

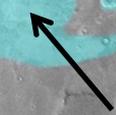
THE HYPANIS FLUVIAL-DELTAIC-LACUSTRINE SYSTEM IN XANTHE TERRA

Workshop Abstract #1051

Sabrina Delta



Hypanis Delta



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P. Grindrod², J. Davis⁷, M. Balme⁵, J.F. Bell III³, C. Stetson⁴, J. Richard⁸

CTX

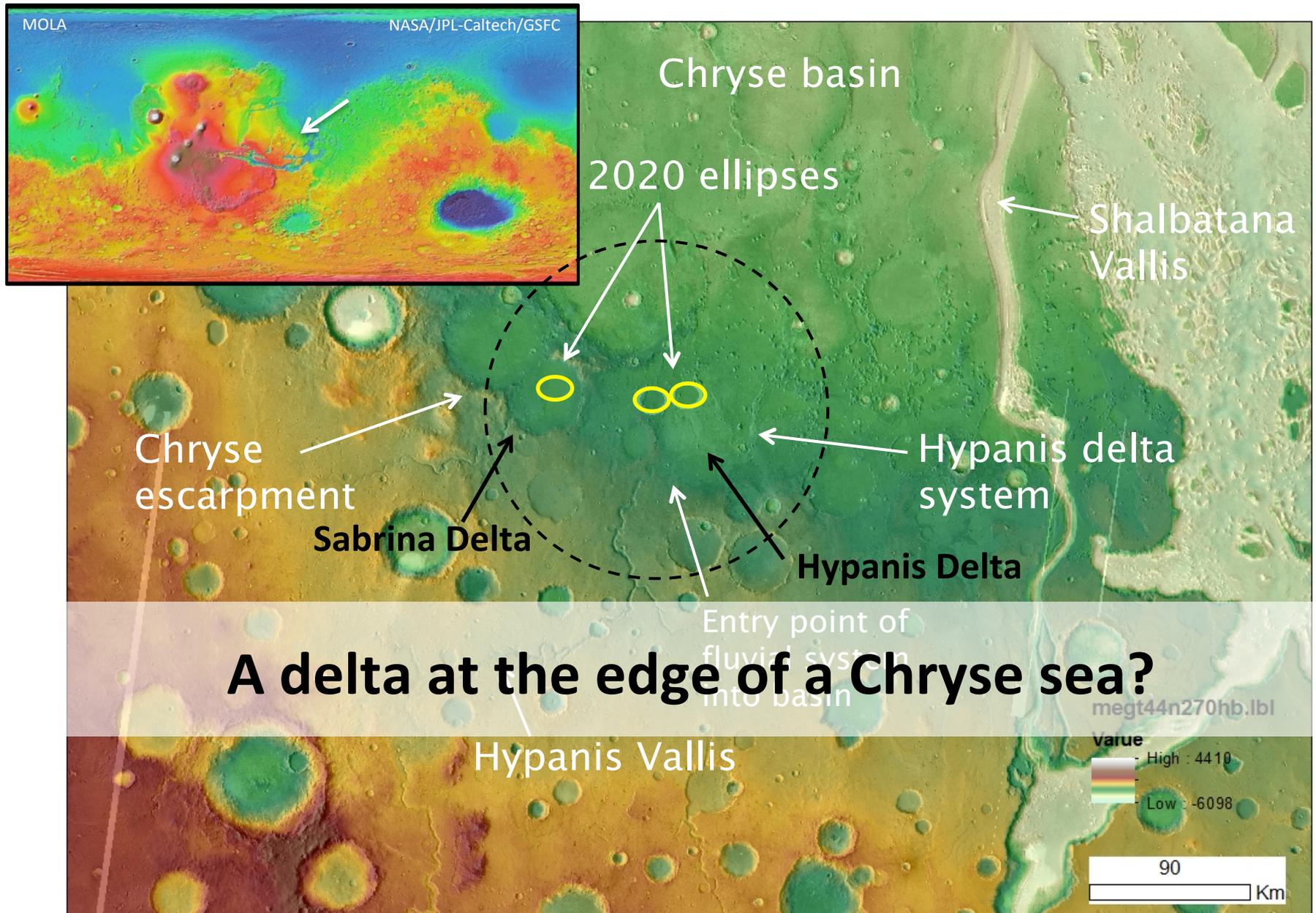
NASA/JPL-Caltech/MSSS

50 km



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MOLA Topography: Overview location of Hypanis system in Xanthe Terra



Exploration Zone Map

Lederberg

SROI	Latitude (°N)*	Longitude (°E)*	Altitude (km)	Dist. (km)**	Description
1	11.523	314.772	-2.695	5	Delta main
2	11.940	314.978	-2.700	30	Delta lobe
3	11.977	315.444	-2.729	53	Delta finger 1
4	11.699	315.708	-2.664	51	Delta finger 2
5	11.741	313.173	-2.553	62	Sabrina
6	11.160	314.342	-2.428	27	Channel 1
7	10.852	314.108	-2.318	57	Channel 2
8	10.634	313.994	-2.236	68	Channel 3
9	12.740	314.461	-2.945	51	Volcanic cones
10	12.231	314.232	-2.634	21	L. Rim 1
11	12.529	314.653	-2.826	38	L. Rim 2
12	13.296	314.864	-3.030	92	Bedrock 1
13	10.624	314.841	-2.331	70	Bedrock 2
14	12.424	315.687	-2.865	76	Ejecta contact

*Approx. coord. of center; ** Approx. line of sight distance from Field Station.

Hypanis
Vallis

50 km

SROI 12

SROI 9

SROI 11

SROI 10

RROI 1

RROI 2

SROI 5

RROI 3

SROI 7

SROI 8

SROI 14

SROI 3

SROI 2

SROI 4

SROI 1

SROI 6

SROI 13

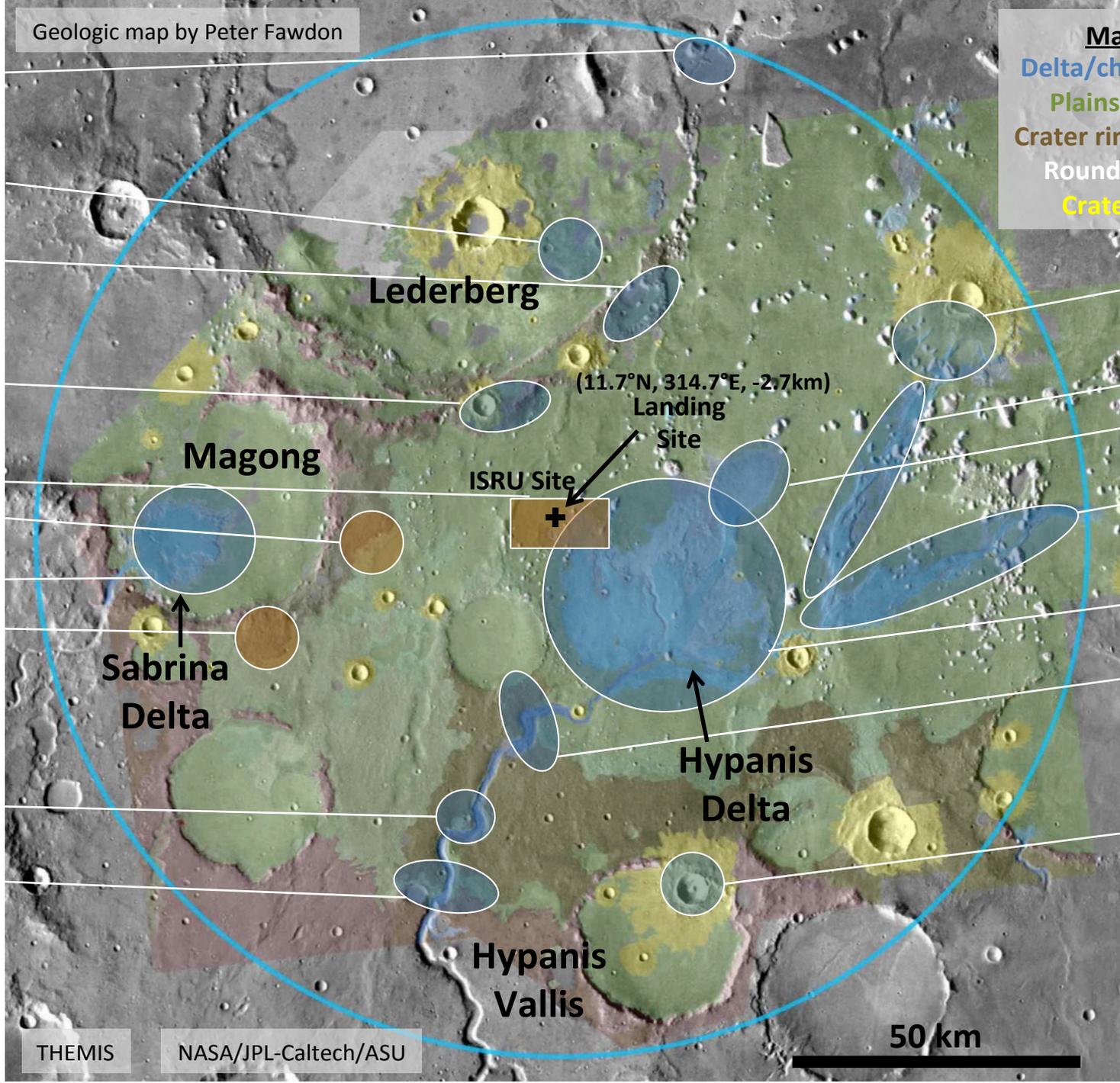
Geologic map by Peter Fawdon

Map Key

- Delta/channel units
- Plains material
- Crater rim/highlands
- Rounded buttes
- Crater ejecta

- SROI 12
- SROI 9
- SROI 11
- SROI 10
- RROI 1
- RROI 2
- SROI 5
- RROI 3
- SROI 7
- SROI 8

- SROI 14
- SROI 3
- SROI 2
- SROI 4
- SROI 1
- SROI 6
- SROI 13



THEMIS

NASA/JPL-Caltech/ASU

50 km

Science ROIs Rubric

1st EZ Workshop for Human Missions to Mars



Site Factors			SROI1	SROI2	SROI3	SROI4	SROI5	iROI6	iROI7	iROI8	iROI9	ROI10	ROI11	SROI12	SROI13	SROI14	EZ SUM		
Science Site Criteria	Astrobio	Threshold	Potential for past habitability	●	●	●	●	●								~	8,1		
		AND/OR	Potential for present habitability/refugia															0,0	
	Qualifying	Potential for organic matter, w/ surface exposure	●															9,1	
	Atmospheric Science	Threshold	Noachian/Hesperian rocks w/ trapped atmospheric gases															2,1	
		Qualifying	Meteorological diversity in space and time	●														6,2	
			High likelihood of surface-atmosphere exchange	●														14,0	
			Amazonian subsurface or high-latitude ice or sediment															0,0	
			High likelihood of active trace gas sources	?													0,0		
	Geoscience	Threshold	Range of martian geologic time; datable surfaces															14,0	
			Evidence of aqueous processes	●														10,3	
			Potential for interpreting relative ages	●														13,1	
		Qualifying	Igneous rocks that date to different times	○															
			Near-surface ice, glacial or permafrost																
			Noachian or pre-Noachian bedrock units																
Outcrops with remnant magnetization																			
Primary, secondary, and basin-forming impact deposits			●																
	Structural features with regional or global context																		
	Diversity of aeolian sediments and/or landforms	●																	

Key Points

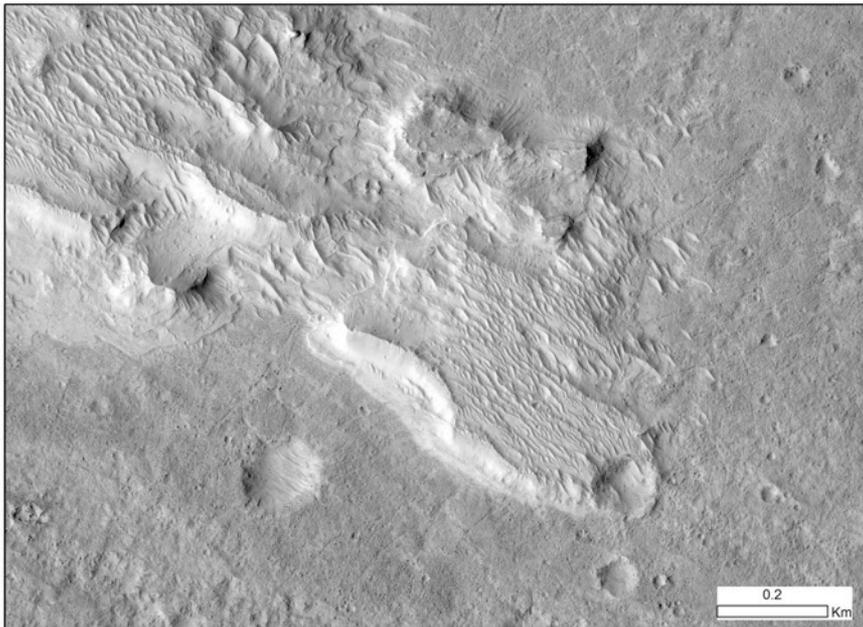
- **GREAT** environment [Summons et al., 2011] for organic formation, concentration, and preservation
- Both Noachian bedrock and Hesperian units
- Clear evidence of aqueous activity
- Addresses major global question: **Was there a Chryse Sea?!**

Science ROIs 1-4,6-7

1st EZ Workshop for Human Missions to Mars



Image of SROI 1



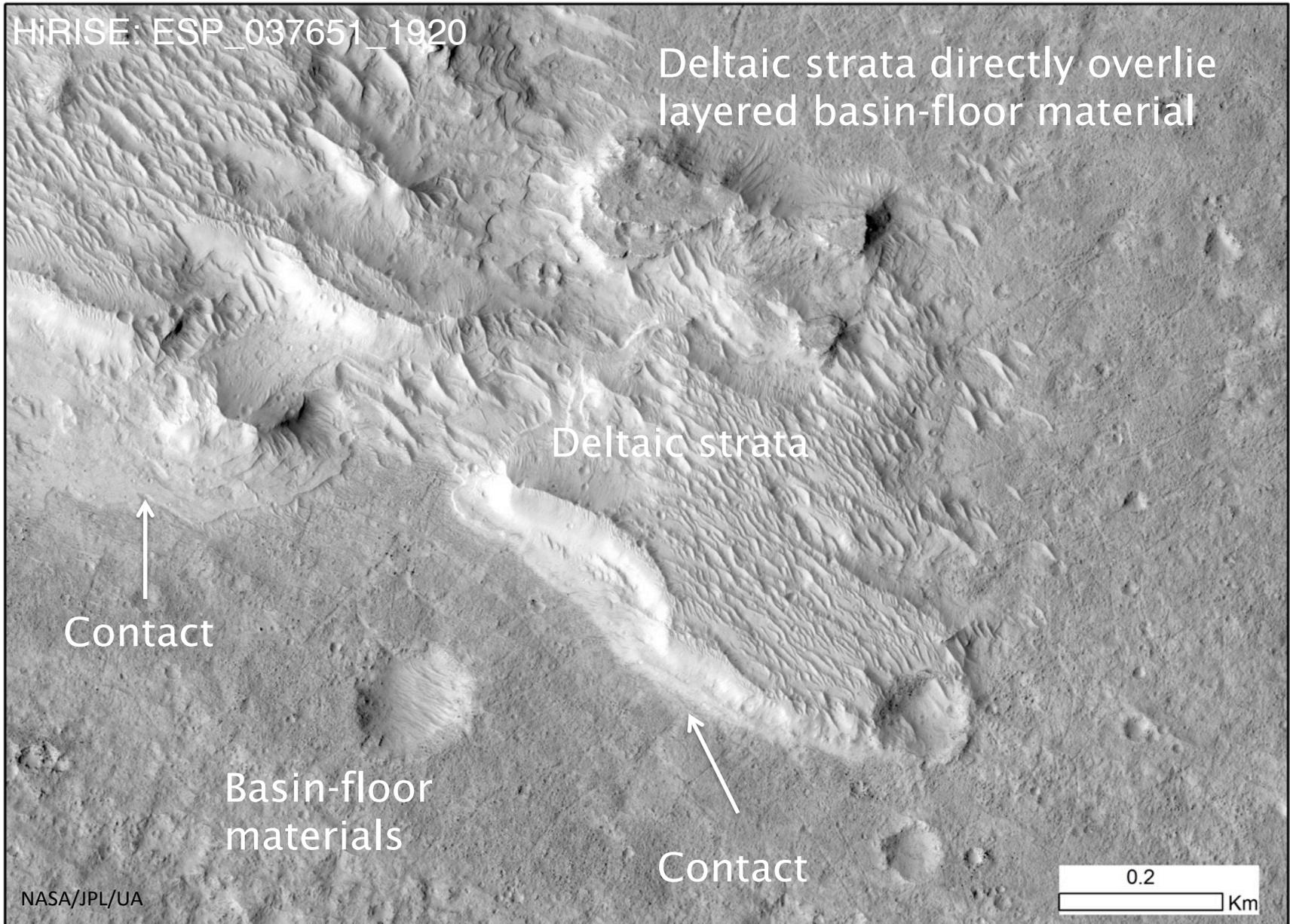
HiRISE: ESP_037651_1920

NASA/JPL/UA

Hypanis Delta and Inverted Fluvial Channels

- Delta main: 11.5°N, 314.8°E, -2.7 km
- Layered sedimentary strata in lobes of Hypanis delta and in inverted channel complexes
- Threshold:
 - Potential for past habitability
 - Clear evidence for aqueous processes
 - Identifiable stratigraphic contacts
 - Outcrop units with a range of ages
- Qualifying:
 - Access to deposits with potential for containing organic matter

