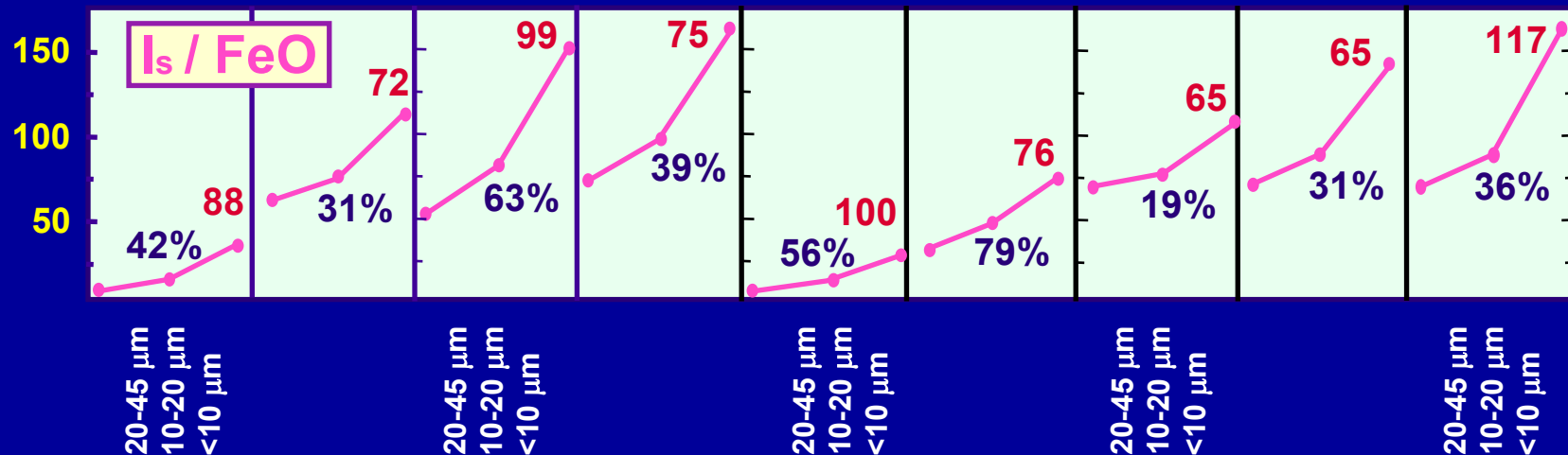
- Measurement of Single-Domain, Nanophase Fe^0 (I_s)
- Normalized I_s for Iron Content: I_s / FeO
- $I_s / \text{FeO} = \text{amount of total iron present as } \text{Fe}^0$
- I_s / FeO is a Function of Agglutinate Abundance
- Agglutinate Abundance is a Function of Maturity

