



Mars Exploration Program

- A Pathway to Future Missions –

Presented at the 1st Human Landing Site Workshop

October 29, 2015

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#JOURNEYTOMARS

Mars – a Frontier Opened by Science Precursors

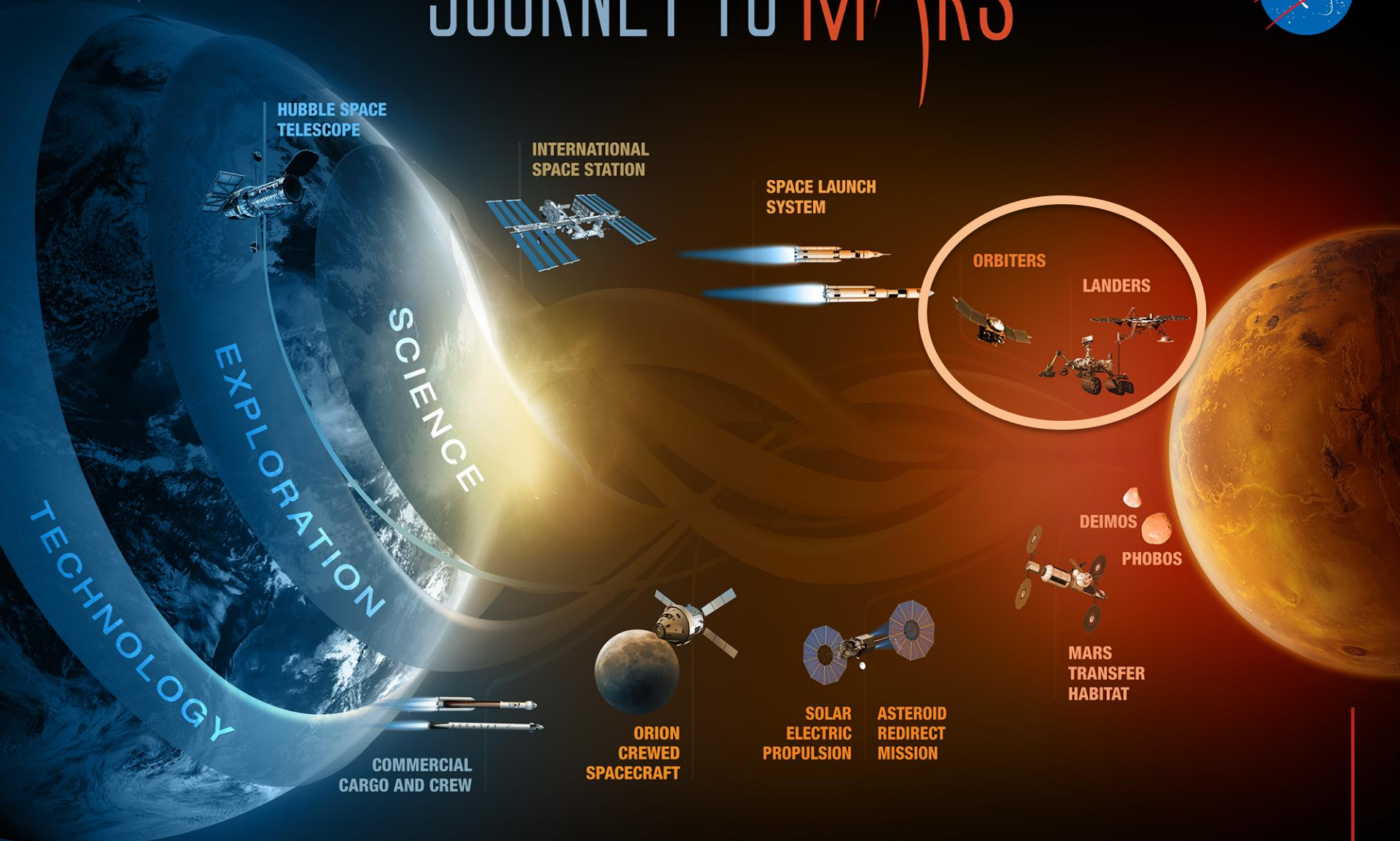
- ❑ Mars Exploration Program (MEP) continues to produce remarkable science and generate public interest in exploring Mars
- ❑ Planning for the future is a pressing priority, as the 2022 launch opportunity is only 5 years from the current budget planning horizon
- ❑ Collaboration on Science/Exploration synergies is an opportunity to excel in the 2020s



“Panorama compliments of the Curiosity rover”