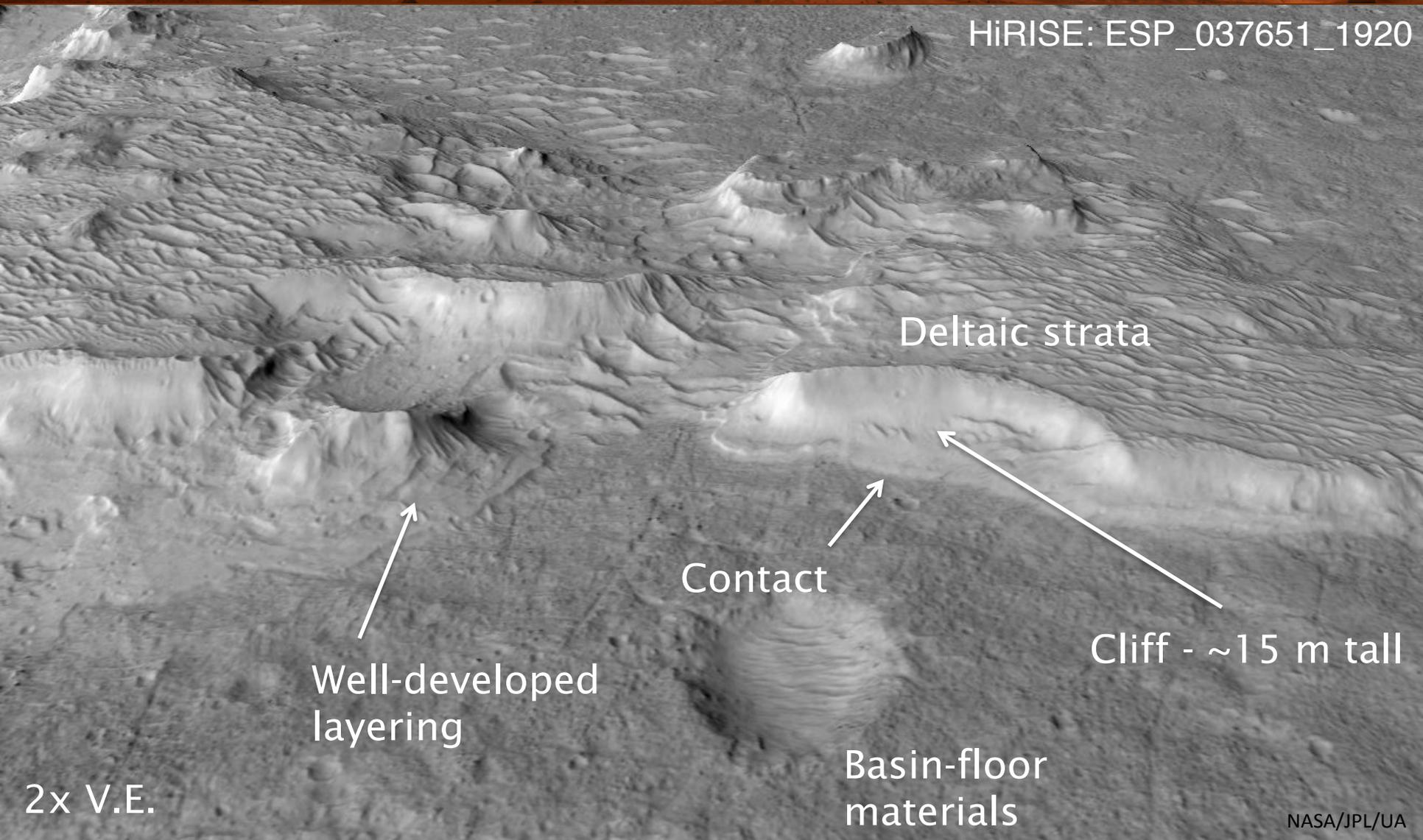
SROI 1 Hypanis Deltaic strata



Perspective view of SROI 1



HiRISE: ESP_037651_1920



Deltaic strata

Contact

Cliff - ~15 m tall

Well-developed layering

Basin-floor materials

2x V.E.

SROI 1 Hypanis Deltaic strata .



HiRISE: ESP_036277_1920

Deltaic strata directly overlie
layered basin-floor material

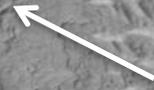
Contact



Deltaic strata

Basin-floor -
lacustrine
strata?

Contact



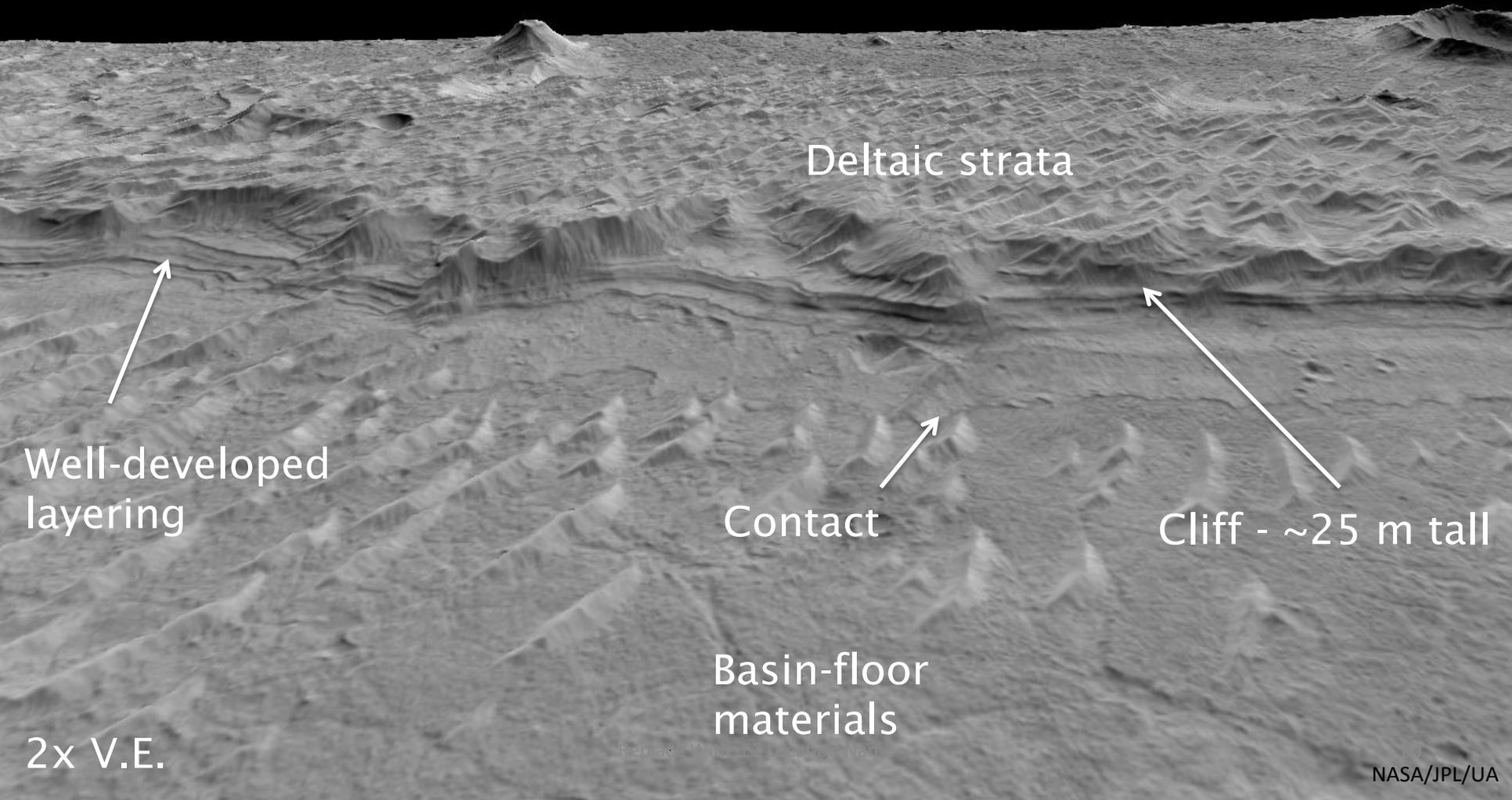
Well-developed
sedimentary
layering



Perspective view of SROI 1



HiRISE: ESP_036277_1920



Deltaic strata

Well-developed layering

Contact

Cliff - ~25 m tall

Basin-floor materials

2x V.E.

Science ROIs 1-4,6-7

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Image of SROI 2

Hypanis Delta

- Prominent lobe: 11.9°N, 314.9°E, -2.7 km
- Layered sedimentary strata in lobes of Hypanis delta and in inverted channel complexes
- Threshold:
 - Potential for past habitability
 - Clear evidence for aqueous processes
 - Identifiable stratigraphic contacts
 - Outcrop units with a range of ages
- Qualifying:
 - Access to deposits with potential for containing organic matter

Delta lobe

Prominent feeder channel

2.5 km



HiRISE: ESP_03494_1920

NASA/JPL/UA

Perspective view of SROI 2



HiRISE: ESP_03494_1920

Deltaic strata

Contact

Cliff - ~40 m tall

Well-developed layering

Abundant TARs

Basin-floor materials

2x V.E.

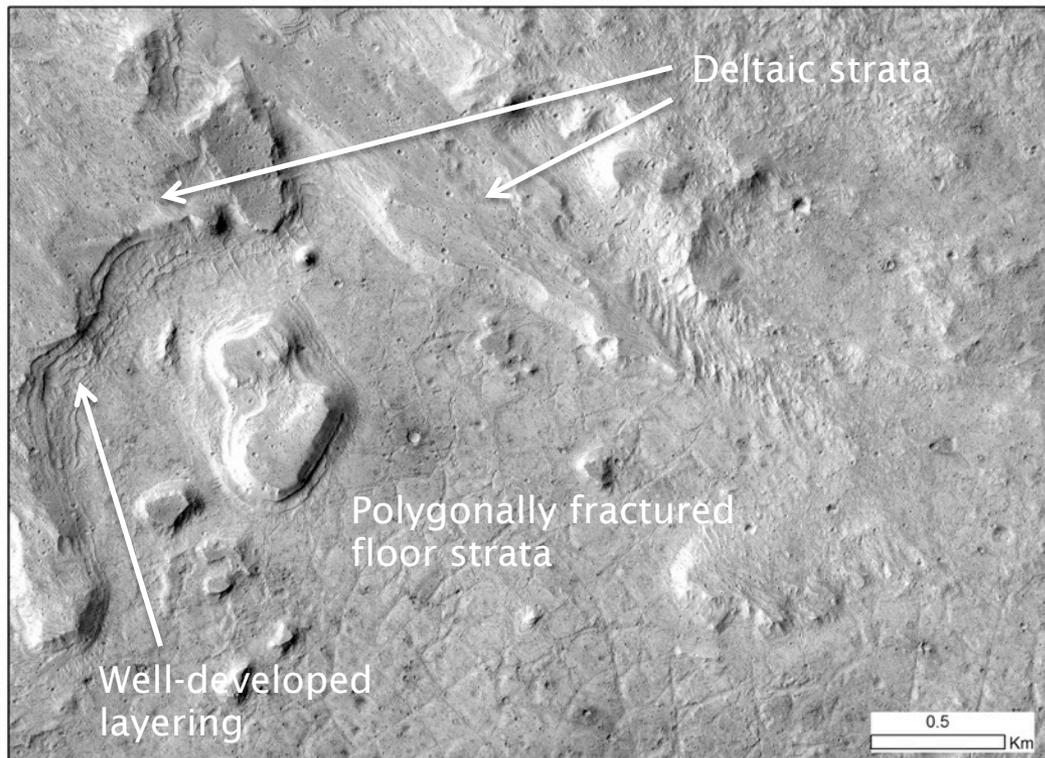
Reproduced with permission from the Mars Science Laboratory team

Science ROI 5

1st EZ Workshop for Human Missions to Mars

Image of SROI 5

Sabrina Delta



- 11.7°N, 313.1°E, -2.5 km
- Layered sedimentary strata in lobes of Hypanis delta and in inverted channel complexes
- Threshold:
 - Potential for past habitability
 - Clear evidence for aqueous processes
 - Identifiable stratigraphic contacts
 - Outcrop units with a range of ages
- Qualifying:
 - Access to deposits with potential for containing organic matter

HiRISE: ESP_026601_1920

NASA/JPL/UA

Hypanis Delta Exploration Zone

13

Perspective view of SROI 5



HiRISE: ESP_026601_1920

Deltaic strata

Contact

Cliff - ~30 m tall

Well-developed layering

Polygonally fractured basin-floor materials

2x V.E.

Hypanis Valles: *Regional Context*

Extensive drainage systems fed Hypanis

Chryse
escarpment

Sabrina
Vallis

Hypanis delta system

Tyras
Valles

Hypanis
Valles

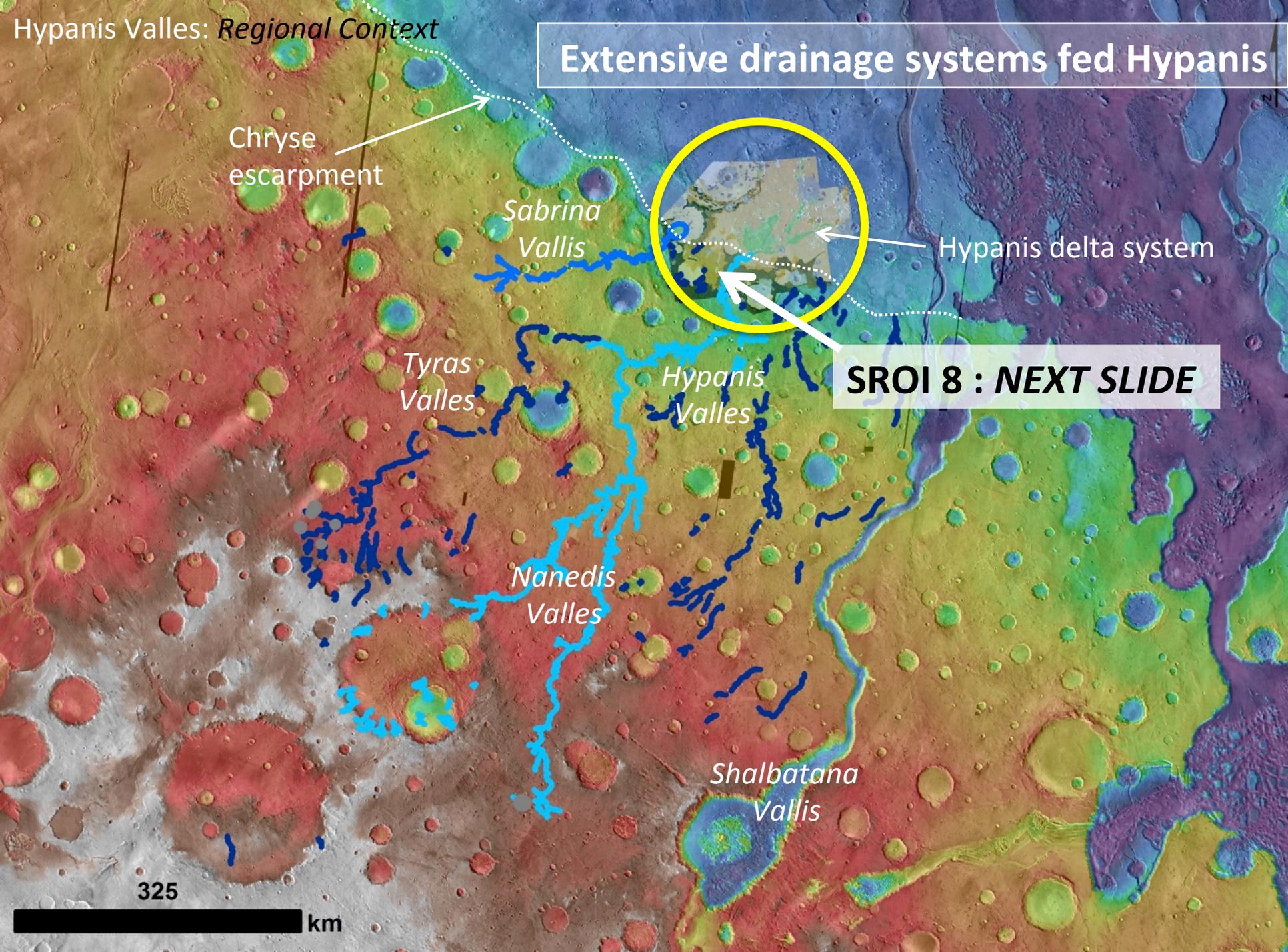
SROI 8 : NEXT SLIDE

Nanedis
Valles

Shalbatana
Vallis

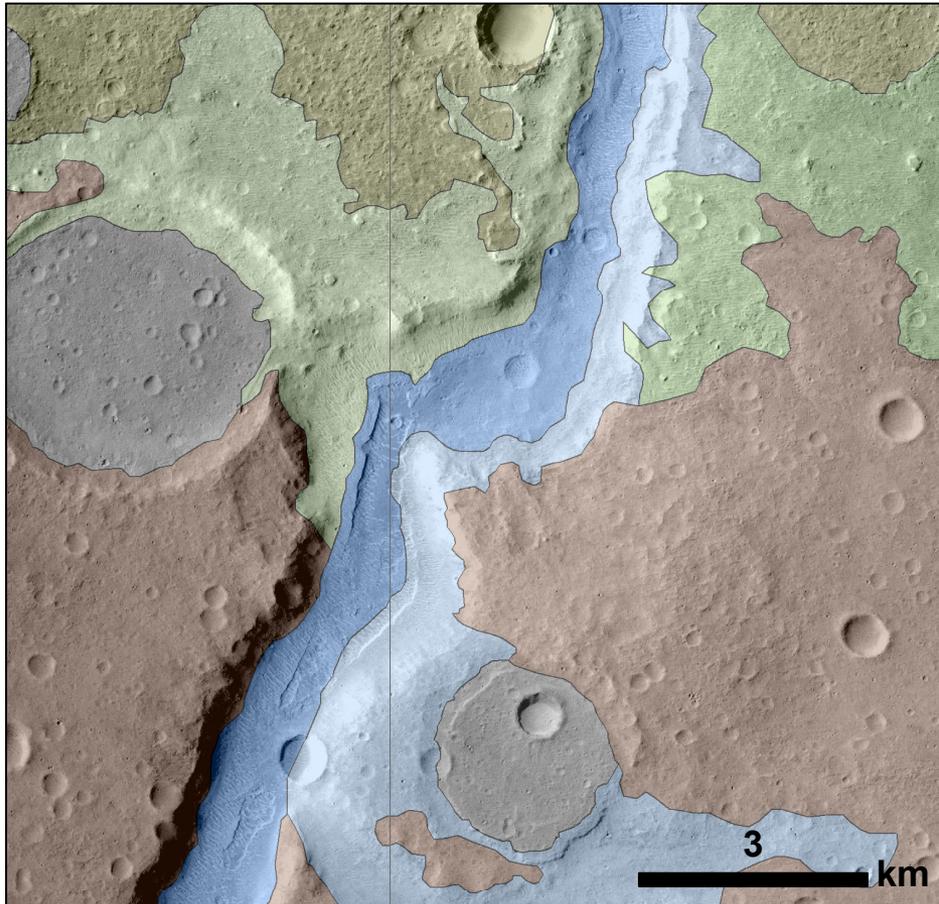
325

km



Science ROI 8

1st EZ Workshop for Human Missions to Mars



Channel Contact on Dichotomy Boundary

- 10.6°N, 313.9°E, -2.2 km
- Hypanis Valles channel exposes stratigraphy along the Martian Dichotomy Boundary
- Threshold:
 - Identifiable stratigraphic contacts
 - Range of datable surfaces (Noachian and Hesperian plains units)
- Qualifying:
 - Evidence of aqueous activity
 - Exposed Noachian bedrock