$$I_s / \text{FeO} = \text{Soil Maturity}$$

Is/FeO Values Versus Agglutinitic Glass Contents

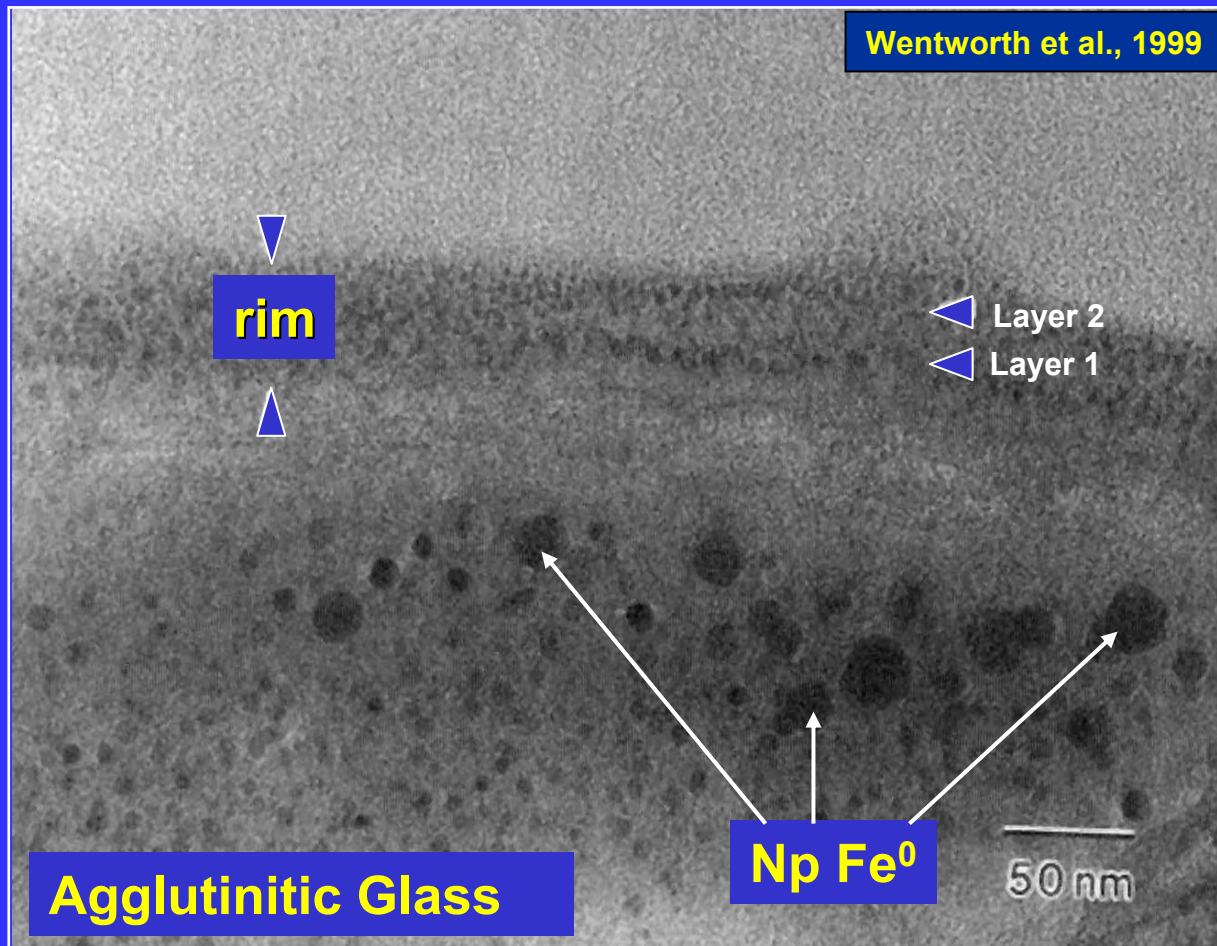


12030-14* 15071-52* 12001-56* 15041-94* 71061-14* 71501-35* 70181-47* 10084-78* 79221-81*

