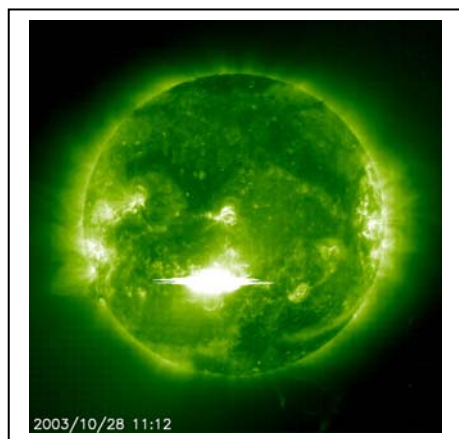
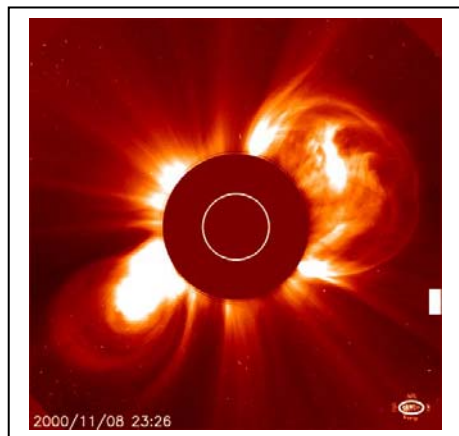


Do Fast CMEs produce intense SPEs?



The sun produces two basic kinds of storms; coronal mass ejections (SOHO satellite: top left) and solar flares (SOHO satellite: bottom left). These are spectacular events in which billions of tons of matter are launched into space (CMEs) and vast amounts of electromagnetic energy are emitted (Flares). A third type of 'space weather storm' can also occur.

Solar Proton Events (SPEs) are invisible, but intense, showers of high-energy particles near Earth that can invade satellite electronics and cause serious problems, even malfunctions and failures. Some of the most powerful solar flares can emit these particles, which streak to Earth within an hour of the flare event. Other SPE events, however, do not seem to arrive at Earth until several days later.

Here is a complete list of Solar Proton Events between 1976-2005: <http://umbra.nascom.nasa.gov/SEP/>

Here is a complete list of coronal mass ejections 1996 - 2006: http://cdaw.gsfc.nasa.gov/CME_list/

Between January 1, 1996 and June 30, 2006 there were 11,031 CMEs reported by the SOHO satellite. Of these, 1186 were halo events. Only half of the halo events are actually directed towards Earth. The other half are produced on the far side of the sun and directed away from Earth. During this same period of time, 90 SPE events were recorded by GOES satellite sensors orbiting Earth. On the next page, is a list of all the SPE events and Halo CMEs that corresponded to the SPE events. There were 65 SPEs that coincided with Halo CMEs. Also included is the calculated speed of the CME event.

From the information above, and the accompanying table, draw a Venn Diagram to represent the data, then answer the questions below.

- Question 1: A) What percentage of CMEs detected by the SOHO satellite were identified as Halo Events?
 B) What are the odds of seeing a halo Event?
 C) How many of these Halo events are directed towards Earth?

- Question 2: A) What fraction of SPEs were identified as coinciding with Halo Events?
 B) What are the odds that an SPE occurred with a Halo CME?
 C) What fraction of all halo events directed towards earth coincided with SPEs?

- Question 3: A) What percentage of SPEs coinciding with Halo CMEs are more intense than 900 PFUs?
 B) What are the odds that, if you detect a 'Halo- SPE', it will be more intense than 900 PFUs?

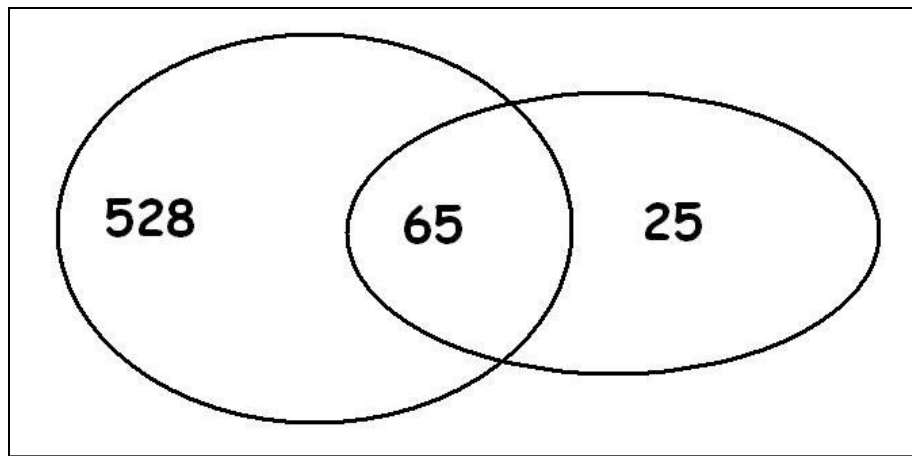
- Question 4: A) What percentage of Halo-SPEs have speeds greater than 1000 km/sec?
 B) What are the odds that a Halo-SPE in this sample has a speed of > 1000 km/sec?