9 Sept 2015

JOURNEY TO MARS



MISSIONS: 6-12 MONTHS
RETURN: HOURS

EARTH RELIANT

MISSIONS: 1-12 MONTHS
RETURN: DAYS

PROVING GROUND

MISSIONS: 2-3 YEARS
RETURN: MONTHS

EARTH INDEPENDENT

What We've Learned and Still Need to Learn at Mars

Orbital environment and operations



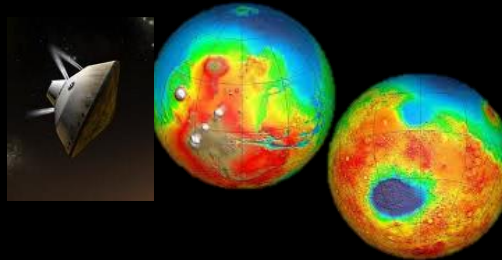
Learned:

- Deep space navigation
- Orbit transfer near low-gravity bodies
- Gravity assist
- Aero-braking
- Gravitational potential
- Mars' moons characteristics
- ISRU potential

To Learn:

- Return flight from Mars to Earth
- Autonomous Rendezvous & Docking
- ISRU feasibility
- Resource characterization of Mars moons
- High-power SEP

Capture, EDL & Ascent at Mars



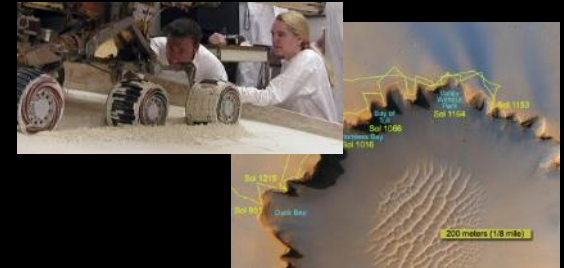
Learned:

- Spatial/temporal temperature variability
- Density and composition variability
- Storm structure, duration and intensity
- 1 mT Payload
- ~10 km Accuracy

To Learn:

- Ascent from Mars
- Large mass EDL
- Precision EDL
- Aero-capture
- Site topography and roughness
- Long-term atmospheric variability

Surface Operations at Mars



Learned:

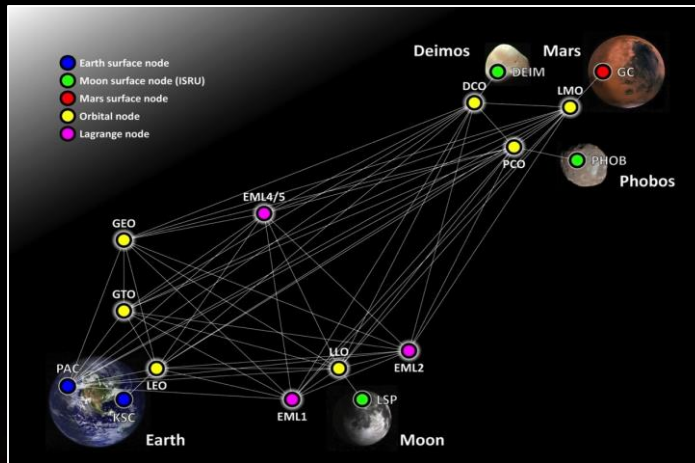
- Global topography: elevation and boulder distributions
- Remnant magnetic field
- Dust impacts on Solar Power / Mechanisms
- Radiation dose
- Global resource distribution
- Relay strategies, operations cadence

To Learn:

- Landing site resource survey
- Dust effects on human health, suits & seals
- Rad/ECLSS in Mars in environment
- Power sufficient for ISRU
- Surface Navigation