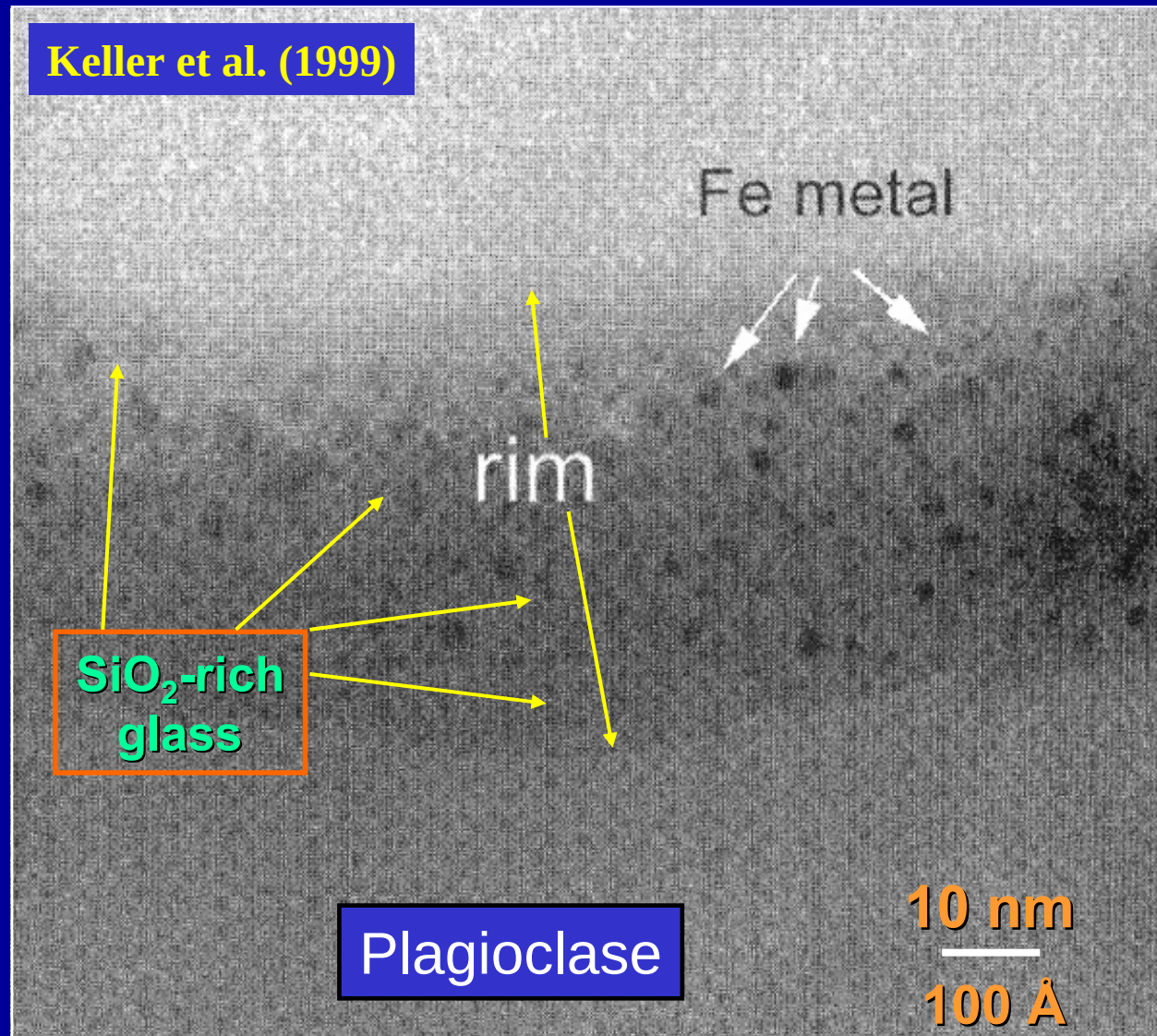


**Small increases in Agglutinitic Glass,
BUT Large increases in Is/FeO (Fe⁰)**

Multiple Layers of Vapor Deposition



Vapor-Deposited Nanophase Fe⁰ on Plagioclase

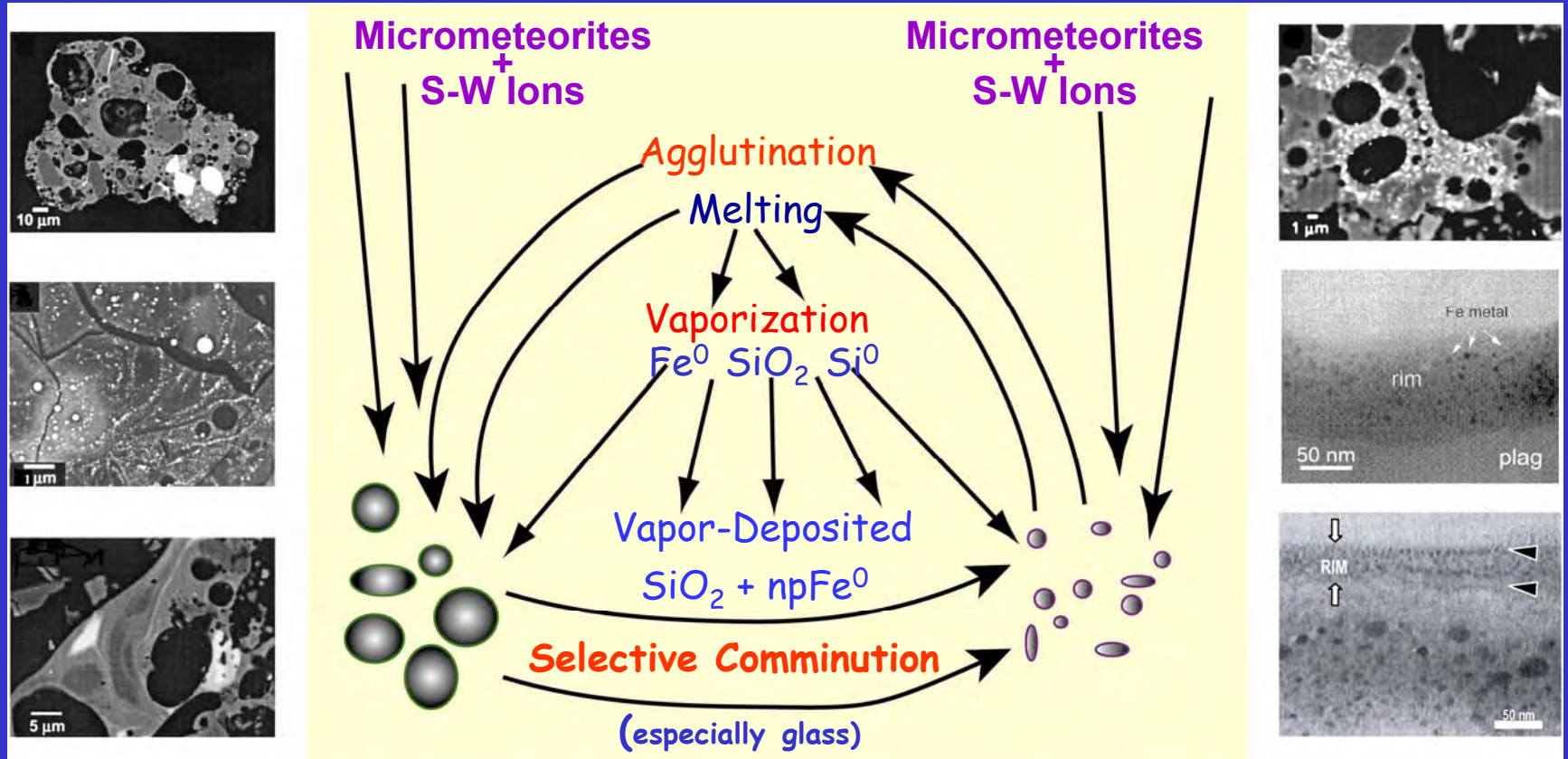


**NANOPHASE Fe^0 in VAPOR-DEPOSITED
PATINAS (RIMS) on VIRTUALLY ALL GRAINS of
a MATURE MARE SOIL PROVIDES an ADDITIONAL
and ABUNDANT SOURCE for the
GREATLY INCREASED I_s/FeO VALUES
for the FINEST GRAIN SIZES.**