Question 5: From what you have calculated as your answers above, what might you conclude about Solar Proton Events and CMEs? How would you use this information as a satellite owner and operator?

Data Tables showing dates and properties of Halo CMEs and Solar Proton Events.

| Date | CME Speed (km/s) | SPE (pfu) | Date | CME Speed (km/s) | SPE (pfu) |
|--------------------|------------------|-----------|--------------------|------------------|-----------|
| November 4, 1997 | 785 | 72 | January 8, 2002 | 1794 | 91 |
| November 6, 1997 | 1556 | 490 | January 14, 2002 | 1492 | 15 |
| April 20, 1998 | 1863 | 1700 | February 20, 2002 | 952 | 13 |
| May 2, 1998 | 938 | 150 | March 15, 2002 | 957 | 13 |
| May 6, 1998 | 1099 | 210 | March 18, 2002 | 989 | 19 |
| May 3, 1999 | 1584 | 14 | March 22, 2002 | 1750 | 16 |
| June 1, 1999 | 1772 | 48 | April 17, 2002 | 1240 | 24 |
| June 4, 1999 | 2230 | 64 | April 21, 2002 | 2393 | 2520 |
| February 18, 2000 | 890 | 13 | May 22, 2002 | 1557 | 820 |
| April 4, 2000 | 1188 | 55 | July 15, 2002 | 1151 | 234 |
| June 6, 2000 | 1119 | 84 | August 14, 2002 | 1309 | 24 |
| June 10, 2000 | 1108 | 46 | August 22, 2002 | 998 | 36 |
| July 14, 2000 | 1674 | 24000 | August 24, 2002 | 1913 | 317 |
| July 22, 2000 | 1230 | 17 | September 5, 2002 | 1748 | 208 |
| September 12, 2000 | 1550 | 320 | November 9, 2002 | 1838 | 404 |
| October 16, 2000 | 1336 | 15 | May 28, 2003 | 1366 | 121 |
| October 25, 2000 | 770 | 15 | May 31, 2003 | 1835 | 27 |
| November 8, 2000 | 1738 | 14800 | June 17, 2003 | 1813 | 24 |
| November 24, 2000 | 1289 | 940 | October 26, 2003 | 1537 | 466 |
| January 28, 2001 | 916 | 49 | November 4, 2003 | 2657 | 353 |
| March 29, 2001 | 942 | 35 | November 21, 2003 | 494 | 13 |
| April 2, 2001 | 2505 | 1100 | April 11, 2004 | 1645 | 35 |
| April 10, 2001 | 2411 | 355 | July 25, 2004 | 1333 | 2086 |
| April 15, 2001 | 1199 | 951 | September 12, 2004 | 1328 | 273 |
| April 18, 2001 | 2465 | 321 | November 7, 2004 | 1759 | 495 |
| April 26, 2001 | 1006 | 57 | January 15, 2005 | 2861 | 5040 |
| August 9, 2001 | 479 | 17 | July 13, 2005 | 1423 | 134 |
| September 15, 2001 | 478 | 11 | July 27, 2005 | 1787 | 41 |
| September 24, 2001 | 2402 | 12900 | August 22, 2005 | 2378 | 330 |
| October 1, 2001 | 1405 | 2360 | | | |
| October 19, 2001 | 901 | 11 | | | |
| October 22, 2001 | 618 | 24 | | | |
| November 4, 2001 | 1810 | 31700 | | | |
| November 17, 2001 | 1379 | 34 | | | |
| November 22, 2001 | 1437 | 18900 | | | |
| December 26, 2001 | 1446 | 779 | | | |

Note: Solar Proton Event strengths are measured in the number of particles that pass through a square centimeter every second, and is given in units called Particle Flux Units or PFUs.



Question 1: A) What percentage of CMEs detected by the SOHO satellite were identified as Halo Events?
 $1186/11031 = 11\%$

B) What are the odds of seeing a halo Event?
 $1 / 0.11 = 1 \text{ chance in } 9$

C) How many of these Halo events are directed towards Earth?
 From the text, only half are directed to Earth so $1186/2 = 593$ Halos.