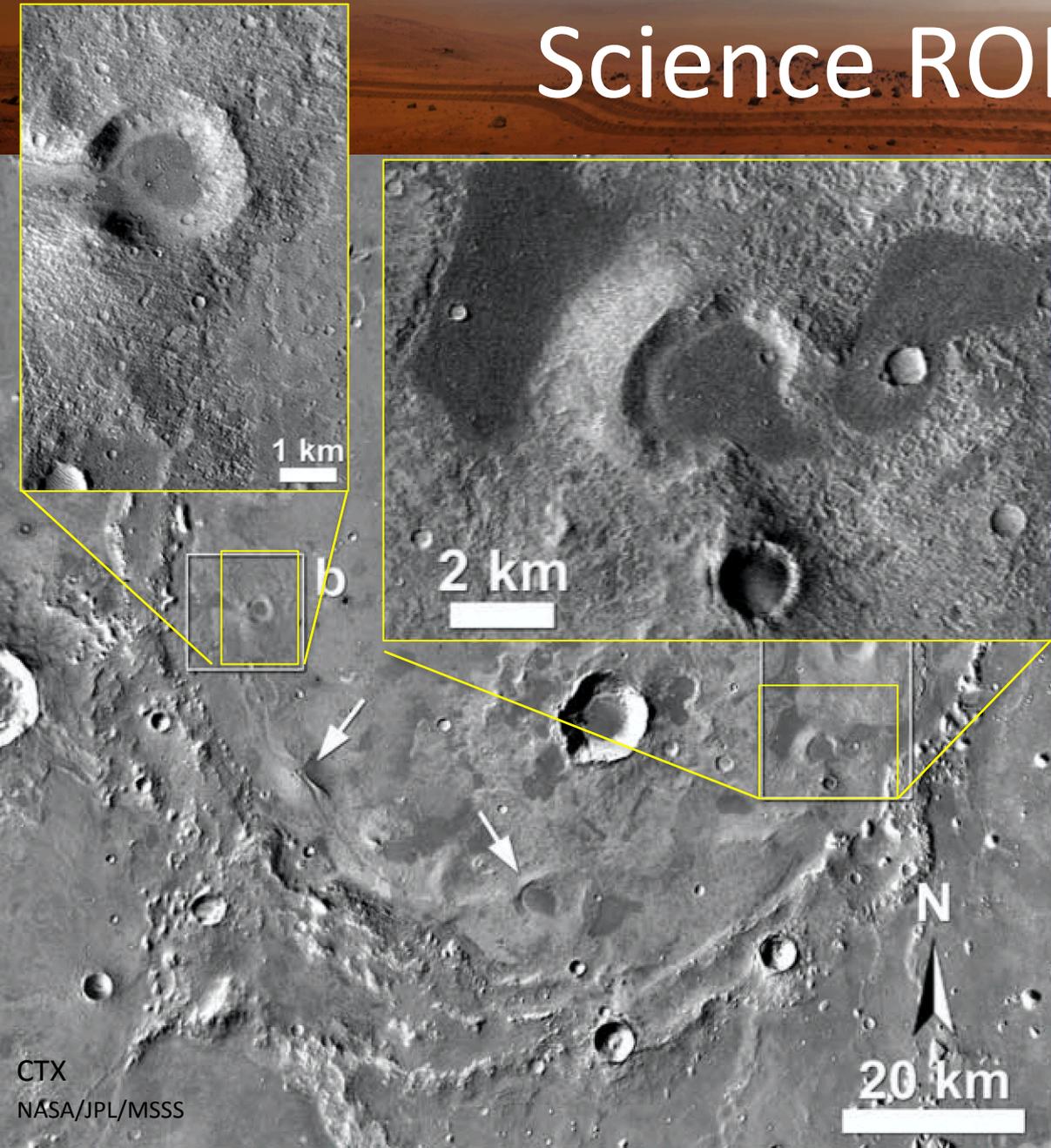
Science ROI 9

1st EZ Workshop for Human Missions to Mars



Possible Hydrovolcanic Features

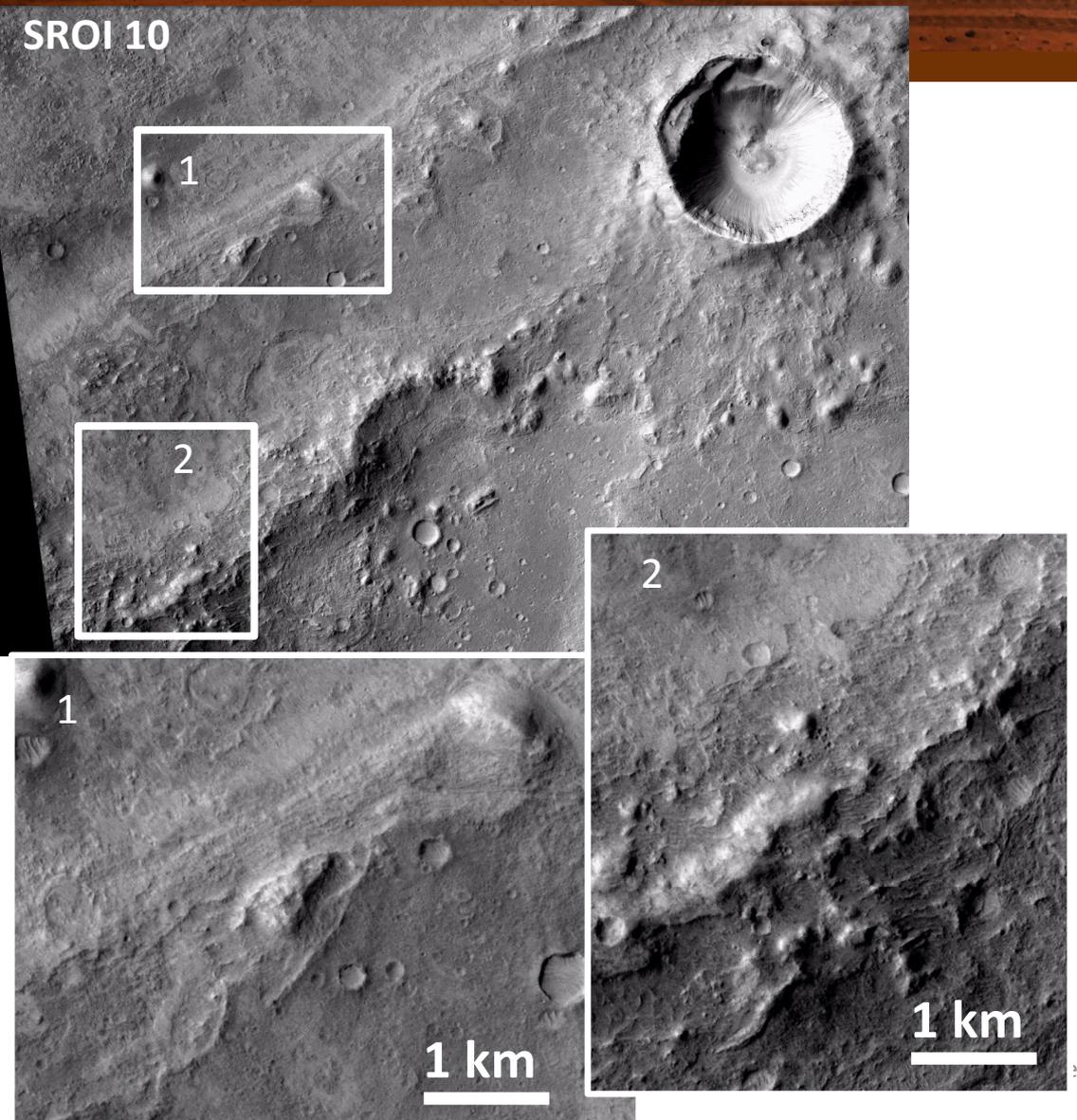
- 11.7°N, 314.4°E, -2.5 km
- Conical landforms in Ledeborg Crater hypothesized to be tuff cones by Brož & Hauber (2013)
- Threshold:
 - Evidence of aqueous activity
 - Stratigraphic context Ledeborg and other units
- Qualifying:
 - Potential Hesperian igneous materials



Science ROIs 10-12



SROI 10



Lederberg Rim Outcrops

- SROI 10: 12.2°N, 314.2°E, -2.6 km
- SROI 11: 12.6°N, 314.6°E, -2.8 km
- SROI 12: 13.3°N, 314.8°E, -3.0 km

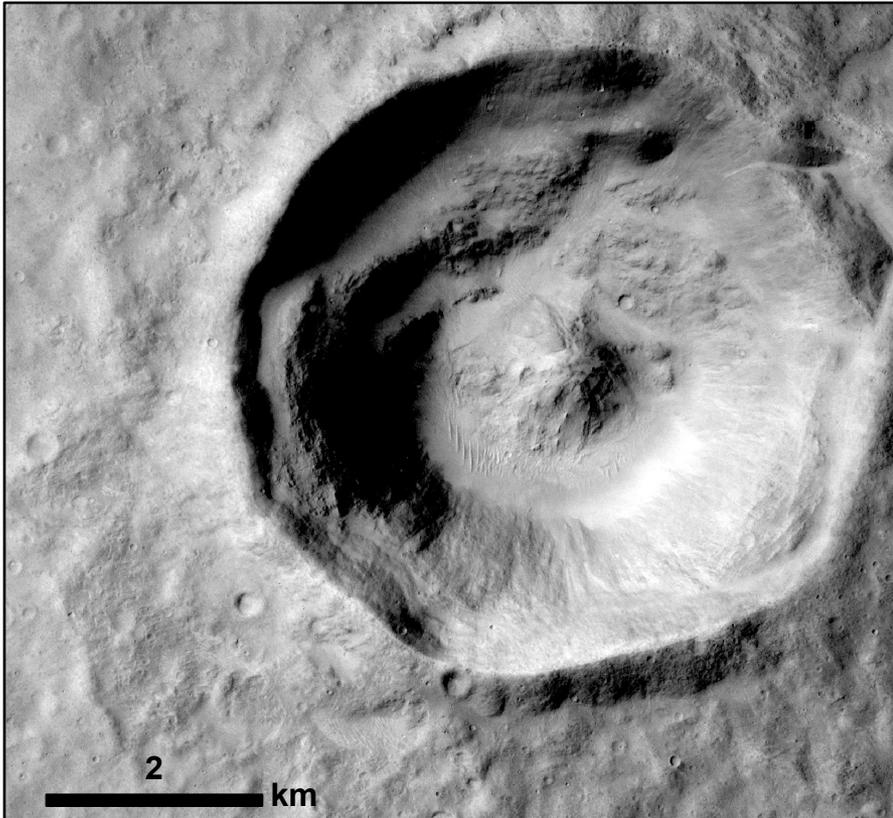
- Noachian bedrock exposed in crater rim
- Multiple units with some indication of layering
- Threshold:
 - Identifiable stratigraphic contacts
 - Potential for trapped gases in impact melts
- Qualifying:
 - Fresh impact structures
 - Exposed Noachian bedrock

Science ROIs 13-14

1st EZ Workshop for Human Missions to Mars



SROI 13: Fresh Crater



Craters and Ejecta

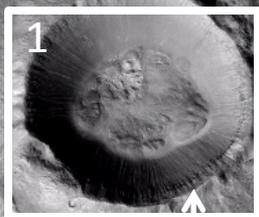
- SROI 13: 10.6°N, 314.8°E, -2.3km
- SROI 14: 12.4°N, 315.7°E, -2.8km
- Threshold:
 - Identifiable stratigraphic contacts between multiple EZ units
 - Potential for trapped gases in impact melts
- Qualifying:
 - Fresh impact structure
 - Exposed Noachian bedrock

SROI 14: Fresh Crater and Ejecta Contacts

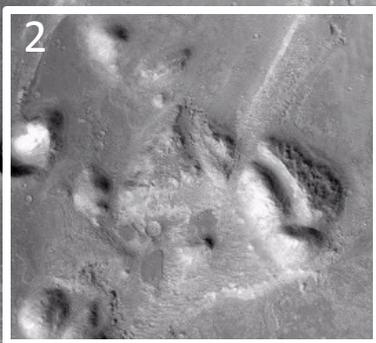
CTX: B09_013296_1920_XI_12N044W

NASA/JPL/MSSS

Delta unit →
continues



↑
Bedrock layer in
crater wall



↑
Erosion of
ejecta lobe

↙
To SROI 3 and
Hypanis delta

5 km



Igneous unit?

Delta unit

Fluvial plains
unit



Collection of floor
morphologies

2 km

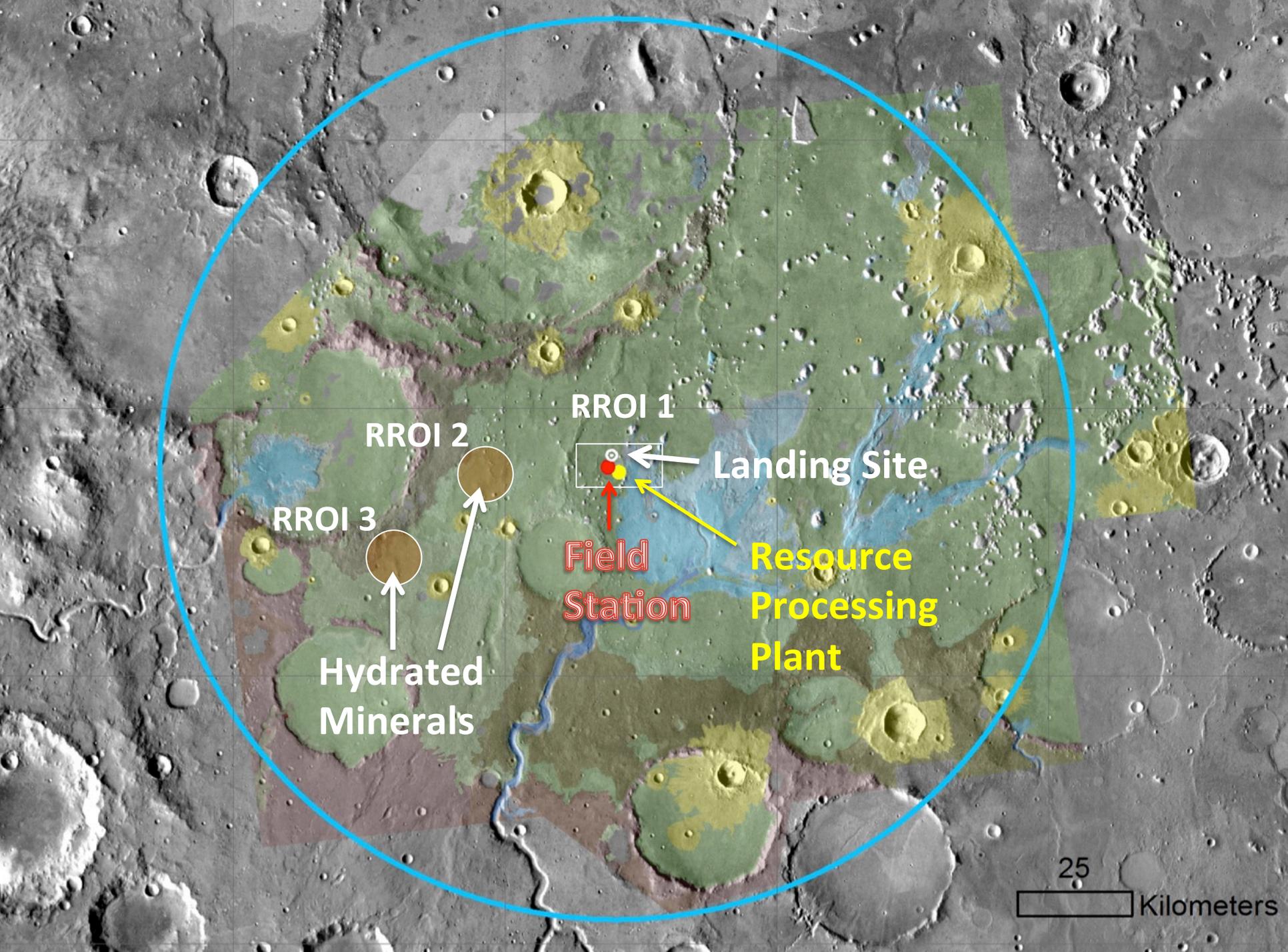
Resource ROIs Rubric

1st EZ Workshop for Human Missions to Mars



Site Factors			RROI1	RROI2	RROI3	EZ SUM	
ISRU and Civil Engineering Criteria	Engineering	Meets First Order Criteria (Latitude, Elevation, Thermal Inertia)	●	●	●	3,3	
	Water Resource	Threshold	AND/OR Potential for ice or ice/regolith mix				0,0
			Potential for hydrated minerals	●	●	●	3,3
			Quantity for substantial production	●	●	●	3,3
			Potential to be minable by highly automated systems	●	●	●	3,3
			Located less than 3 km from processing equipment site	●	●	●	3,3
			Located no more than 3 meters below the surface	●	●	●	3,3
			Accessible by automated systems	●	●	●	3,3
			Qualifying	Potential for multiple sources of ice, ice/regolith mix and hydrated minerals			
		Distance to resource location can be >5 km	●	●	●	3,3	
		Route to resource location must be (plausibly) traversable	●	●	●	3,3	
	Civil Engineering	Threshold	~50 sq km region of flat and stable terrain with sparse rock distribution	●	●	●	3,3
			1-10 km length scale: <10°	●	●	●	3,3
			Located within 5 km of landing site location	●			1,0
		Qualifying	Located in the northern hemisphere	●	●	●	3,3
		Evidence of abundant cobble sized or smaller rocks and bulk, loose regolith	●			1,0	
		Utilitarian terrain features	●			1,0	
	Food Production	Qualifying	Low latitude	●	●	●	3,3
			No local terrain feature(s) that could shadow light collection facilities	●	●	●	3,3
			Access to water	●	●	●	3,3
		Access to dark, minimally altered basaltic sands	●			1,0	
	Metal/Silicon Resource	Threshold	Potential for metal/silicon	●	●	●	3,3
			Potential to be minable by highly automated systems	●	●	●	3,3
			Located less than 3 km from processing equipment site	●	●	●	3,3
			Located no more than 3 meters below the surface	●	●	●	3,3
			Accessible by automated systems	●	●	●	3,3
		Qualifying	Potential for multiple sources of metals/silicon	●			1,0
			Distance to resource location can be >5 km				0,0
		Route to resource location must be (plausibly) traversable	●	●	●	3,3	