**NANOPHASE Fe^0 in VAPOR-DEPOSITED RIMS
PROVIDES GREATLY INCREASED I_s / FeO VALUES
for the FINEST GRAIN SIZES.**

**The Amount of Ferromagnetic NP- Fe^0
Increases as the Grain Size Decreases.**

Lunar Soil Cycle

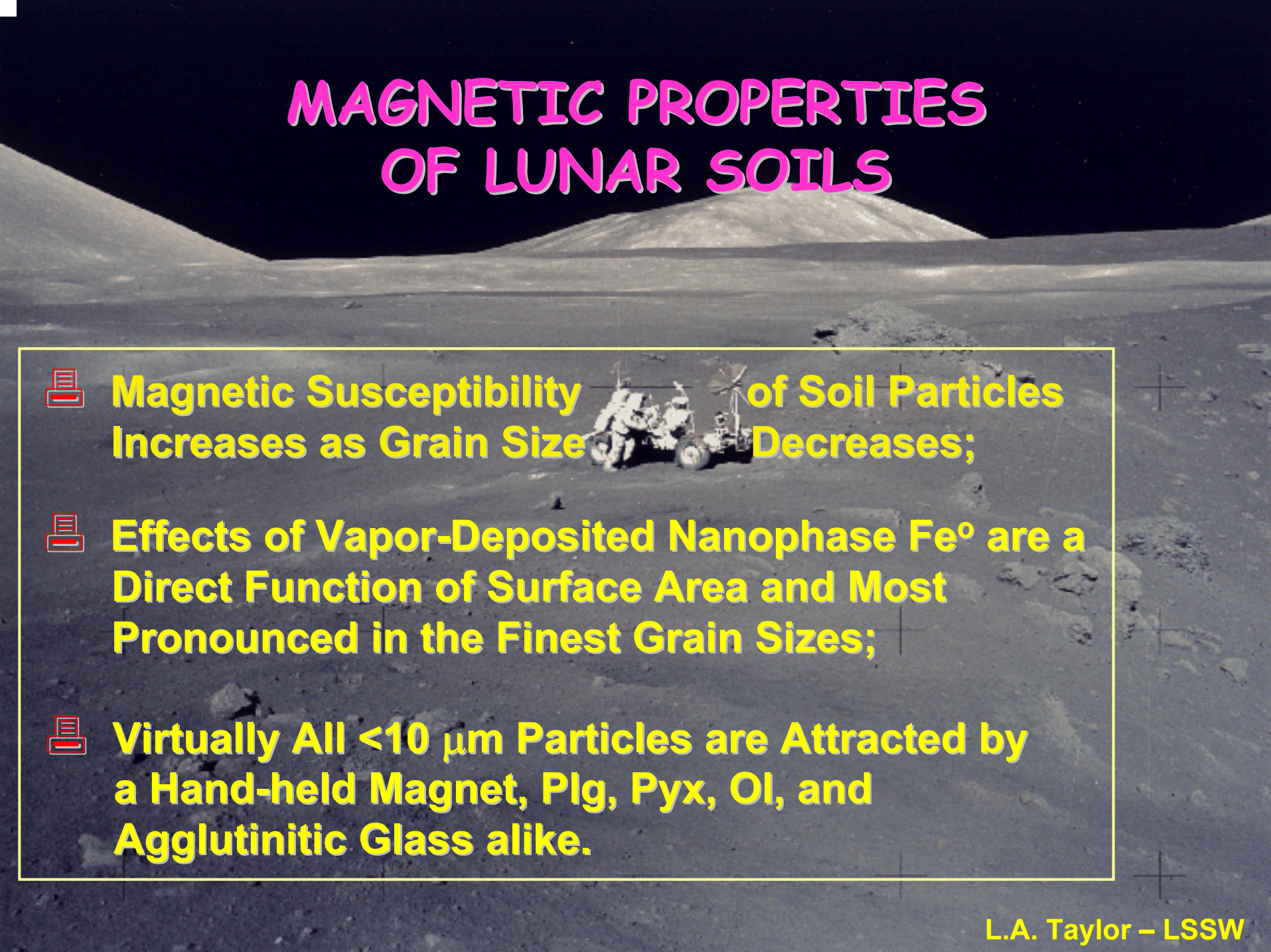


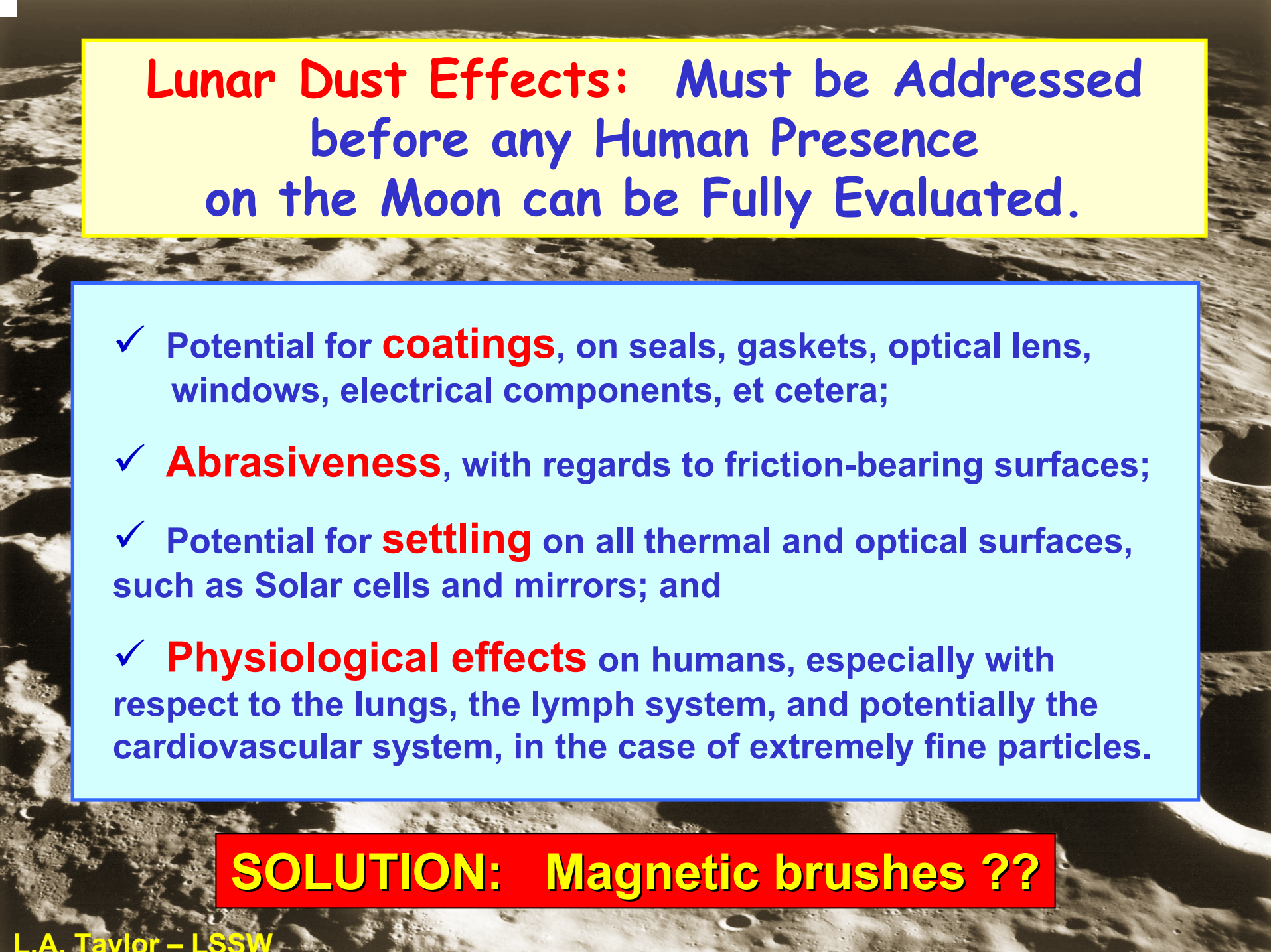
Course ($>100 \mu\text{m}$)

