Reduced Order Modeling for Non-equilibrium Radiation Hydrodynamics of Base Flow and Wakes: Enabling Manned Missions to Mars

PI: Dr. Marco Panesi (University of Illinois, UIUC) Team Members: (NASA AMES and LaRC)

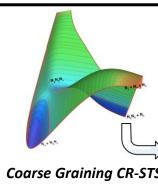
Drs. C.O. Johnston (*LaRC*), R.L. Jaffe (*ARC*), A. Wray (*ARC*), Y. Liu (*ARC*)

Approach: *Coarse-Grain Method* The methodology of reduction for **Collisional and Radiative STS** models consists of two distinct steps:

- Local Representation and Reconstruction: lumping of the internal energy levels in energy groups, and the reconstruction of population of each group, using macroscopic quantities.
- Macroscopic Moment Equations and Rate Coefficients: macroscopic governing equations are obtained by taking moments of the master equations.

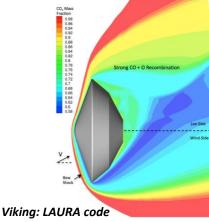
Radiation and Kinetics Coupling

- Operator Splitting Method: enabling different grouping strategies for Radiation and Kinetics.
- Semi-Implicit Time integration Validation/Calibration (EAST DATA)
- **Bayes Inference:** Calibration of the reduced model.



2. Coarse Graining CR-STS Model:

Grouping of ground and electronic excited states, **stochastic** and **deterministic** Master equation.



Research Objective:

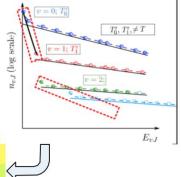
- To reduce predictive uncertainty in the heating predictions for the back shell region, due to CO₂ IR radiation.
- Construct a computationally efficient physics/chemistry model for CFD that is physics based and accurate for Earth (10-16km/s) and Mars (5-8km/s) entries.
- Develop a new framework for the construction of **Coarse Grained Collisional and Radiative State To State model** to enable their application to current CFD models.
 - *Time accurate, tight coupling* between radiation and chemistry without relying on the QSS assumption.
 - The starting TRL is 1, and the final is TL is 3.

Potential Impact:

- **Predictive modeling** of strong non-equilibrium environments (Backshell).
- Modeling of Entry heating for all hypervelocity return missions such as OSIRIS-Rex and Insight, MAV, Lunar/Asteroid Sample Return & Crewed asteroid rendezvous.
- This innovative approach will transform the way nonequilibrium flows can be understood and modeled by the community.

1. Ab-initio Calculation of Kinetic parameters

Develop a new model for Mars Chemistry and Radiation **CO** and **CO**₂ **STS** Kinetics and Radiation



- 3. Application to CFD: Different Grouping strategies for radiation and kinetics. US3D or LAURA.
- 4. Calibration and Validation (V&V and UQ) Using EAST data, and Bayes inference