Key	
●	Yes
○	Partial Support or Debated
	No
?	Indeterminate



RROI 1

Landing Site

RROI 2

Resource
Processing
Plant

RROI 3

Field
Station

Hydrated
Minerals

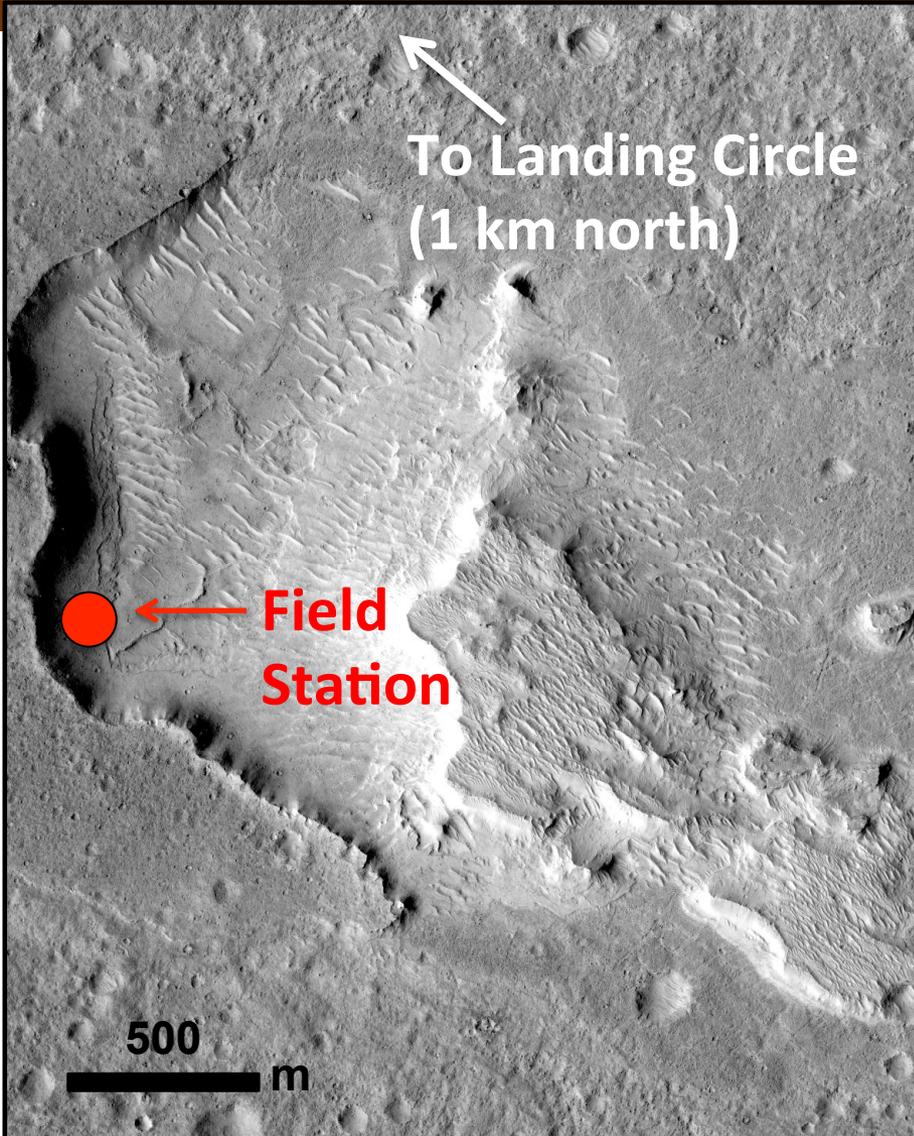
25

Kilometers

Resource ROI 1



s to Mars



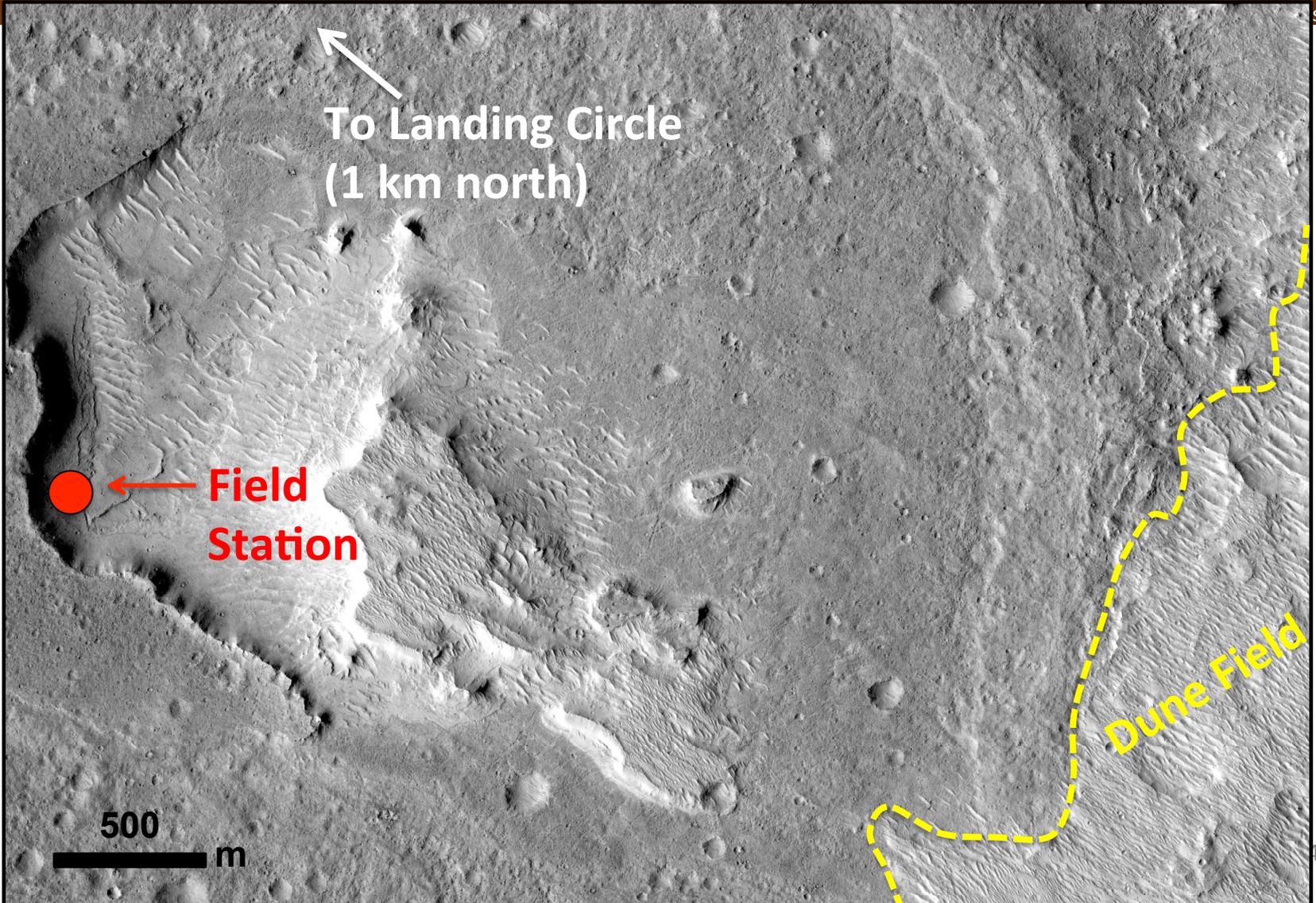
Field Station

- Depression with steep cliffs forms a natural shelter with some protection from wind and radiation
- Easily traversable from landing site (3 km north)
- Easy traverse on smooth, compact terrain to processing plant and resources (2.5 km SE)

Resource ROI 1



s to Mars



To Landing Circle
(1 km north)

Field
Station

500
m

Dune Field

Resource ROI 1

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SECTIONS



HOME



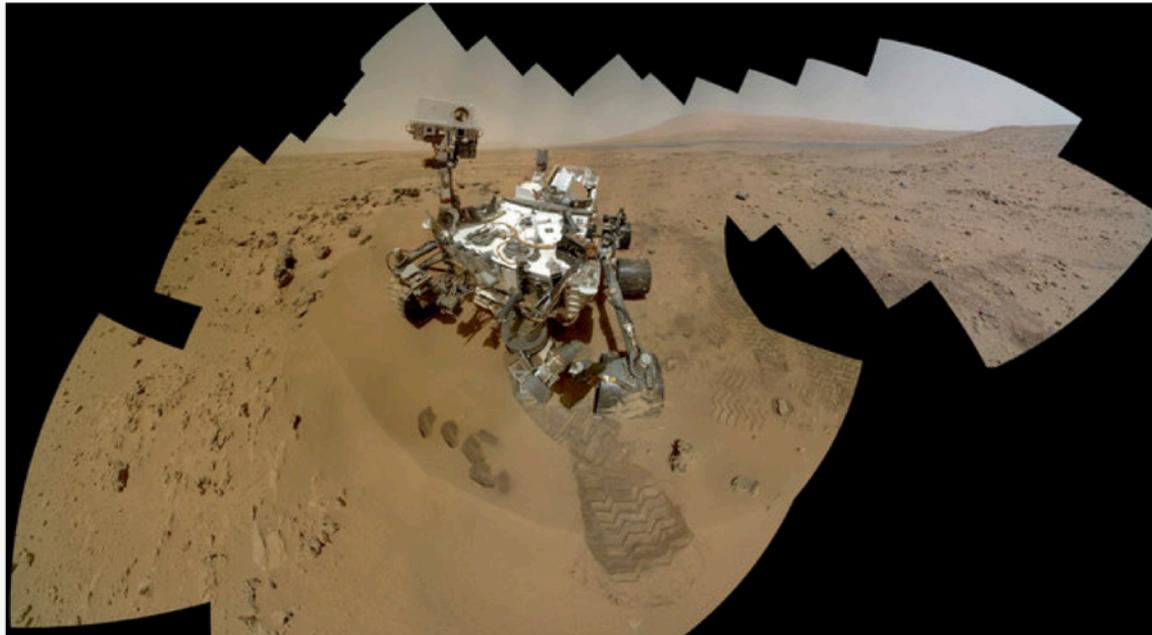
SEARCH

The New York Times

SPACE & COSMOS

Hitting Pay Dirt on Mars

By KENNETH CHANG SEPT. 30, 2013



Hypanis Delta Exploration Zone

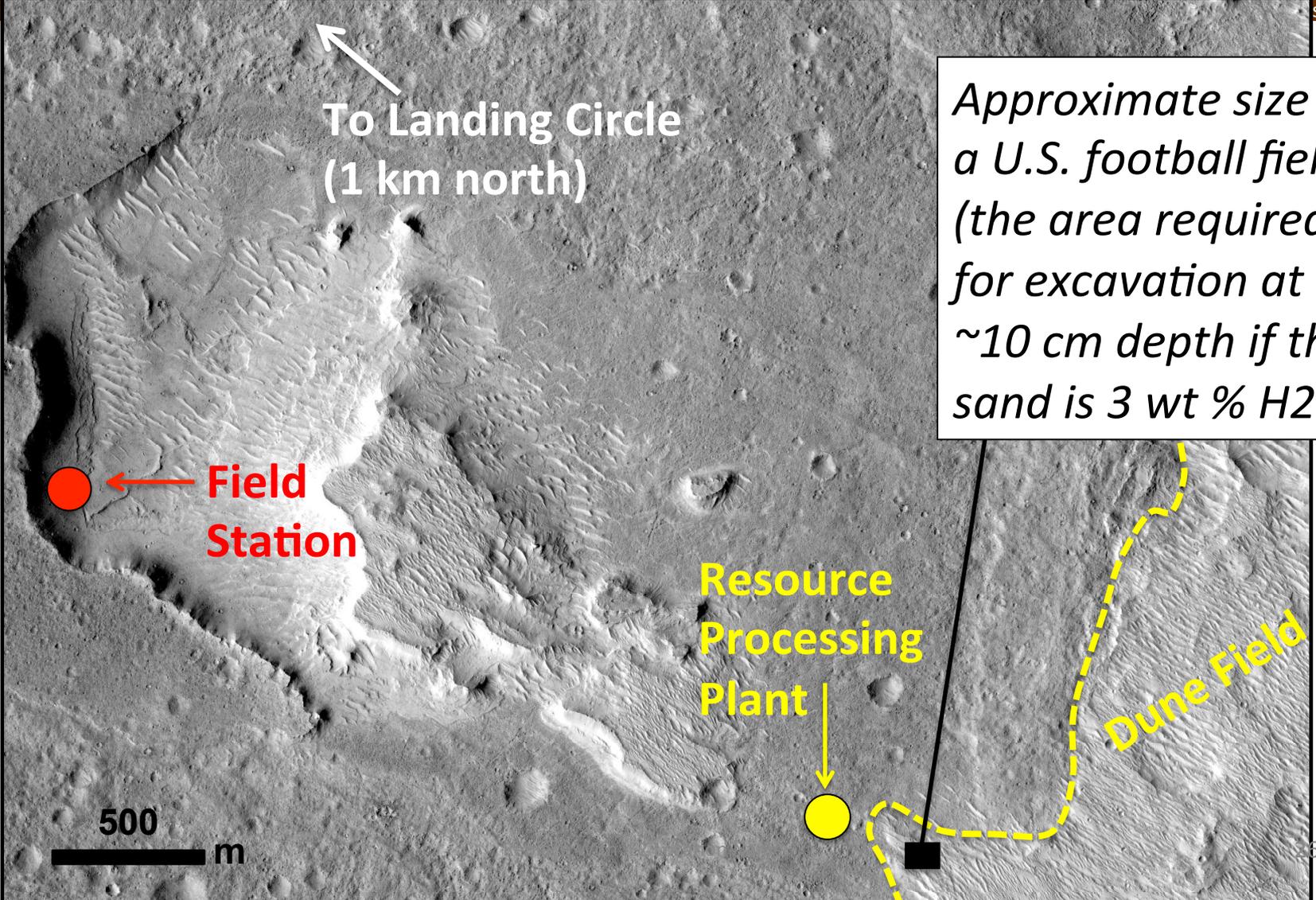
Water and Metal Resource: Dunes

- In-situ water resources are potentially provided by the extensive aeolian dune sands
- Observations by MSL of atmospheric water adsorbed in dune sands (Meslin et al., 2013) and 3-6 wt. % H₂O in sands at Rocknest (Leshin et al., 2013)
- Absence of crater retention indicates that the TARs are weakly indurated and could be easily mobilized to a processing facility constructed on adjacent, level bedrock
- Could provide metal and silicon resources and building material

Resource ROI 1



s to Mars



To Landing Circle
(1 km north)

Field Station

Resource Processing Plant

Dune Field

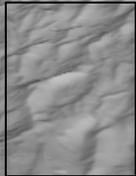
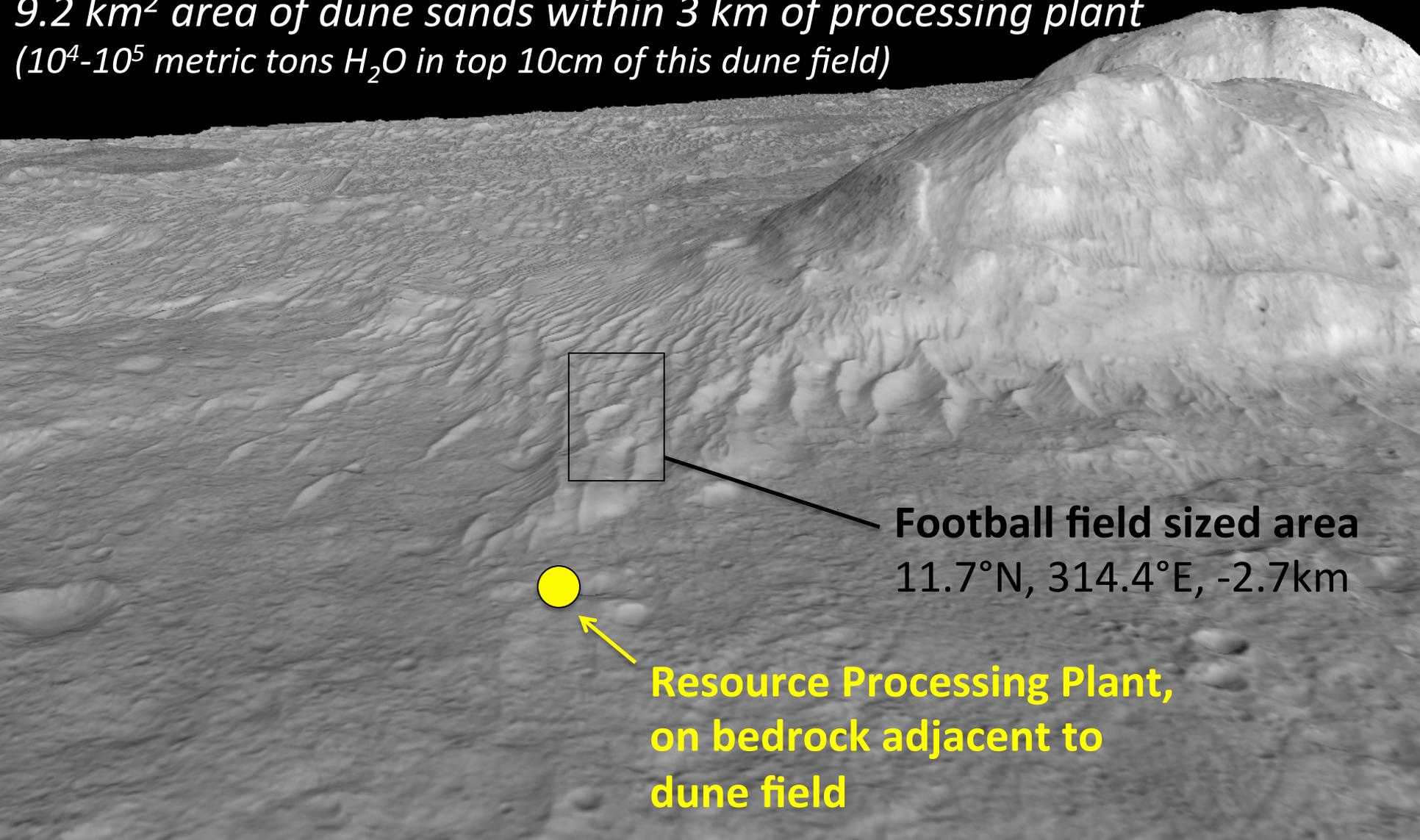
500 m

Approximate size of a U.S. football field (the area required for excavation at ~10 cm depth if the sand is 3 wt % H₂O)

Dune Sands as Resources



*9.2 km² area of dune sands within 3 km of processing plant
(10⁴-10⁵ metric tons H₂O in top 10cm of this dune field)*



**Football field sized area
11.7°N, 314.4°E, -2.7km**



**Resource Processing Plant,
on bedrock adjacent to
dune field**

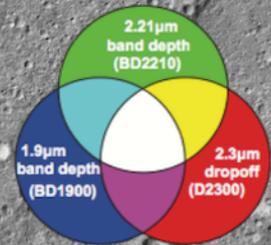
Resource ROIs 2-3



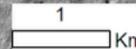
Water Resource: Hydrated minerals in multiple EZ locations