Fine ($<10 \mu\text{m}$)

Grain size $\xrightarrow{\text{decreases}}$

MAGNETIC PROPERTIES OF LUNAR SOILS

- 
- A photograph of the lunar surface, showing a dark, sandy terrain with a lunar rover in the background. The rover is a small, four-wheeled vehicle with a solar panel and an antenna. The background features a large, rounded lunar hill under a black sky.
- 📄 **Magnetic Susceptibility of Soil Particles Increases as Grain Size Decreases;**
 - 📄 **Effects of Vapor-Deposited Nanophase Fe⁰ are a Direct Function of Surface Area and Most Pronounced in the Finest Grain Sizes;**
 - 📄 **Virtually All <10 μm Particles are Attracted by a Hand-held Magnet, Plg, Pyx, Ol, and Agglutinitic Glass alike.**



Lunar Dust Effects: Must be Addressed before any Human Presence on the Moon can be Fully Evaluated.

- ✓ Potential for **coatings**, on seals, gaskets, optical lens, windows, electrical components, et cetera;
- ✓ **Abrasiveness**, with regards to friction-bearing surfaces;
- ✓ Potential for **settling** on all thermal and optical surfaces, such as Solar cells and mirrors; and
- ✓ **Physiological effects** on humans, especially with respect to the lungs, the lymph system, and potentially the cardiovascular system, in the case of extremely fine particles.

SOLUTION: Magnetic brushes ??

APOLLO 17



Jack



Gene

Microwave Heating of Lunar Soil:

NanoPhase Fe^0 in Silicate Glass

Fe^0 grain size is so small as to be below the effective “skin depth” of microwave penetration;

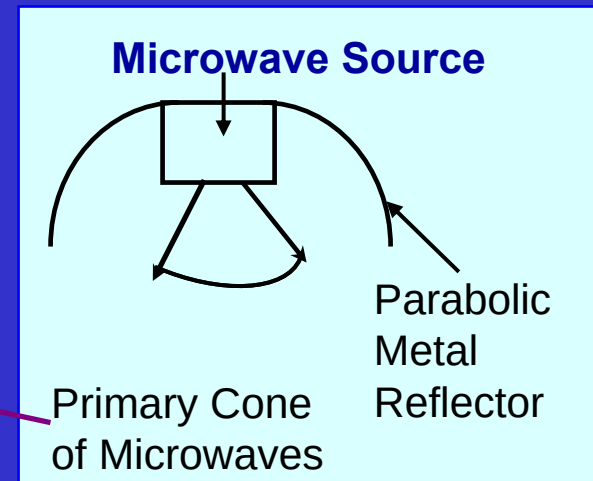
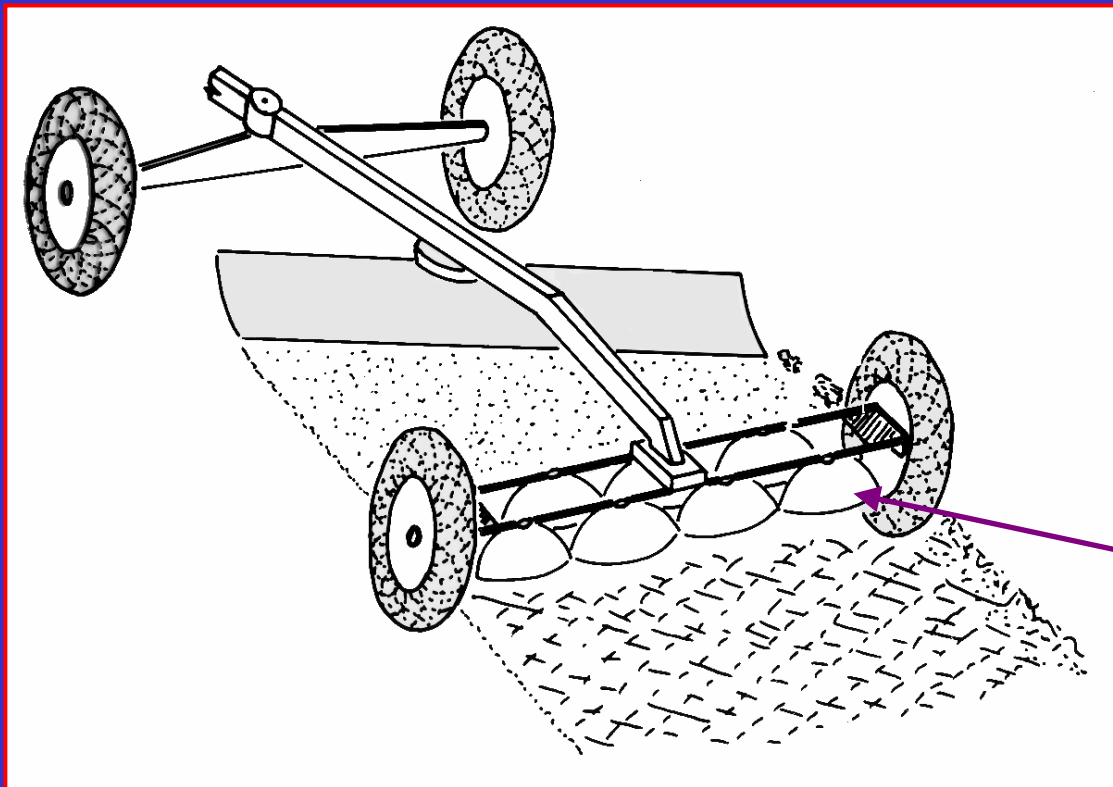
System is basically one of :
*Small conductors of Fe^0 insulated by
Intervening dielectric glass*

GREAT MICROWAVE COUPLING!