Strong Science and Exploration synergies motivate future Precursor collaboration

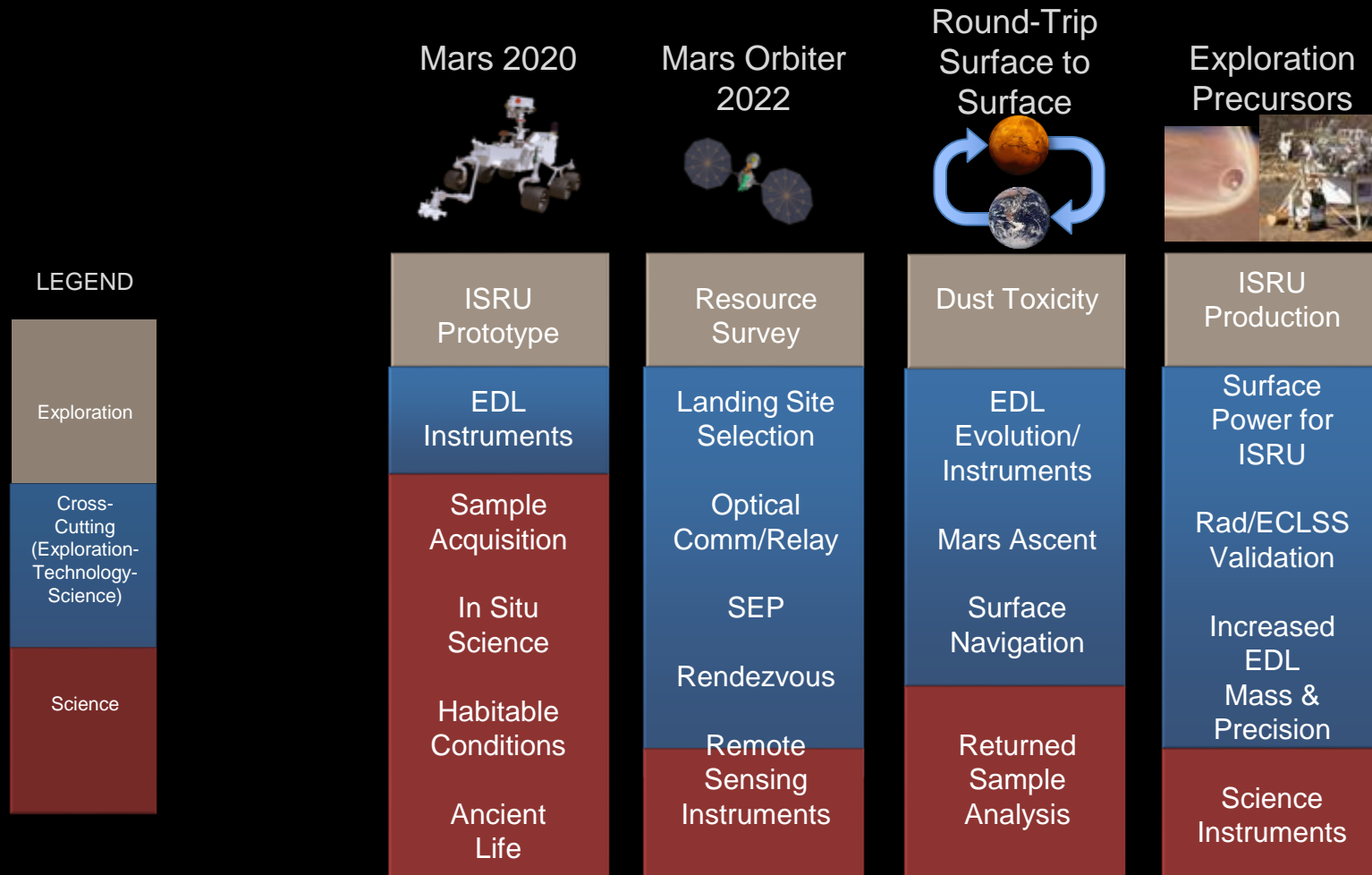
Resilient Architectures for Mars Exploration



Graphic used courtesy of de Weck et al

- There are many different architectures and implementation approaches that can be employed on the Journey to Mars
 - The first step of each architecture is the same – develop/validate common required capabilities
- NASA is studying precursor mission concepts for the 2020s to reduce the risk for these architectures and acquire relevant operational experience for the Journey to Mars

Conceptual Integrated Campaign for Mars Precursors “in the 2020’s”



Robotic precursors pursuing round-trip objectives intrinsically inform strategic exploration planning by providing invaluable flight experience

Science Exploration Integration

How can these objectives be pursued?

HSO-SAG

Human Science Objectives

Co-Chairs: D. Beaty, P. Niles
Ex Officio: Bussey, Davis, Meyer

HLS²

Human
Landing Site Study
Coordinators:
Davis, Bussey, Meyer

ICE Working Group

ISRU & Civil Engineering

Co-Chairs: S. Hoffman, R. Mueller
Ex Officio: Bussey, Davis

What are the Base & Exploration Zone
criteria? What & where are the
resources needed?