FRS0003134F - PHY

Fe/Mg phyllosilicates detected by CRISM



BD1900 Bound molecular H₂O
BD2210 Al-OH minerals
D2300 Hydroxylated Fe/Mg silicates



Clays and Hydrated Units

- RROI 2: 11.8°N, 313.8°E, -2.5 km
- RROI 3: 11.5°N, 313.4°E, -2.3 km
- Fe/Mg phyllosilicates identified in CRISM FRT and MSP observations at multiple locations on the plains units
- Threshold:
 - Hydrated minerals
 - Multiple large exposures
 - Accessible and potentially easily-minable
- Qualifying:
 - Routes to all exposures are easily traversable from landing site

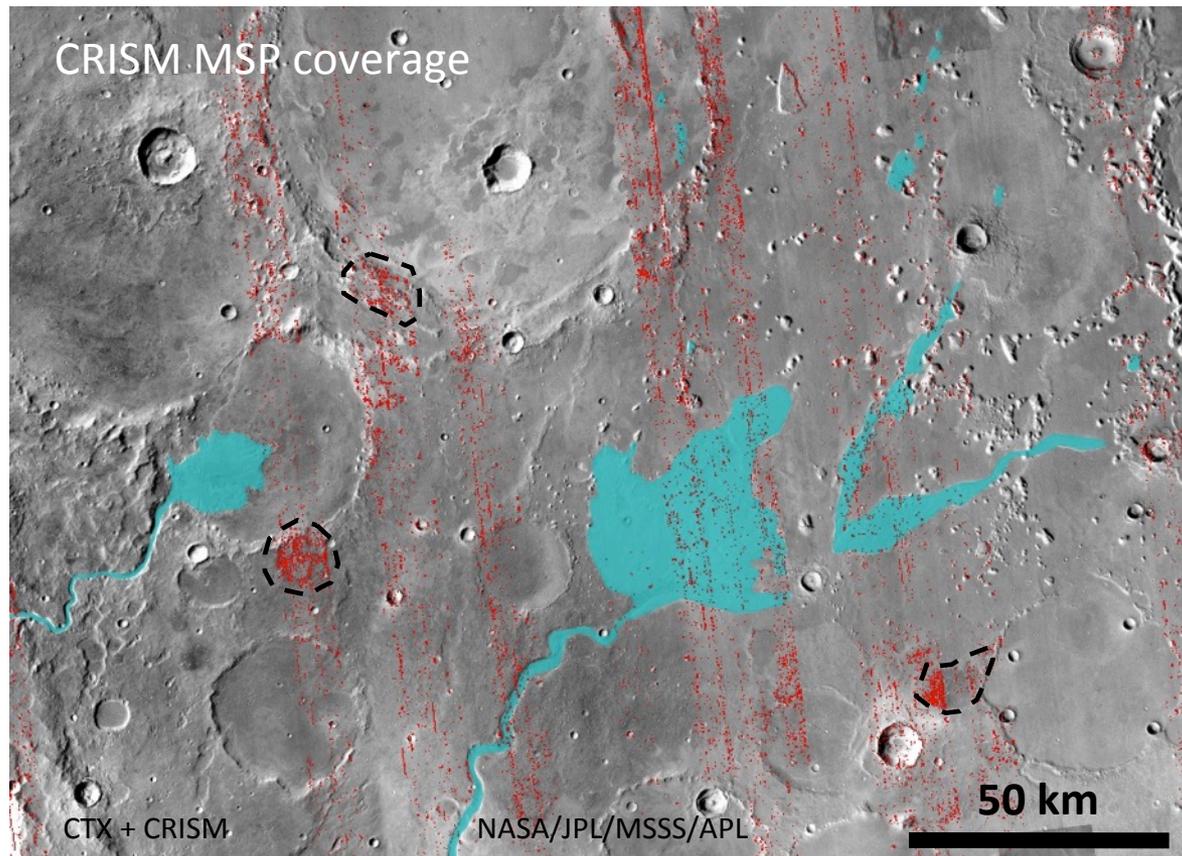
Highest Priority EZ Data Needs



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Most important to assess science: **CRISM FRTs**

Most important to assess resources: **CRISM FRTs**



Band depth at 2.3 μ m

0.5% < **D2300** < 3%

Summary

1st EZ Workshop for Human Missions to Mars

- Hypanis delta is a great environment for potential habitability and astrobiological preservation
- Key stratigraphic contacts and ROIs are ideal for a human mission since can actually access these distant sites
- Resources are relatively easy to mine – processing and accessing is right next to landing site
- Science is fascinating and has GLOBAL implications: a delta along the dichotomy boundary, **Chryse sea?**

BACKUP SLIDES

Prioritization List of EZ Data Needs

1st EZ Workshop for Human Missions to Mars



1 – CRISM FRT images

- CRISM signatures in D2300 parameter low, but correlate with morphology
- The best potentially phyllosilicate bearing deposits have not been targeted in higher resolution

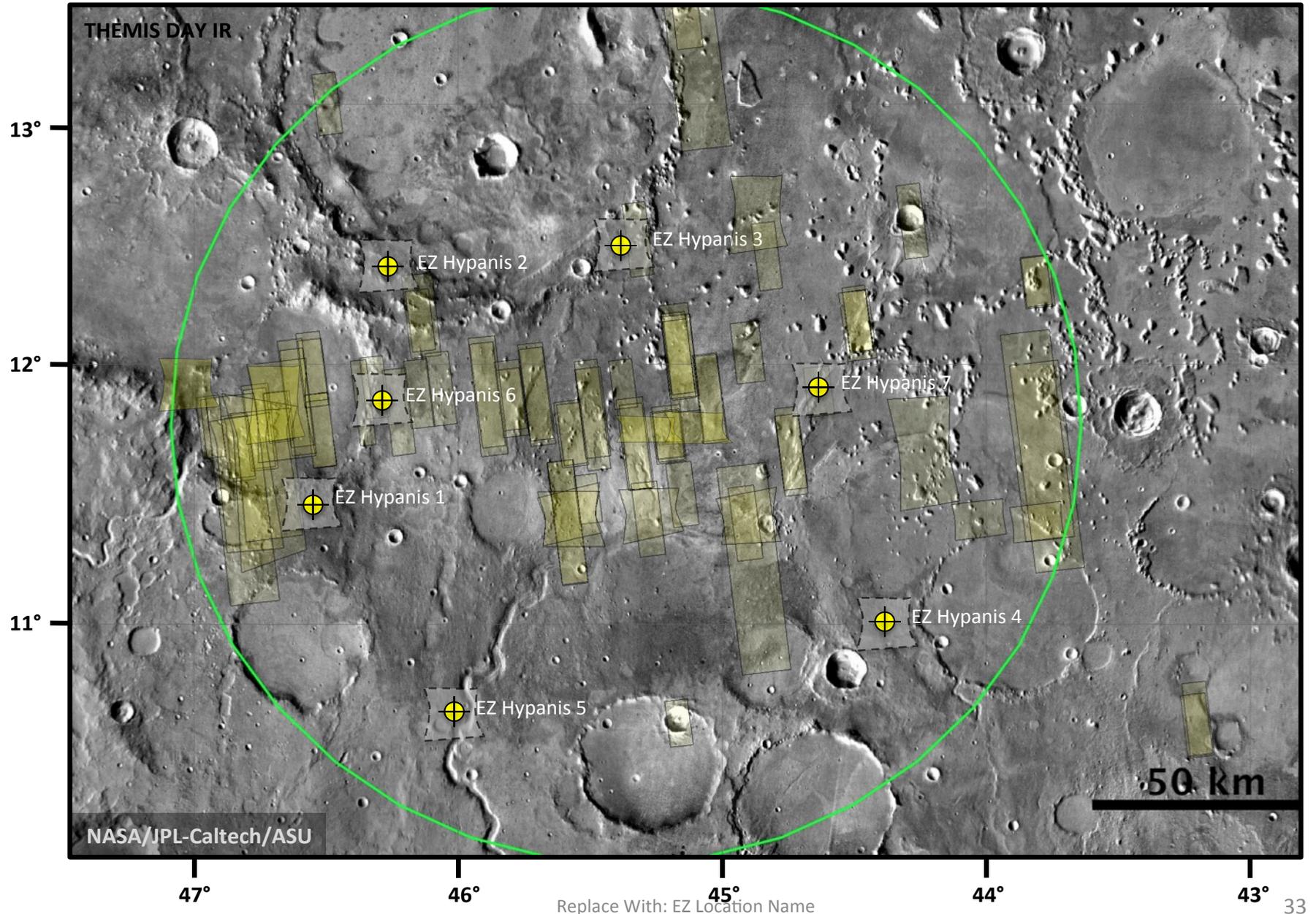
2 – HiRISE images

- Delta stratigraphy
- To couple with CRISM FRT spectral analysis of phyllosilicates

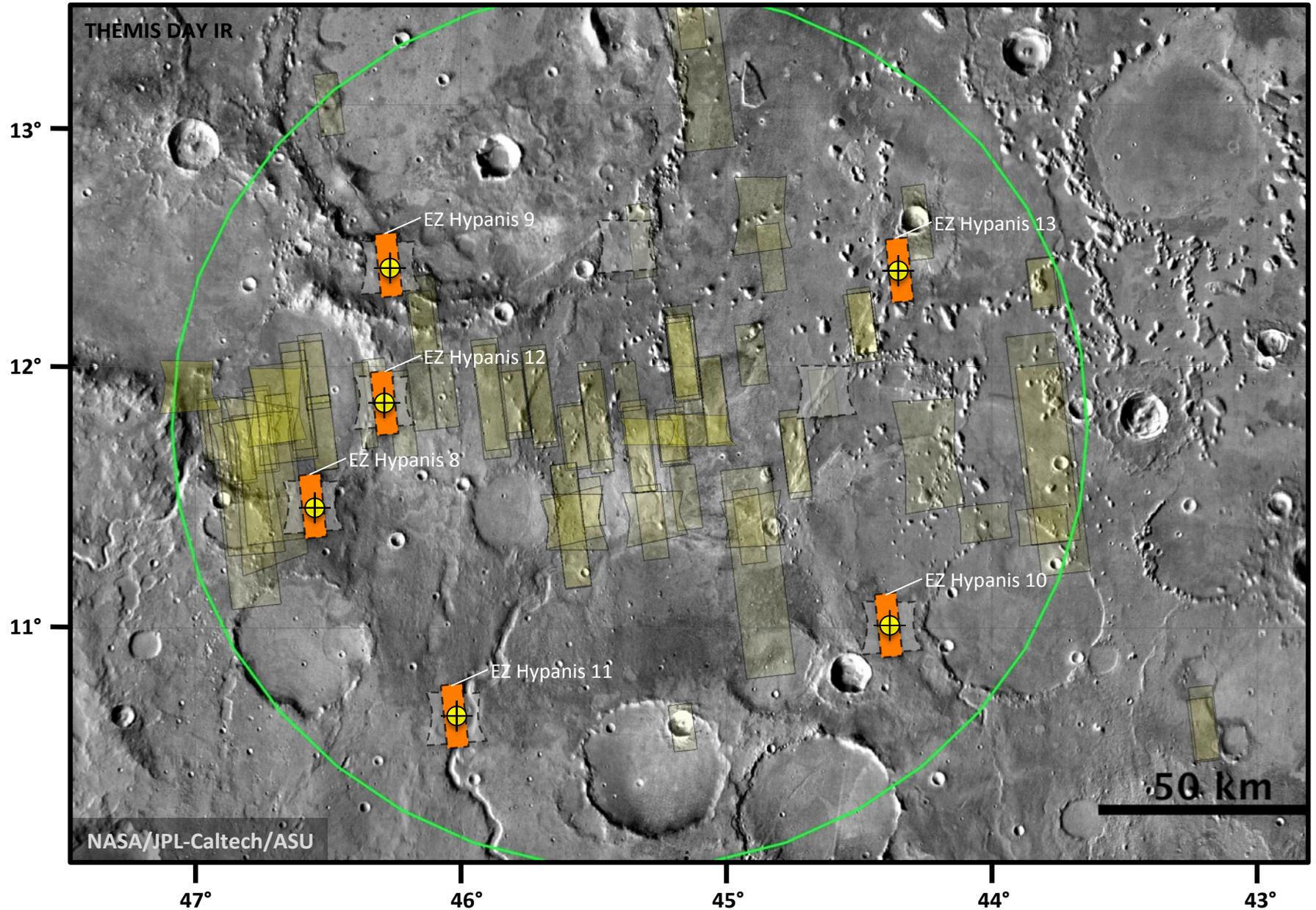
3 – CTX stereo pairs

- Delta slope analysis along proximal and distal lobes using CTX DTM

Hypanis Human Exploration Zone CRISM Requests



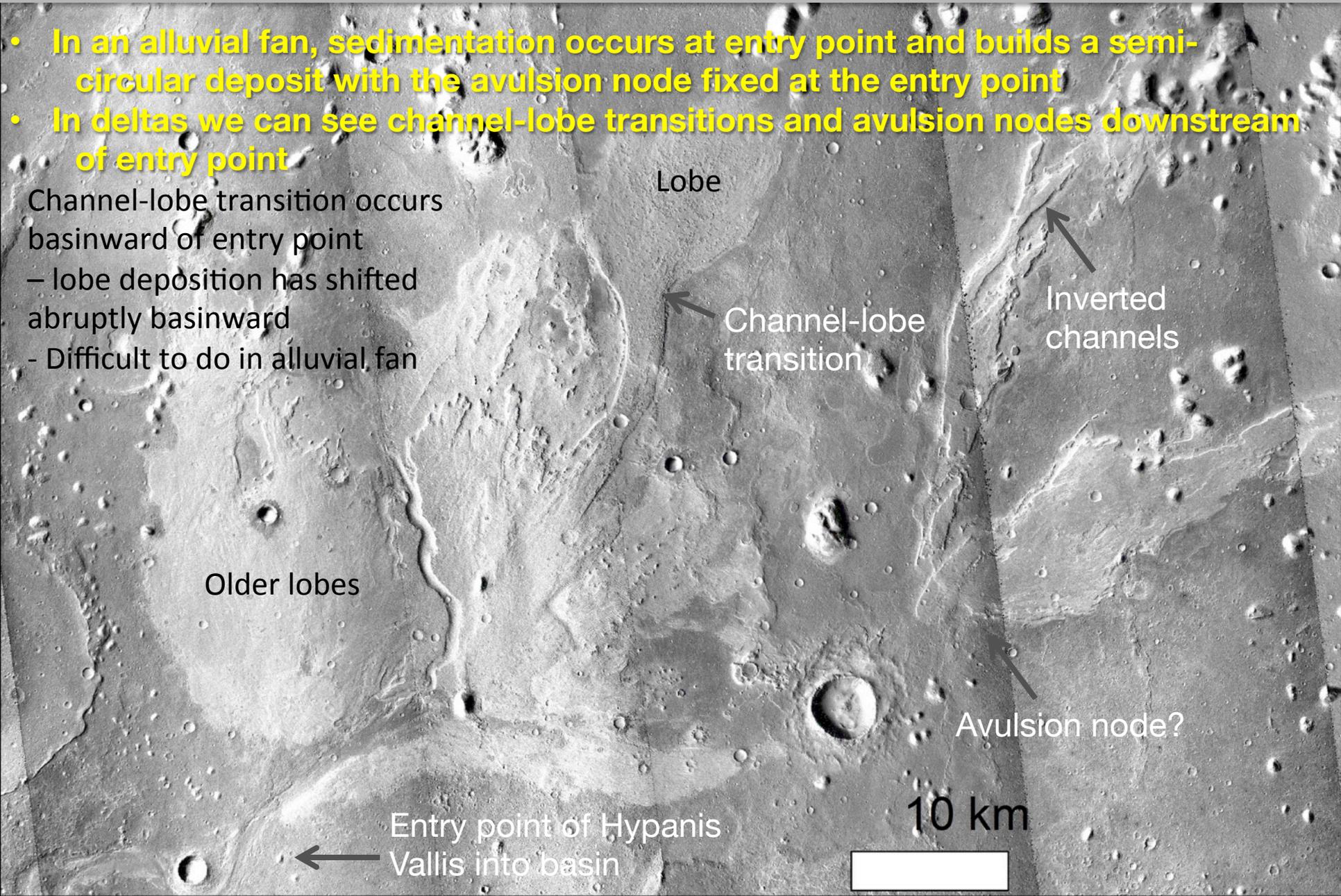
Hypanis Human Exploration Zone HiRISE Requests



Hypanis – why is it a delta and not an alluvial fan?