LUNAR SOIL PROCESSING & PRODUCTS

SINTERING and MELTING

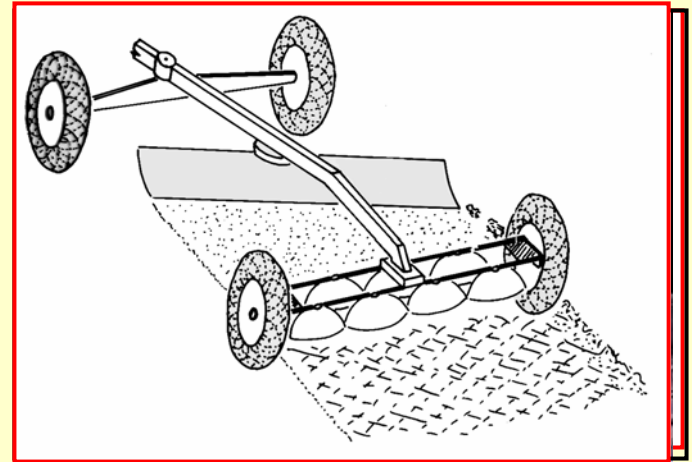
Creating Smooth-Sintered to Glassy Surfaces on the Moon



LUNAR SOIL PROCESSING & PRODUCTS

SINTERING and MELTING

- ➔ Roads
- ➔ Satellite Dishes
- ➔ Shielding
- ➔ Welding
- ➔ Recovery of Volatiles
- ➔ Glass Fiber Production
- ➔ Solar Cells (Ilmenite)



SUGGESTIONS ??



Polar Regions

Two cases:

❖ **Hydrogen enrichment: solar-wind hydrogen only**

Properties of regolith may be similar to elsewhere on the Moon, from our Apollo studies; new information may not be needed; possibility of unexpected effects of extreme cold (25-100 K); major considerations for processing tools at 50 K.

❖ **Hydrogen enrichment: H₂O-dominated ice**

In this case, physical properties of the regolith might be very different from those at Apollo sites

Concentrations of Solar-Wind Volatile Species in Lunar Regolith Samples, in ppm

	H	He	C	N	Ne	Ar
Apollo 11	20-100	20-84	96-216	45-110	2-11	1.3-12
Apollo 12	2-106	14-68	23-170	46-140	1.2-6	0.5-4.6
Apollo 14	67-105	5-16	42-225	25-130	0.14-1.6	0.4-2.2
Apollo 15	13-125	5-19	21-186	33-135	0.6-108	0.5-2.7
Apollo 16	4-146	3-36	31-280	4-209	0.4-1.2	0.6-3
Apollo 17	0.1-206	13-41	4-200	7-94	1.2-2.7	0.6-2.6

Haskin and Warren, 1991

Regolith Properties

■ Liquidus Temperature

- Maria: 1200 - 1250°C
- Highlands: 1350 - 1380°C

■ Viscosity (at liquidus)

- Maria: 9 Pa s (90 poise)
- Highlands: 5 Pa s (50 poise)