Question 2: A) What fraction of SPEs were identified as coinciding with Halo Events?
 $65 \text{ table entries} / 90 \text{ SPEs} = 72\%$

B) What are the odds that an SPE occurred with a Halo CME?
 $1 / 0.72 = 1 \text{ chance in } 1.38 \text{ or about } 2 \text{ chances in } 3$

C) What fraction of all halo events directed towards Earth coincided with SPEs?
 $65 \text{ in Table} / (528+65) \text{ Halos} = 11\%$

Question 3: A) What percentage of SPEs coinciding with Halo CMEs are more intense than 900 PFUs?
 From the table, there are 12 SPEs out of 65 in this list or $12/65 = 18\%$

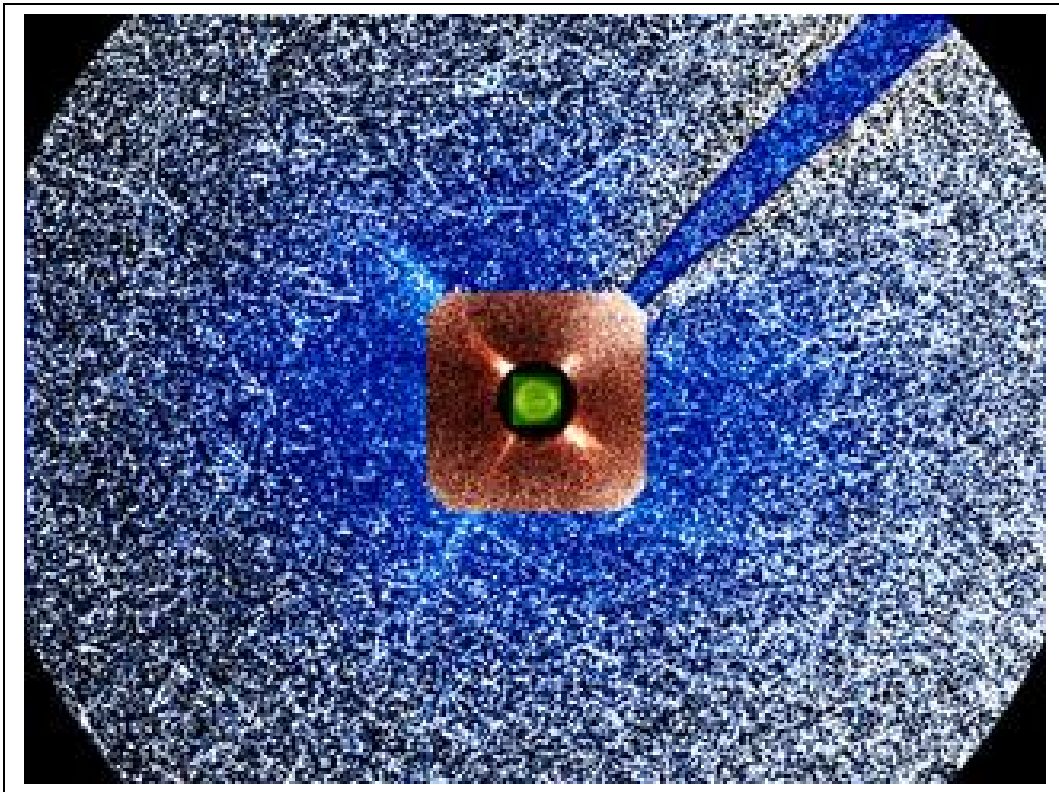
B) What are the odds that, if you detect a 'Halo- SPE', it will be more intense than 900 PFUs?
 $1 / 0.18 = 1 \text{ chance in } 5.$

Question 4: A) What percentage of Halo-SPEs have speeds greater than 1000 km/sec?
 There are 50 out of 65 or $50/65 = 77\%$

B) What are the odds that a Halo-SPE in this sample has a speed of > 1000 km/sec?
 $1 / 0.77 = 1 \text{ chance in } 1.3 \text{ or } 2 \text{ chances in } 3.$

Question 5: From what you have calculated as your answers above, what might you conclude about Solar Proton Events and CMEs? How would you use this information as a satellite owner and operator?

A reasonable student response is that Halo CMEs occur only 11% of the time, and of the ones directed towards Earth only 1 out of 9 coincide with SPEs. However, in terms of SPEs, virtually all of the SPEs coincide with Halo events (2 out of 3) and SPEs are especially common when the CME speed is above 1000 km/sec. As a satellite owner, I would be particularly concerned if scientists told me there was a halo CME headed towards Earth AND that it had a speed of over 1000 km/sec. Because the odds are now 2 chances out of 3 that an SPE might occur that could seriously affect my satellite. I would try to put my satellite in a safe condition to protect it from showers of high-energy particles that might damage it.



The January 20, 2005 solar proton event (SPE) was by some measures the biggest since 1989. It was particularly rich in high-speed protons packing more than 100 million electron volts (100 MeV) of energy. Such protons can burrow through 11 centimeters of water. A thin-skinned spacesuit would have offered little resistance, and the astronaut would have been radiation poisoned, and perhaps even killed.

The above image was taken by the SOHO satellite during this proton storm. The instrument, called LASCO, was taking an image of the sun in order for scientists to study the coronal mass ejection (CME) taking place. Each of the individual white spots in the image is a track left by a high-speed proton as it struck the imaging CCD (similar to the 'chip' in your digital camera). As you see, the proton tracks corrupted the data being taken.

The high-speed particles from these proton storms also penetrate satellites and can cause data to be lost, or even false commands to be given by on-board computers, causing many problems for satellite operators.