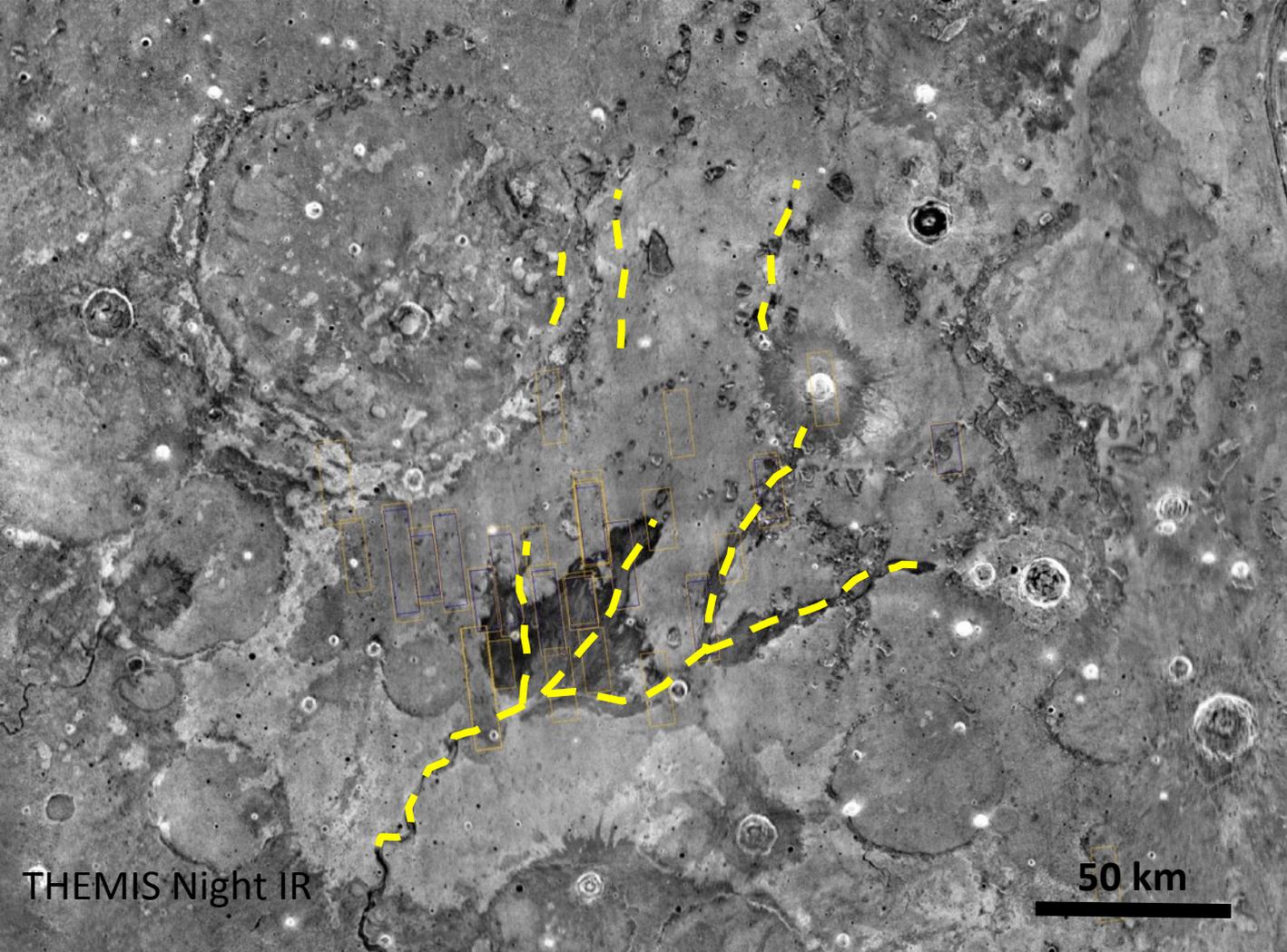
- In an alluvial fan, sedimentation occurs at entry point and builds a semi-circular deposit with the avulsion node fixed at the entry point
- In deltas we can see channel-lobe transitions and avulsion nodes downstream of entry point

Channel-lobe transition occurs basinward of entry point
– lobe deposition has shifted abruptly basinward
- Difficult to do in alluvial fan



Delta vs. Fan

1st EZ Workshop for Human Missions to Mars



- Need DTMs spanning downstream profiles
- Examine:
 - channel lobe transitions
 - avulsion node migration
 - slopes connecting farthest deposits
- Some HiRISE, CTX DTMs made
- HRSC will work, but CTX DTM of entire delta extent is better



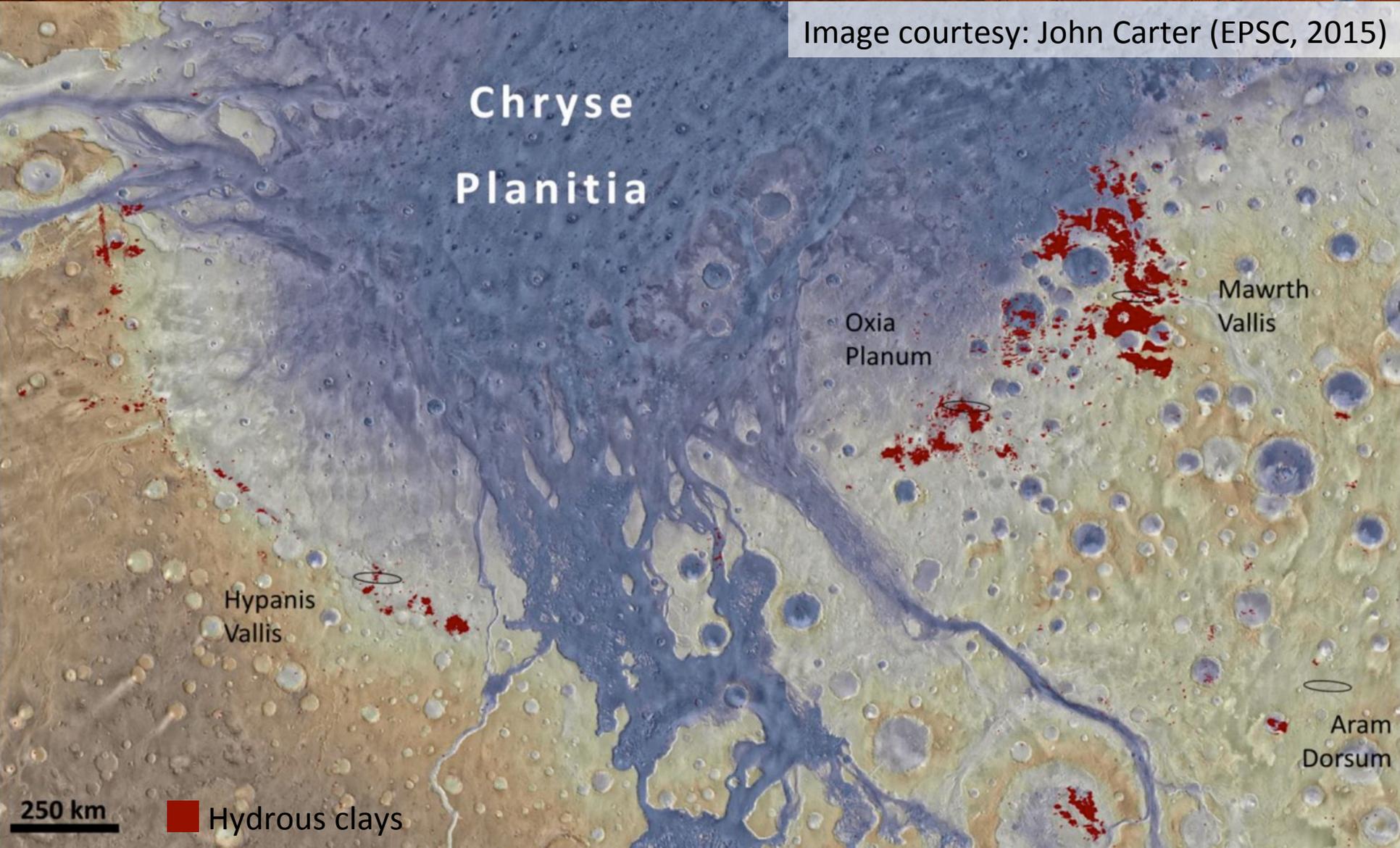
MINERALOGY

Hypanis Valles: Mineralogy - OMEGA

- Hydrated signatures (hydrous clays) at Hypanis form part of the wider, circum-Chryse basin, 'bathtub' ring of deposits → *related to large water body in Chryse/beyond?*



Image courtesy: John Carter (EPSC, 2015)



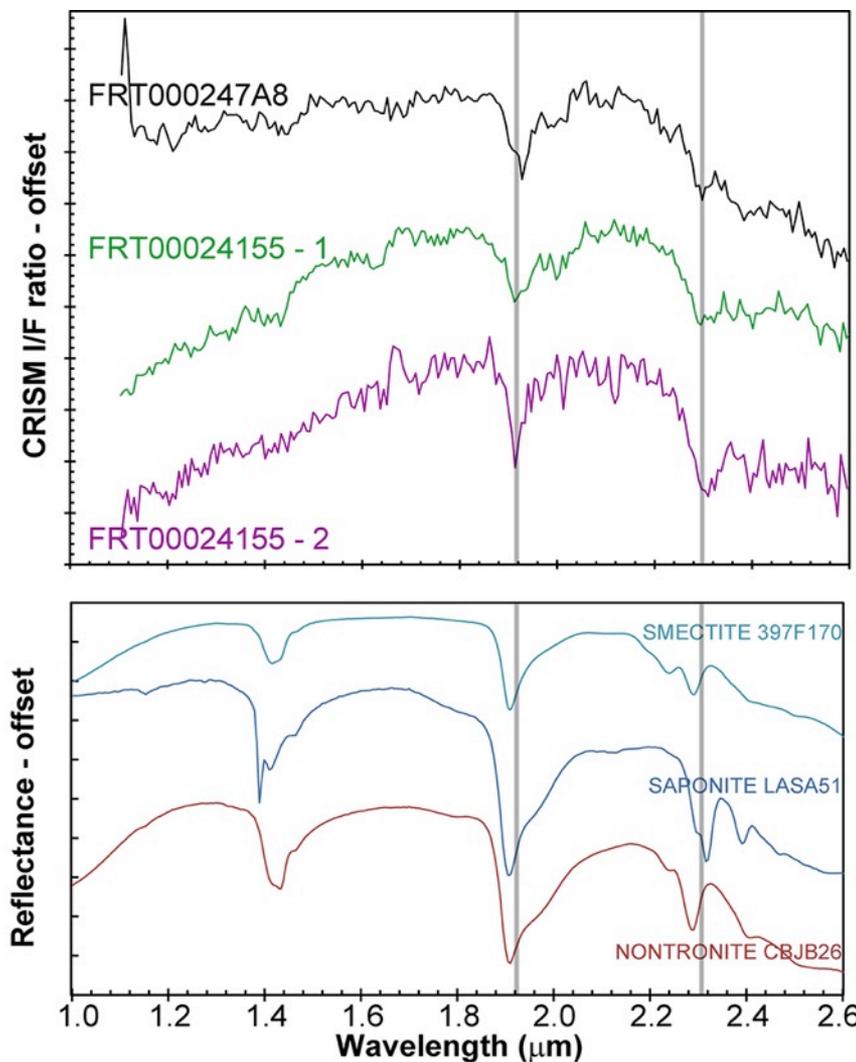
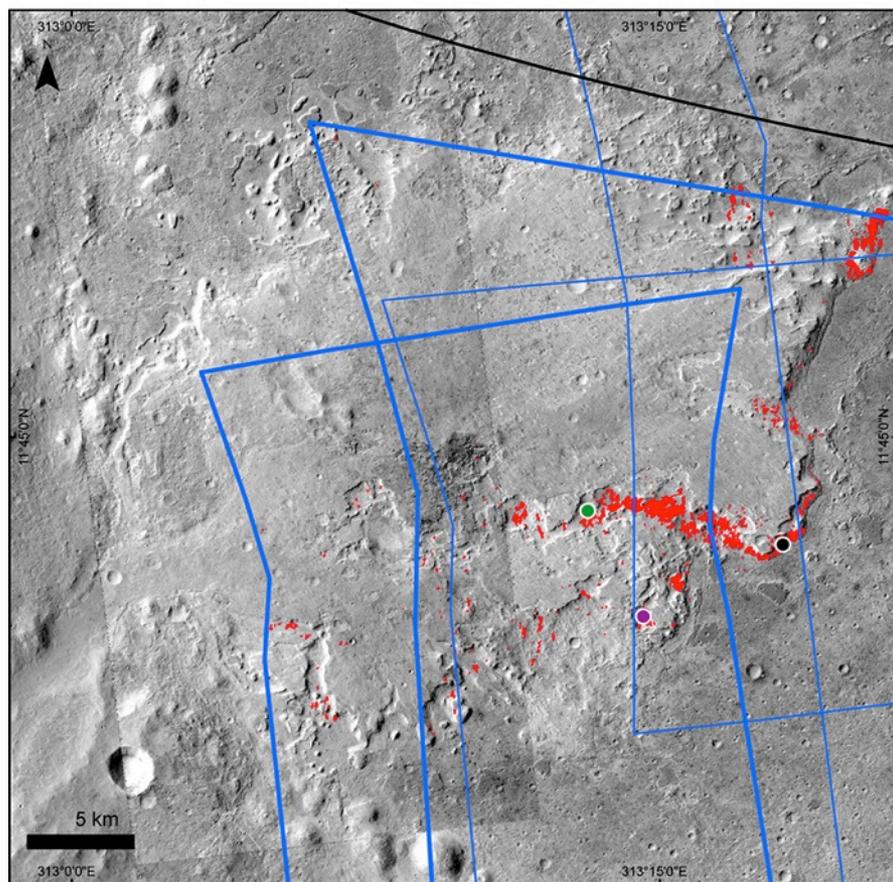
What is the bath-tub that ponded water?



Was the Chryse basin the bathtub?

Detrital Phyllosilicates

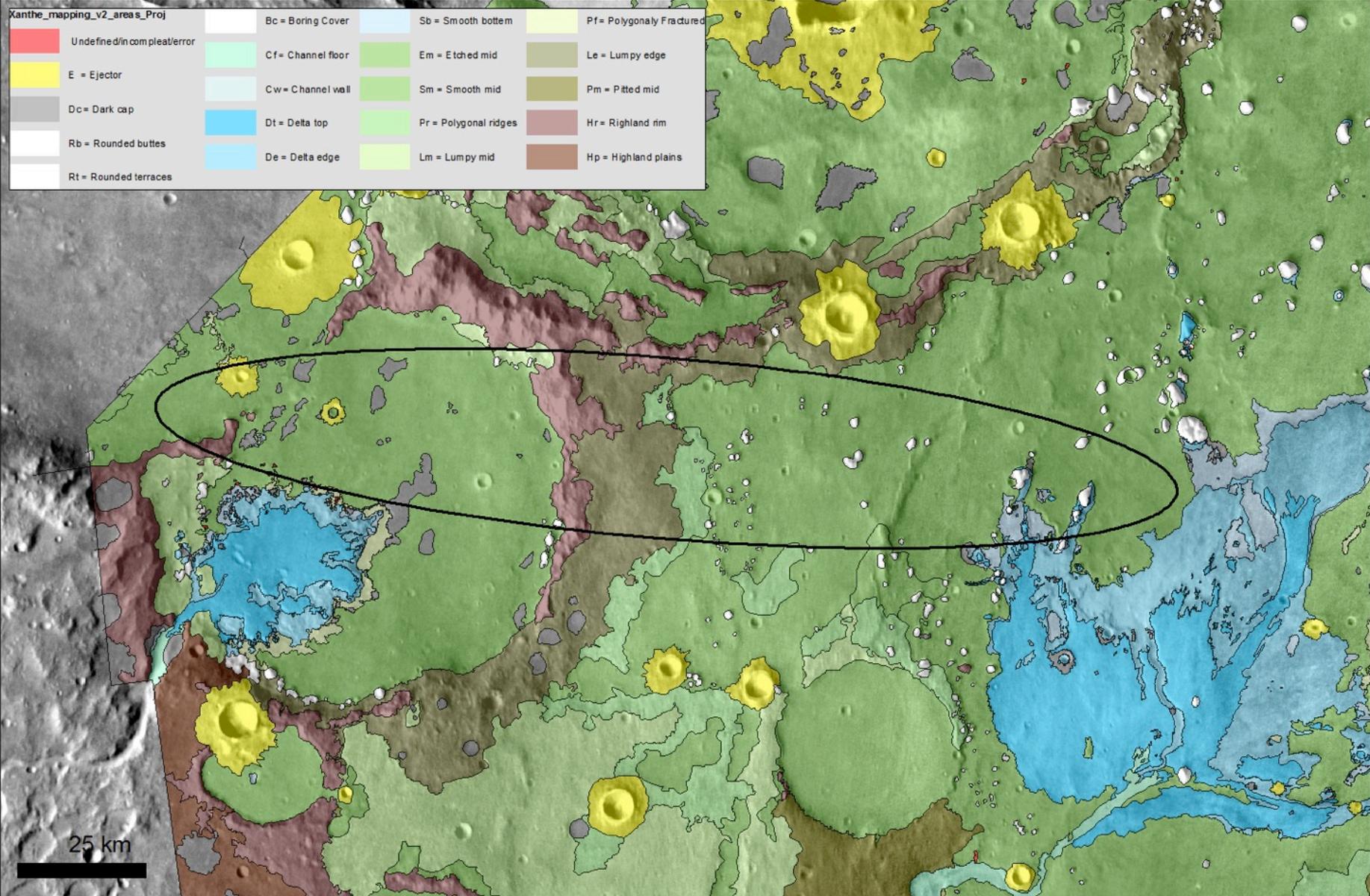
- Fe/Mg phyllosilicates (nontronite?) at low stratigraphic level in fan deposits
- Outside the ellipse to SW
- Across multiple (high SNR) CRISM observations



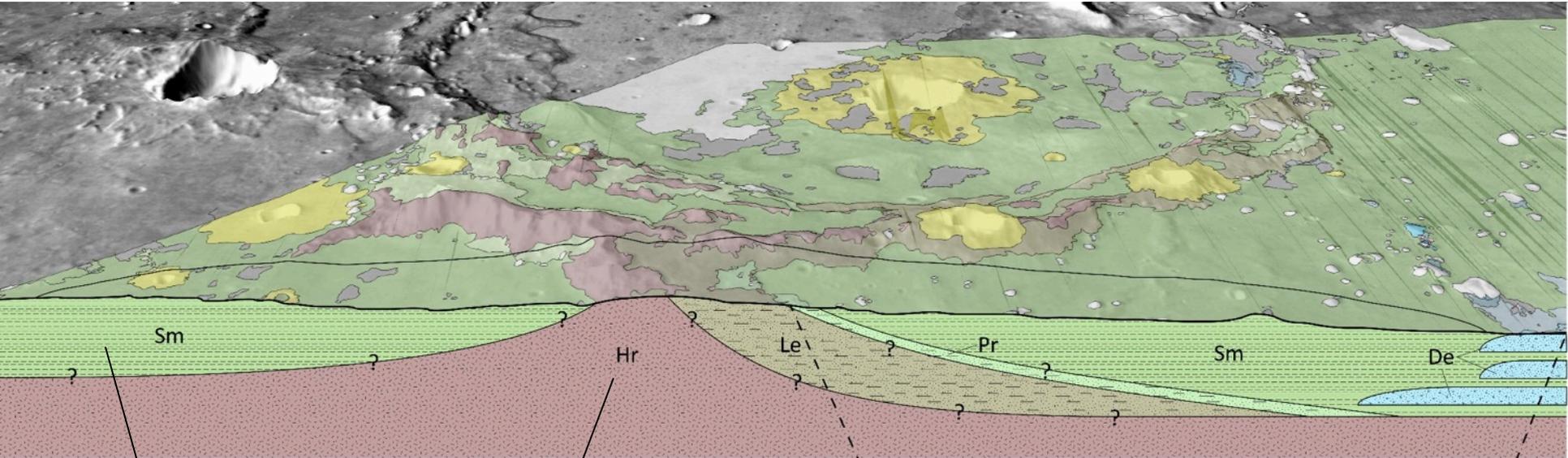


GEOLOGIC MAP

Hypanis Valles: Geological mapping – CTX basemap

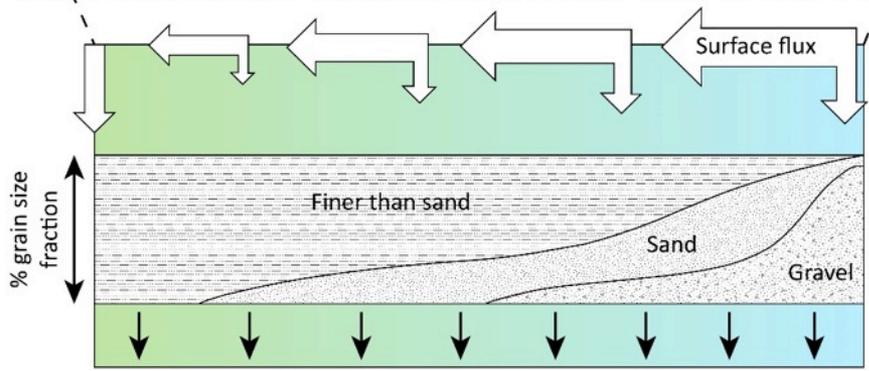


Hypanis Valles: Geological mapping – CTX basemap



Fine-grained distal
deltic-deposits

Noachian bedrock
Fe/Mg phyllosilicates



After Michael et al. (2014)

Finer ← Grain size → Coarser



BIOSIGNATURE PRESERVATION POTENTIAL

What makes deltaic-lacustrine systems good for biosignature formation and



1st EZ Workshop for Human Missions to Mars

- Rapid deposition due to abrupt change in sediment transport efficiency => rapid burial => enhanced preservation of organics
- Enhanced settling of fines (silts & clays) from suspension – include clay minerals
- Input of mineral ‘resources’ – nutrients
- Low energy environments – reduced destruction by high energy fluid flows

Delta toe deposits interfingering with basinal mudstones



Mud glorious mud.....