■ Bulk Density

- Upper few mm: 800-1000 kg/m³
- 10-20 cm: 1500-1800 kg/m³

■ Porosity: 35-45% at 10-20 cm depth

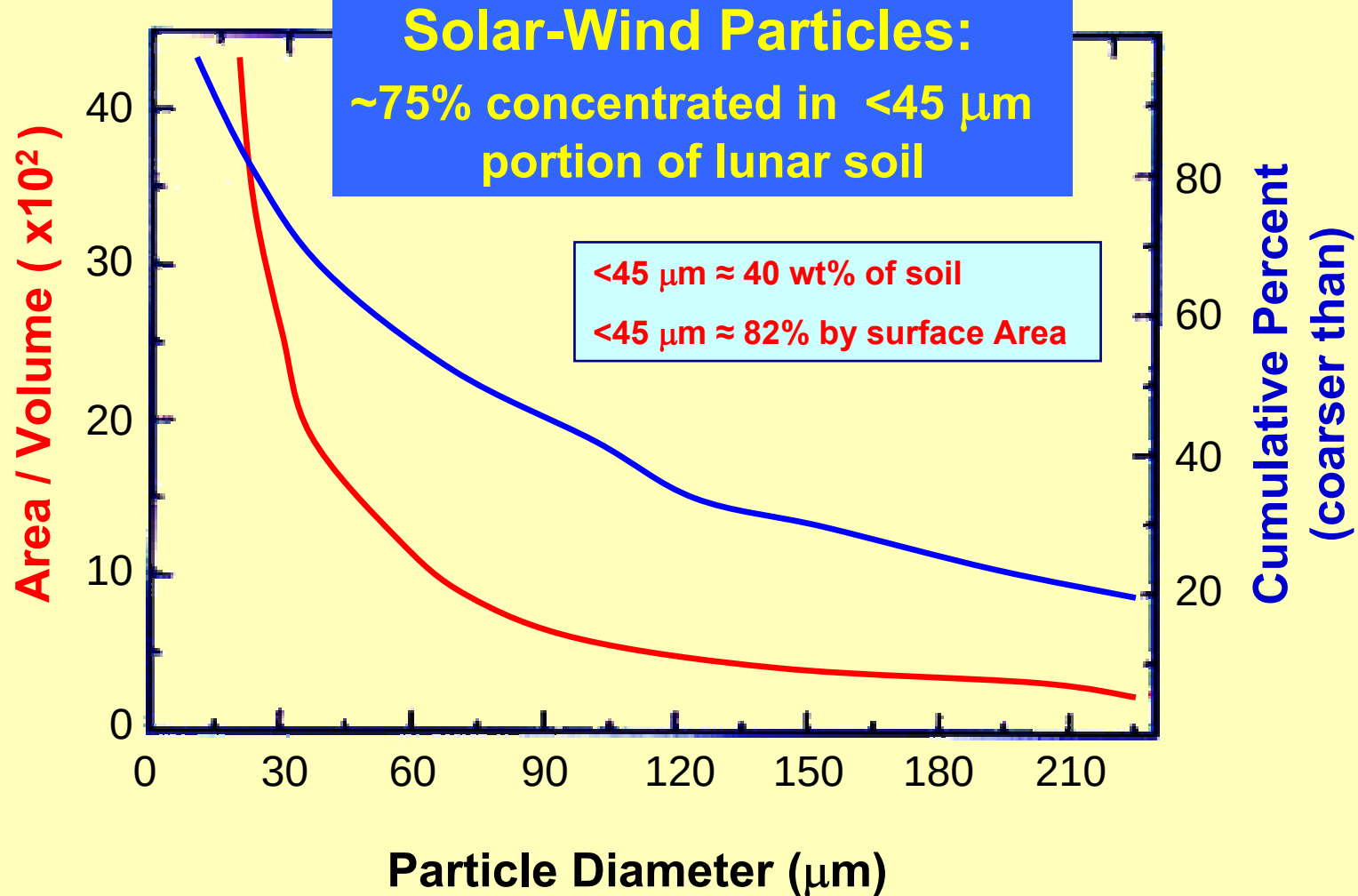
■ Thermal diffusivity

- At 30 cm: $0.7-1.0 \times 10^{-8} \text{ m}^2/\text{s}$
- Very good insulator

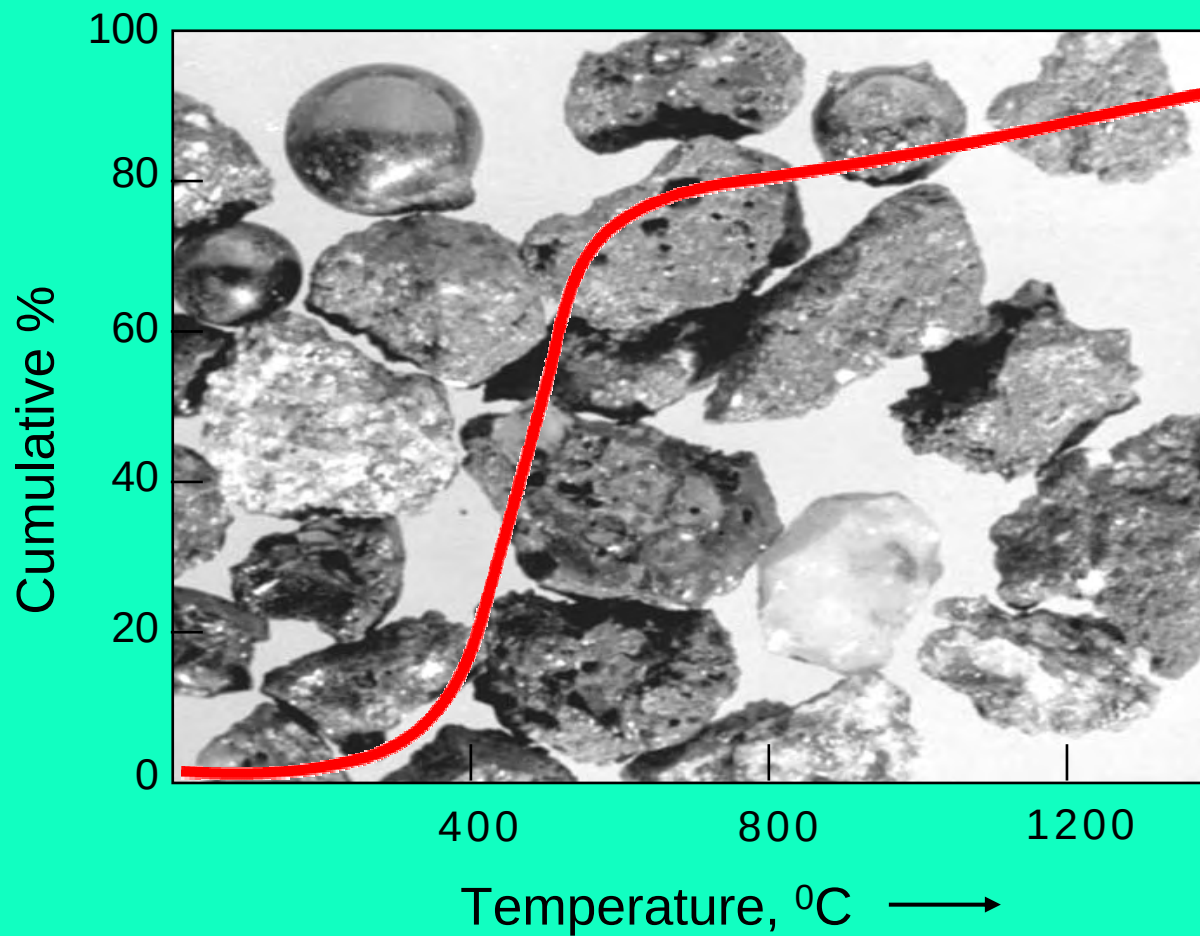
■ CONCENTRATIONS OF SOLAR-WIND PARTICLES

- H, He, C, N;
- Released as gases (H₂, O₂, CO₂, N₂, H₂O) when soil is **agitated** and/or heated;
- **Perhaps as much as 40-50% of SW lost by manipulation alone.**
- Peak gas release between 500 and 700°C