Where & what should humans explore

NRC Planetary Decadal Survey

MEPAG
Goals

Science
Objectives

NEX-SAG

Next Orbiter Options

OR

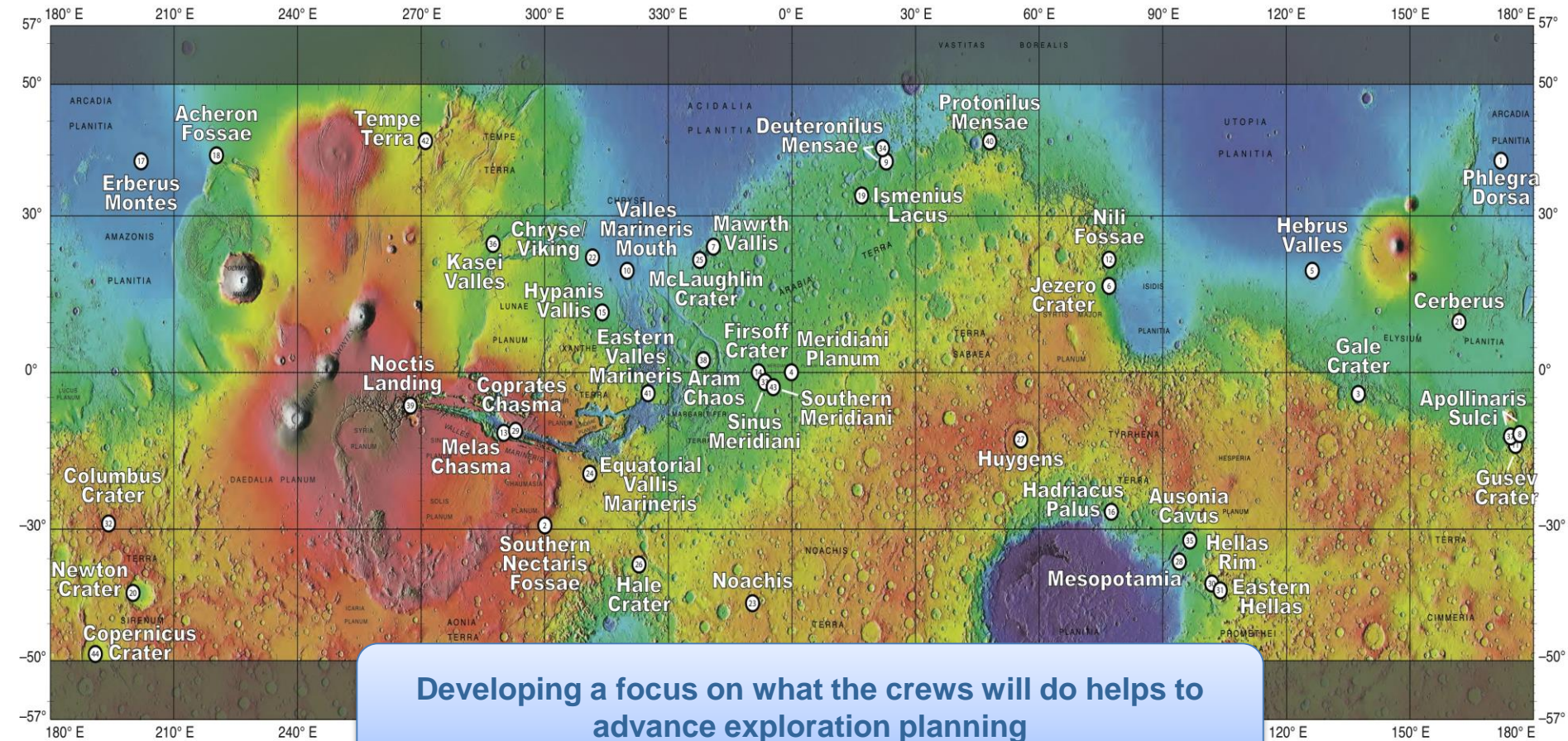
OR

Co-Chairs: R. Zurek
Ex Officio: Meyer, Bussey

Next Orbiter (NEX-SAG) Findings

- SEP brings the advantages of orbit flexibility and increased payload mass & power
- Advanced telecom provides necessary coverage for high-resolution data
- Considerable overlap between science goals and human exploration resource prospecting interests & derived objectives yield similar, mature instrument approaches
 - **Visible imaging** of HiRISE-class or better ($\sim 15\text{-}30\text{ cm/pixel}$)
 - **Polarimetric synthetic aperture radar** imaging with penetration depth of a few (<10) meters and spatial resolution of $\sim 15\text{ m/pixel}$ to search for shallow ground ice and crustal structure
 - **Short-wave IR spectral** mapping with a spatial resolution of $\sim 6\text{ m/pixel}$ ($3 \times \text{CRISM}$) with sufficient spectral resolution to detect key minerals
 - **Long-wave atmospheric sounding** for wind, temperature, & water vapor profiles with 5 km vertical resolution
 - **Thermal IR sounding** for aerosol (dust & ice) profiles
 - **Multi-band thermal IR** mapping of thermo-physical surface properties (e.g., ice overburden) and surface composition
 - **Wide-angle imaging** to monitor weather and surface frosts (global, km-scale)

Potential Exploration Zones for Human Missions to the Surface of Mars



Exploration Zones proposed for humans to Mars.
At the equator, circles are ~100km radius

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Lunar footprints



Mars wheelprints

Time for the next step!