Preservation of deltaic deposits in the stratigraphic record



10 m

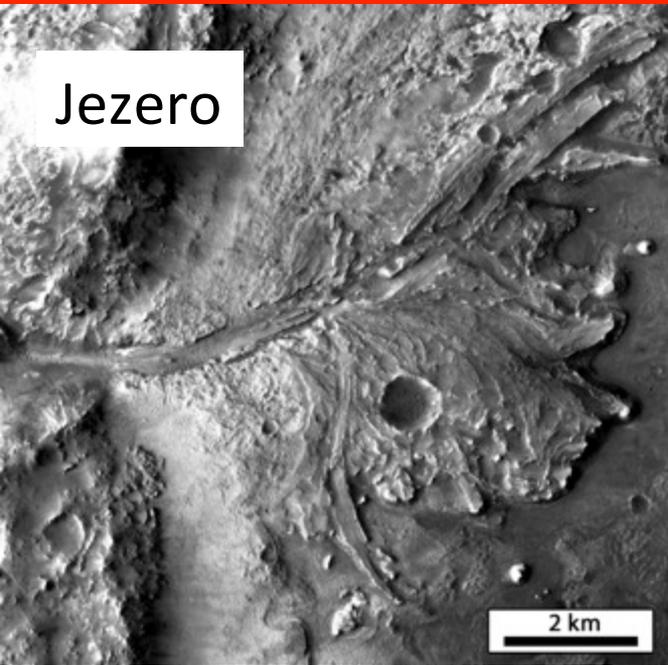
Mouth-bar
sandstones

Laminated delta
front mudstones

Tullig Cycle, Clare, Ireland,
SJ Davies

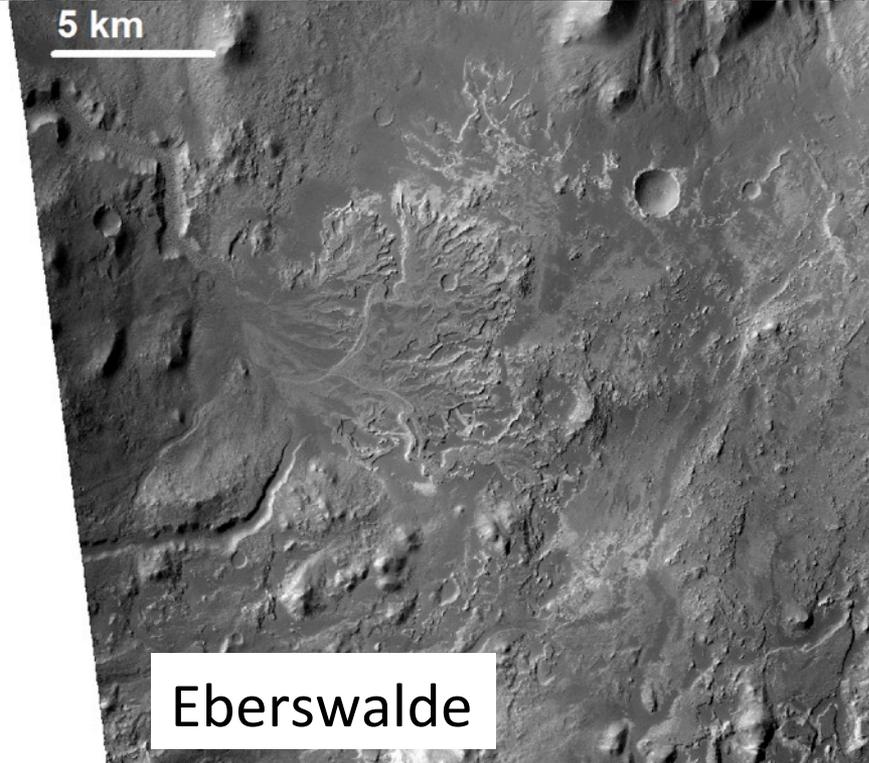
Scales of deltas

Jezero

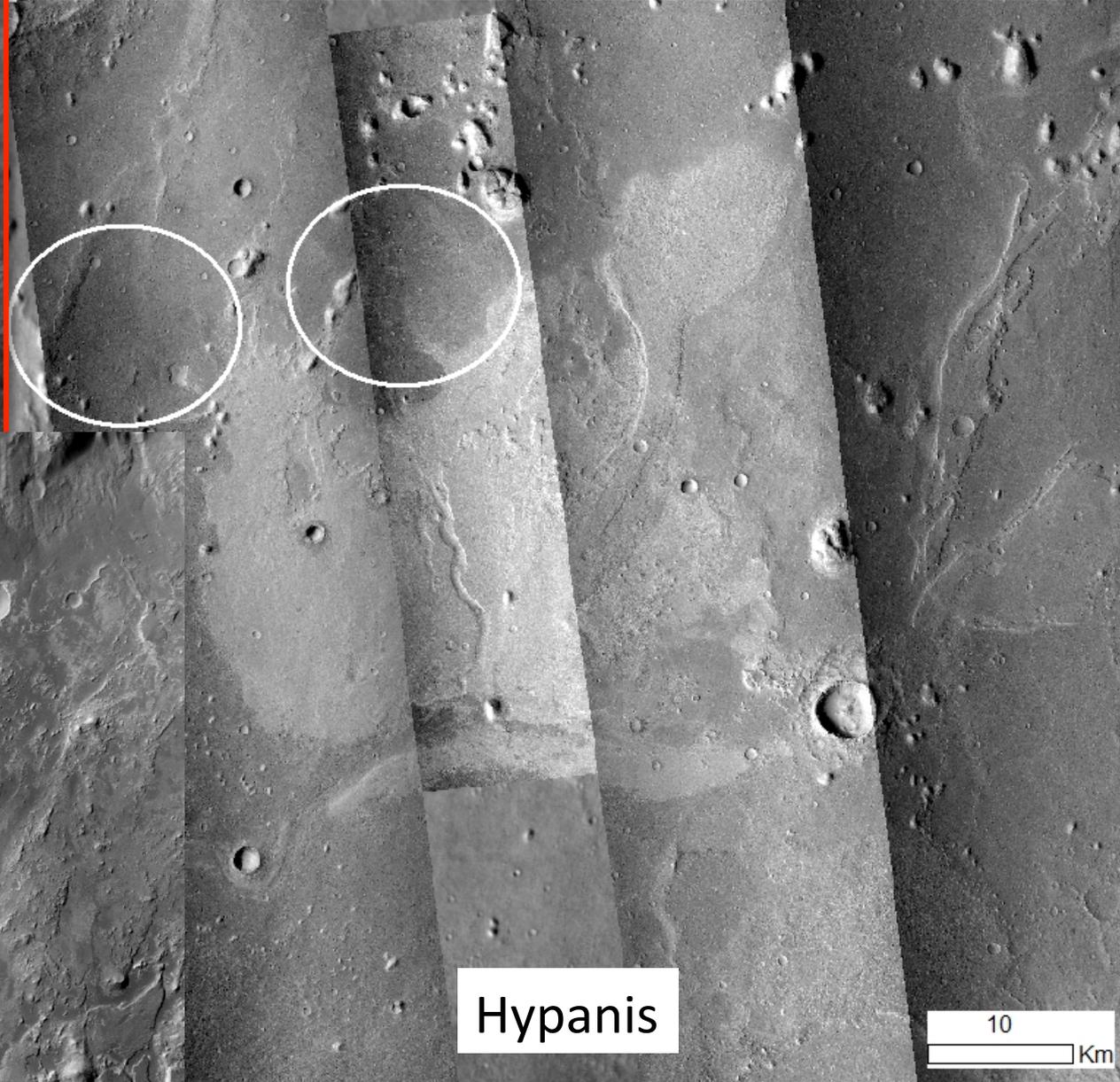


2 km

5 km



Eberswalde



Hypanis

10

Km

Hypanis Valles: *Regional Context*

- Hypanis fan system is fed by a very extensive bedrock valley – Hypanis Valles – several hundred kilometres long (?), ~75 m deep
 - Hypanis Valles connected to Nanedi Vallis? NO
- Large volume of rock has been removed by fluvial erosion
- Valley network is very different to drainages sourced from crater rims (c.f. Gale crater fan and other crater-rim fed fans)
 - Hypanis is a much bigger system
- Hypanis and Sabrina deltas are located at margin of Chryse escarpment – abrupt transition from erosional to depositional realm
- Timescale of fluvial erosion
 - Difficult to estimate precisely
 - Valley form is different to outflow channels
 - Narrow width, sinuous valley form
 - Suggests long-lived erosion – not instantaneous or short-lived erosion