Lunar Mare Soils



Gas Release from Lunar Mare Soil



Abundance of Hydrogen in Mare Soil

Premise : Need 20 tonnes of LLH hydrogen per year

Hydrogen in lunar soil = 200 ppm; 50% recovery = 0.01 wt%

H in 1 m³ = [2.0 g/cc] x [10⁶ cc/m³] X [1 x 10⁻⁴] = **2.0 x 10² g/m³**

20 tonnes = 20t x 10³ kg / t x 10³ g/kg = 20 x 10⁶ g

20 tonnes = 20 x 10⁶g / 2.0 x 10² g/m³ = 10 x 10⁴ m³ = **10⁵ m³**

1 Football Field (Depth of 3m) = 5 x 10³ m² x 3 = **15 x 10³ m³**

20 t LH = 6.2 Football Fields to 3 m depth

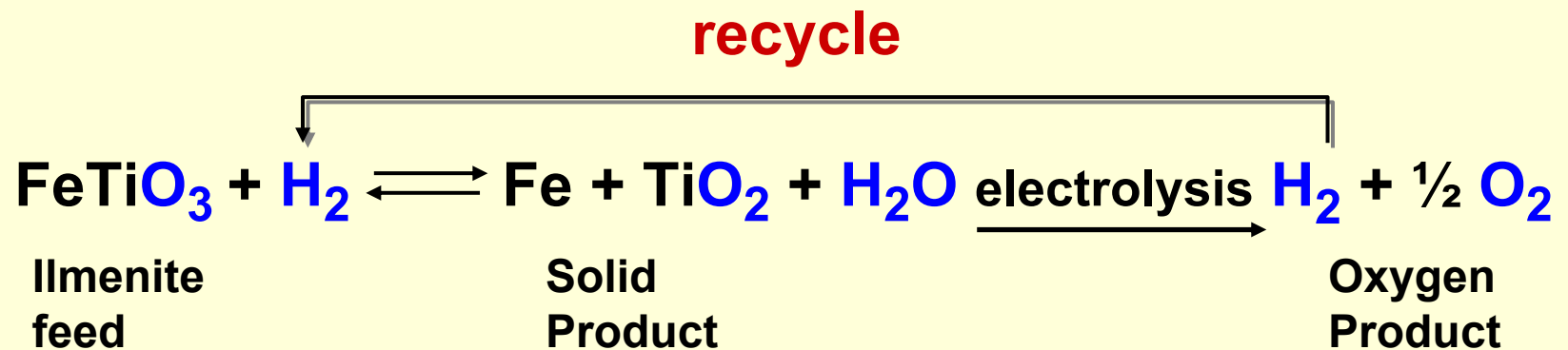
20 tonnes of LLH = ~6 Football Fields
0.03 km² (1 / 30th km²)

= 180 TONNES OF WATER

Potential Processes for Oxygen Production on the Moon

Process	Reference
<u>Solid/Gas Interaction</u> X Ilmenite Reduction with Hydrogen Ilmenite Reduction with C/CO Ilmenite Reduction with Methane x Glass Reduction with Hydrogen Reduction with Hydrogen Sulfide Extraction with Fluorine X Carbothermal Reduction Chlorine Plasma Reduction	<u>Solid/Gas Interaction</u> Gibson & Knudsen (1985) Chang (1959); Shadman & Zhou (1988) Friedlander (1985) McKay et al. (1991) Dalton & Hohmann (1972) Burt (1988) Lynch (1989) Lynch (1989)
<u>Silicate/Oxide Melt</u> X Molten Silicate Electrolysis x Fluxed Molten Silicate Electrolysis Caustic Dissolution & Electrolysis Carbothermal Reduction Magma Partial Oxidation Li or Na Reduction Ilmenite	<u>Silicate/Oxide Melt</u> Haskin (1985) Keller (1986) Dalton & Hohmann (1972) Rosenburg (1966); Cutler & Krag (1985) Waldon (1989) Sammells & Semkow (1987)
<u>Pyrolysis</u> Vapor Pyrolysis Ion Plasma Pyrolysis Plasma Reduction of Ilmenite	<u>Pyrolysis</u> Steurer & Nerad (1983) Steurer & Nerad (1983) Allen et al., (1988)
<u>Aqueous Solutions</u> HF Acid Dissolution H ₂ SO ₄ Acid Dissolution	<u>Aqueous Solutions</u> Waldron (1985) Christiansen et al. (1988); Sullivan (1991)
<u>Co-Product Recovery</u> Hydrogen-Helium-Water from Soil.	<u>Co-Product Recovery</u> Christiansen et al. (1988)

Production of LLOX by H₂ Reduction of Ilmenite



Mining Water on the Moon Living- Off the-Land Concept

- H_2 is abundant in the Lunar Soil !
- O_2 is easily produced from Lunar Resources !
- Therefore, we can readily obtain rocket fuels and make H_2O !!!

BUSH FULL-SERVICE GAS STATION: LLOX & LLH

- At BMB *
- L1 Station,
- LEO,
- Soon in a neighborhood near you!

* BUSH MOON BASE



STRAWMAN LIST OF SIMULANTS NEEDED FOR STUDIES OF:

	<u>Chemistry</u>	<u>Geotech/Engr</u>	<u>Simulant</u>
Facilities Construction		XX	JSC-2; MLS-2
Regolith Digging and Movin		XX	JSC-2; MLS-2
Trafficability (e.g., Roads)	XX	XX	JSC-2; MLS-2
Microwave Processing	X	X	NP-1; JSC-2; MLS-2
Conventional Heat Treatment	X	X	JSC-2; MLS-2
Oxygen Production	X	X	JSC-2; MLS-1; MLS-2
Dust Abatement	X	X	NP-1; JSC-2
Mineral Beneficiation	X	X	???
Solar-Wind Gas Release	XX	X	JSC-2; MLS-2; NP-1
Cement Manufacture	X		JSC-2; MLS-2
Radiation Protection		X	JSC-2; MLS-1; MLS-2

Mare Soil: **JSC-2** = JSC-1 in chemistry + Geotech Prop.
 MLS-1 = Chemistry only of Apollo 11 soil (no glass)
Highland Soil: **MLS-2** = Anorthosite = Chemistry only
Magnetic Soil: **NP-1** = Nanophase Magnetic properties only

JSC = Johnson Space Center; MLS = Minnesota Lunar Simulant