Hypanis Valles: *Regional Context*

Extensive drainage systems fed Hypanis

Chryse
escarpment

Sabrina
Vallis

Hypanis delta system

Tyras
Valles

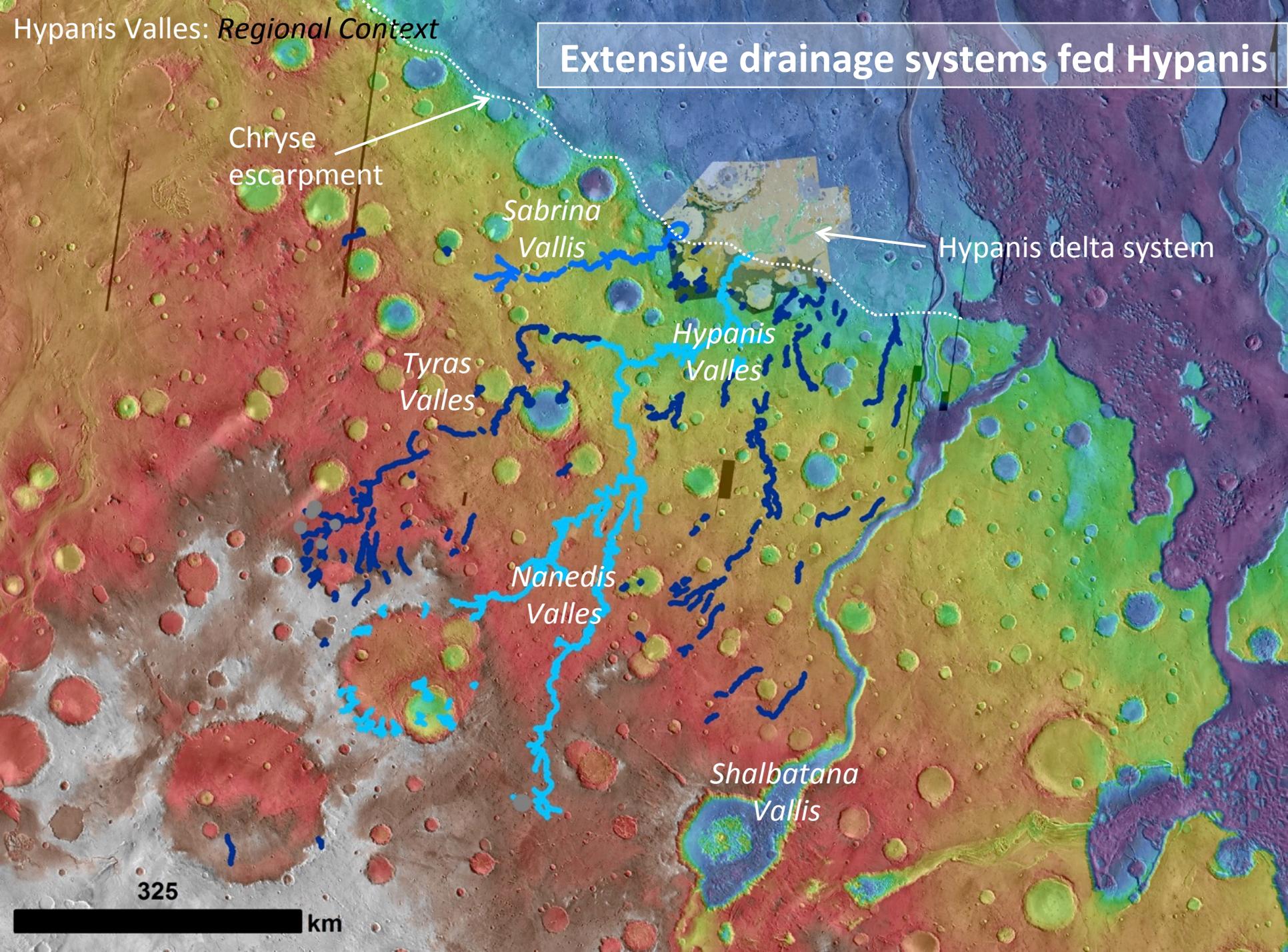
Hypanis
Valles

Nanedis
Valles

Shalbatana
Vallis

325

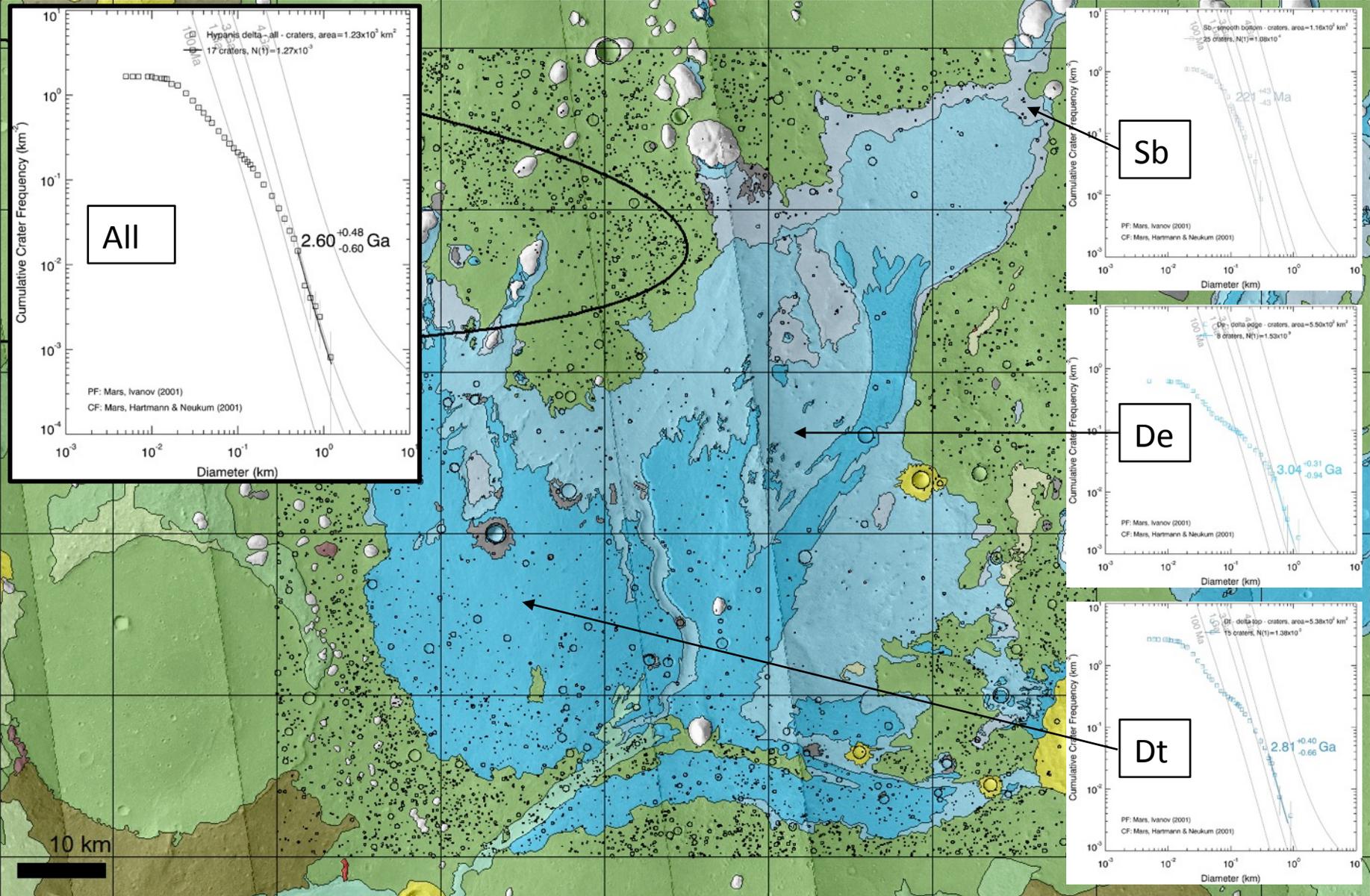
km





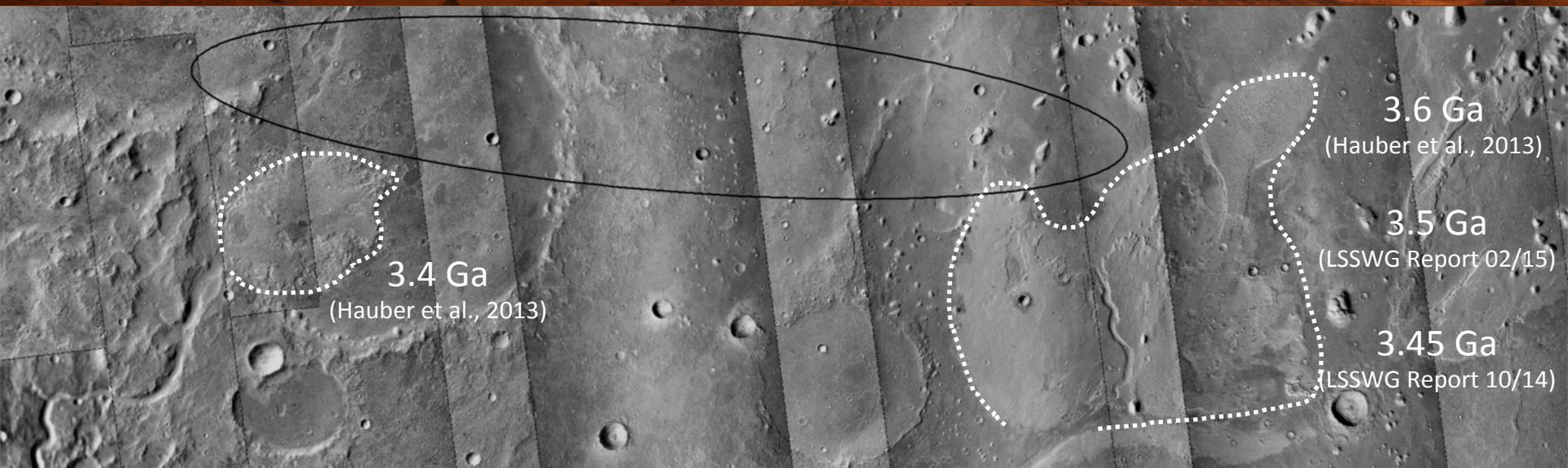
CRATER COUNT AGE DATING

Hypanis Valles: Age - local



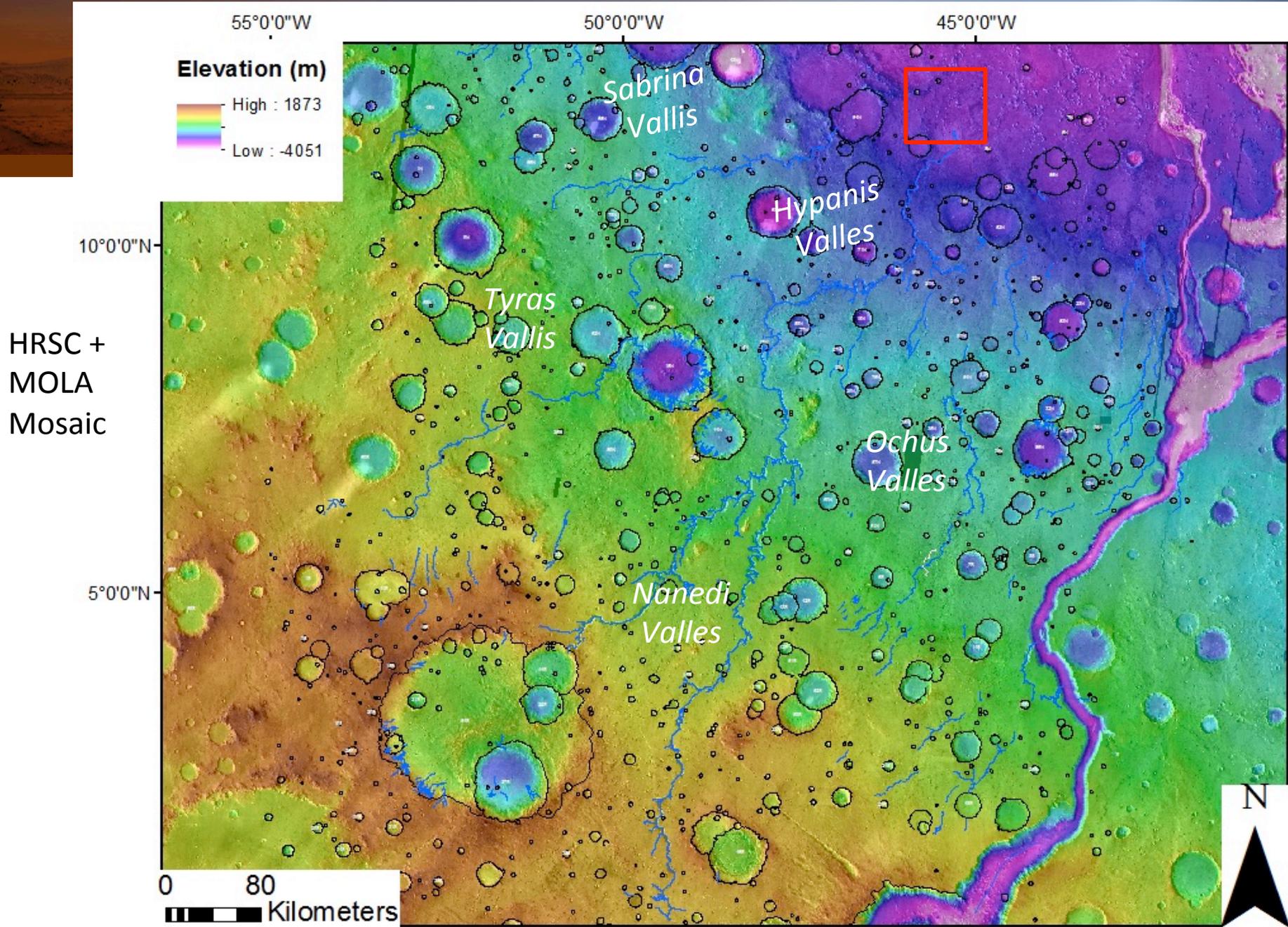
Crater Counts on Deltas? Can we trust them?

- Crater counting directly on deltas in Xanthe Terra (e.g. Hypanis Delta) has limitations.
 - Are Hypanis targets “*more recent than those at other sites*”? (LSSWG Report, 02/15)

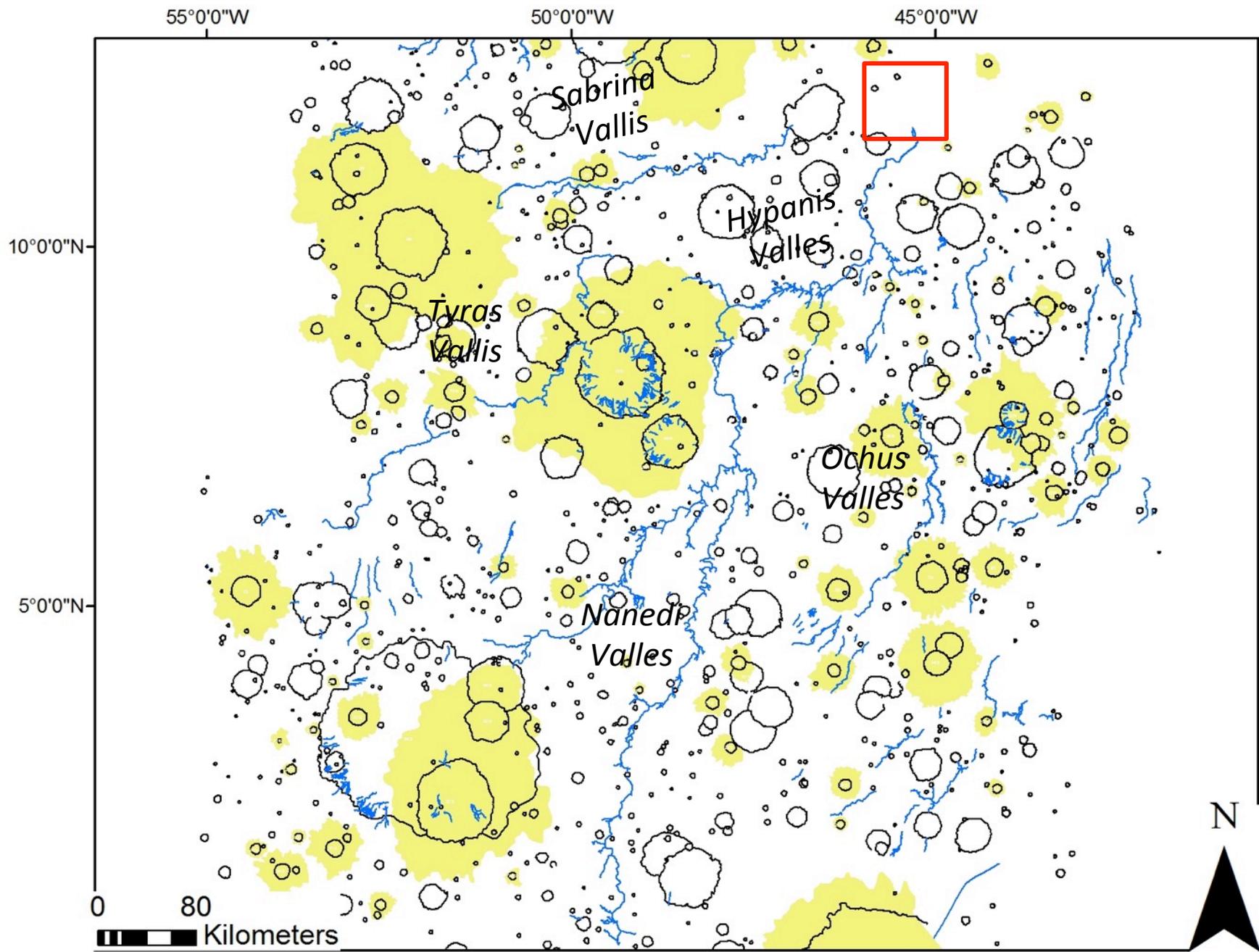


- Area of typical deltas = $10^1 - 10^2$ km², possibly too small for accurate crater counting given
 - (1) crater resurfacing
 - (2) the spatial variability of cratering as a random process.
- Deltas in this region show evidence for significant resurfacing (inverted landforms, isolated layered mesas and buttes, degraded craters) and do not preserve craters well.
- **Another more regional method of dating has also been applied.**

Hypanis Valles: Age - regional



Hypanis Valles: Age - regional



Objective: How old is Hypanis Valles and its delta?

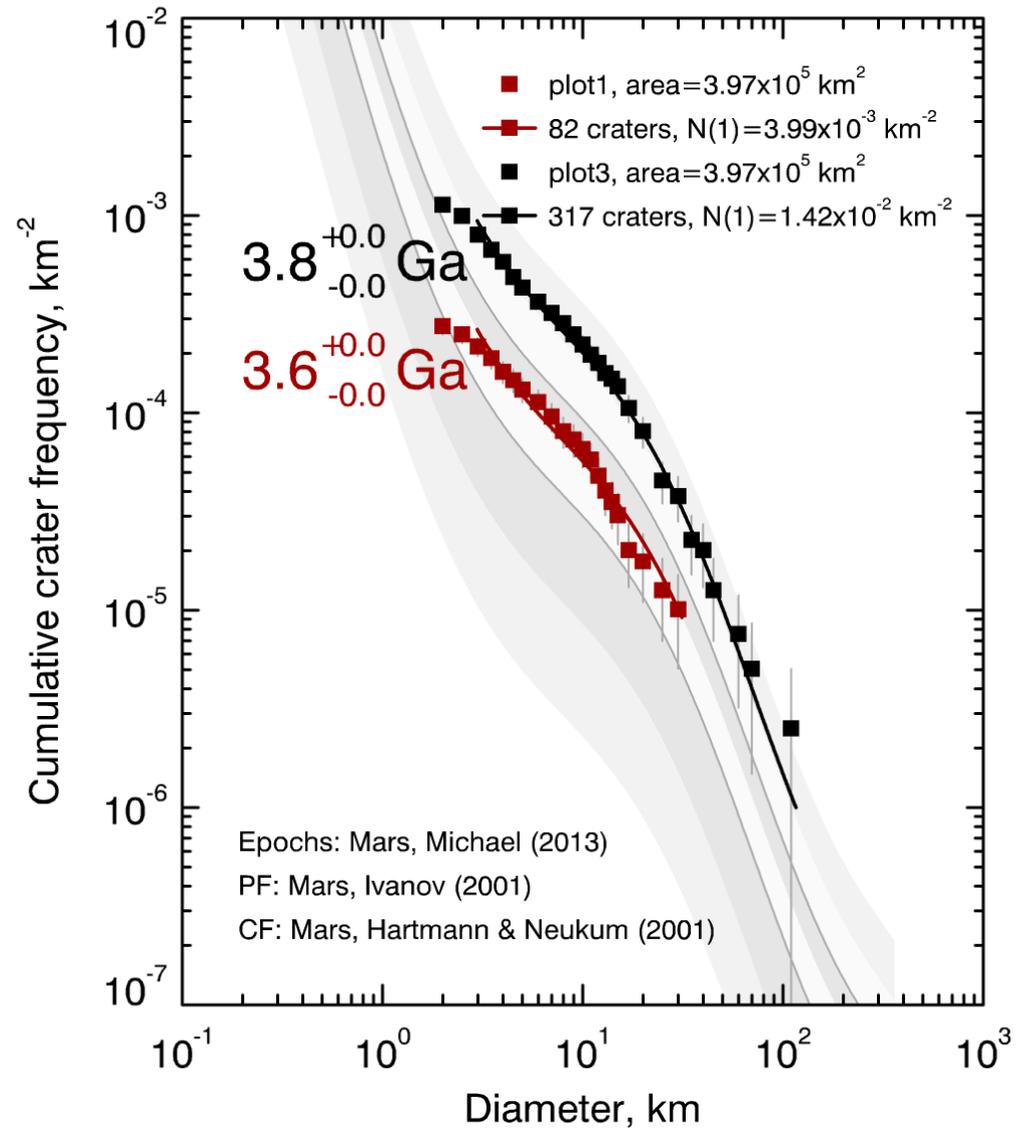
Method:

- Constructed a near complete CTX mosaic of Xanthe Terra upstream of Hypanis Delta.
- Mapped major river networks and > 2 km diameter craters at 1:40,000 using CTX.
- Our focus: relative age of impact craters compared to river networks.
- Timing methods
 1. Crater counts on ejecta that pre and post-date valley networks.
 2. Size frequency distribution of > 2 km craters at relative morphologic states (timescales of degradation of Xanthe craters).

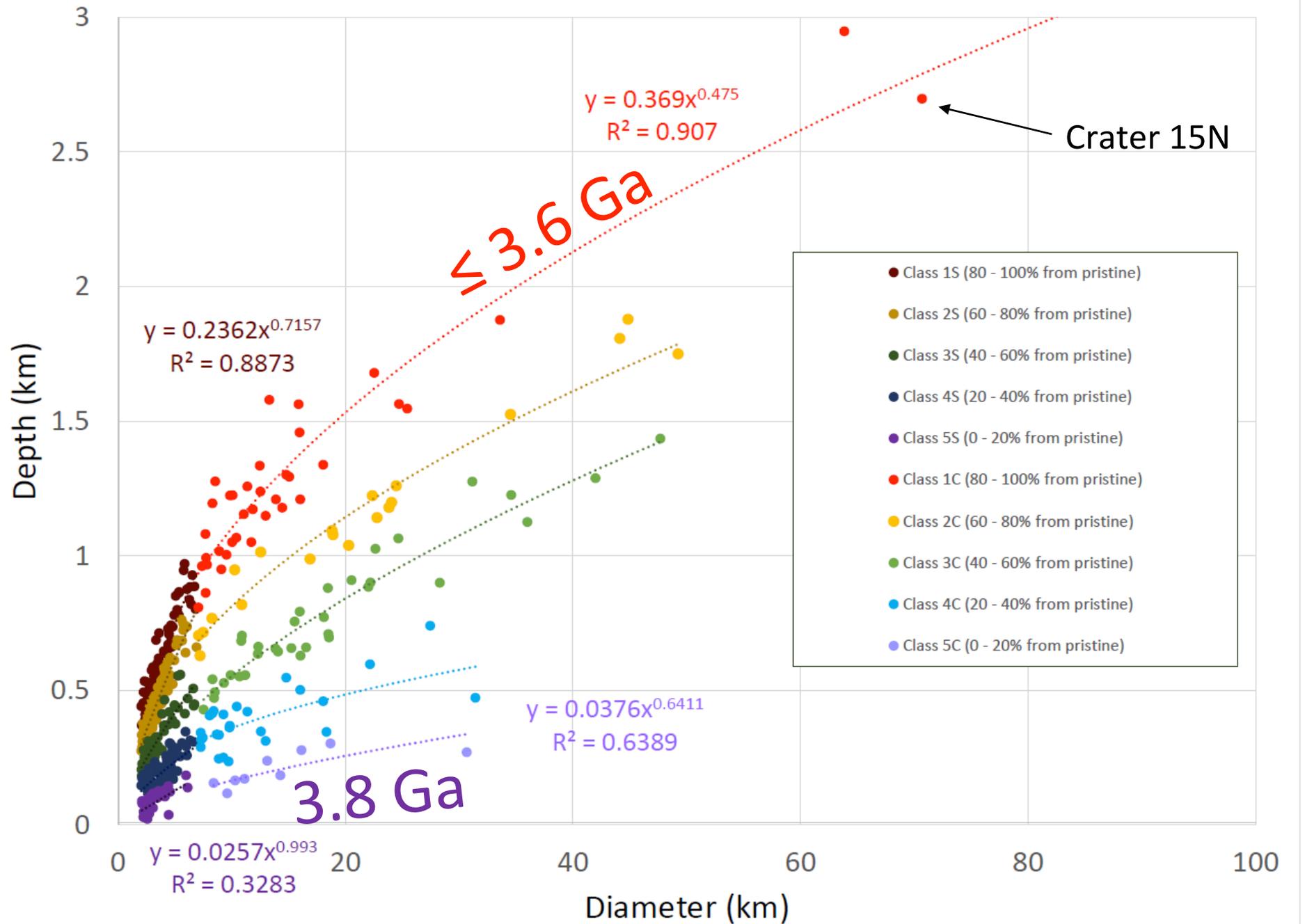


Size Frequency Distributions (SFD)

- SFD of all 317 craters = age of Xanthe Terra.
 - 3.8 Ga (Late Noachian)
- SFD of only the most pristine craters (<80% from pristine model depth)
 - 3.6 Ga (Early Hesperian)
- Consistent with rapid decrease in surface erosion rates in Hesperian.
- Most degradational work done over short interval from 3.8 – 3.6 Ga.



Hypanis Valles: Age - regional





Hypanis Valles Relative Age

1. Older than Nanedi Valles

- Suggests multiple fluvial events in this region.
- Consistent with fluvial channel superposition relationships elsewhere in Xanthe Terra.
- Nanedi Valles likely does not connect to Hypanis delta.

2. Hypanis Valles older than crater 15N

- Ejecta crater count = 3.7 Ga
- Crater degradation data: 15N is 80% from pristine d/D (≤ 3.6 Ga)
- Hypanis Valles is likely Early Hesperian or older.

3. Nanedi Valles is younger than 15 N

- Nanedi Valles is ≤ 3.6 Ga
- Suggests multiple fluvial events spanning relatively long period of time (separated by large impact events).