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Press Conference:

"SHUTTLE RETURN TO HUBBLE"

SPEAKERS:

MICHAEL GRIFFIN, Administrator, NASA
BILL GERSTENMAIER, Associate Administrator
for Space Operations
MARY CLEAVE, Associate Administrator
for Science Mission Directorate
PRESTON BURCH, Project Manager,
Goddard Space Flight Center
DAVID LECKRONE, Senior Project Scientist,
Hubble Space Telescope Project

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[TRANSCRIPT PREPARED FROM A TELEPHONIC RECORDING.]

1 P R O C E E D I N G S

2 MR. ACOSTA: Good afternoon, and welcome to our
3 Shuttle Return to Hubble press conference.

4 Let me introduce our participants today. To my
5 left, the NASA Administrator, Dr. Michael Griffin; to his
6 left, the Associate Administrator for Space Operations,
7 Bill Gerstenmaier; to his left, Science Mission Directorate
8 Associate Administrator, Dr. Mary Cleave. To her left, we
9 have Preston Burch, and then to his left, Dave Leckrone.

10 All right. We are going to go through some short
11 prepared remarks by our individuals and participants, and
12 then we will open it up to questions. I ask that you
13 please identify yourself and identify who your question is
14 for before asking it.

15 All right. Dr. Griffin?

16 ADMINISTRATOR GRIFFIN: Well, I don't really have
17 any prepared remarks. So I will skip on down the line, and
18 we can go to Gerst.

19 MR. GERSTENMAIER: Okay. I don't have any
20 prepared remarks either. So we will just move on down the
21 line.

22 [Laughter.]

1 MR. ACOSTA: I shouldn't have said prepared
2 remarks. Just remarks in general, if you have anything to
3 say.

4 MR. GERSTENMAIER: I would like to say it is
5 great to be here representing the Shuttle team that is
6 ready to go do this mission. Again, I think through the
7 Return to Flight activities that was just done, we have
8 proven the basic concepts and the basic techniques that are
9 needed to go do the Hubble mission, and it is great to be
10 able to be here today and to be able to be in the process
11 of starting the plan and starting to train for the Hubble
12 mission that is coming up.

13 MR. ACOSTA: Mary?

14 DR. CLEAVE: Well, on behalf of the scientists in
15 the area, we are thrilled at the opportunity to potentially
16 have 4 more years of science out of Hubble, brilliant
17 science out of Hubble, looking forward to the new
18 instruments, looking forward to potential repair, and I
19 would just like to thank Bill and his team and Mike and
20 everyone for providing this opportunity. We really didn't
21 think it was going to happen, and we are thrilled that it
22 is.

1 MR. BURCH: I can say that looking back on the
2 last 4 years, without reservation, today is my happiest day
3 to be at the office, and I think the same is true for
4 everyone who works on the Hubble Space Telescope.

5 We owe an enormous debt of gratitude to the folks
6 at the Johnson Space Center, Kennedy Space Center, the
7 Marshall Space Center and others for their hard work in
8 making the needed improvements to the Space Shuttle, so
9 that we can now go back to Hubble safely and make the
10 upgrades and repairs that will enable us to continue our
11 amazing voyage of the scientific discovery with the Hubble
12 Space Telescope.

13 This next servicing mission can be likened to
14 those extreme make-over reality shows on TV that are so
15 popular today.

16 On Servicing Mission 4, we are going to give
17 Hubble another extreme make-over. This make-over will be
18 the best one yet because we will outfit Hubble with the
19 most powerful and advanced imaging and spectrographic
20 instruments available, and we will extend Hubble's
21 operating lifetime for 5 additional years, which should
22 keep us operating until well into 2013 and possibly longer.

1 Work on these new science instruments and the
2 replacement hardware has been underway for several years.
3 The Congress and our NASA Administrator asked the Hubble
4 team to continue working on the mission during the past few
5 years to enable us to do this mission as soon as possible,
6 if and when the Space Shuttle became ready again, which we
7 now know it will be.

8 All of the Hubble servicing missions make
9 extensive use of the Shuttle's enormous cargo bay and
10 spacewalking astronauts. Servicing Mission 4 will be the
11 heaviest servicing mission to date. We will be carrying
12 approximately 22,000 pounds of hardware up to Hubble to do
13 Servicing Mission 4.

14 We will be using four carriers inside the Shuttle
15 cargo bay to carry all the new science instruments,
16 replacement hardware, and tools for the astronauts, and to
17 attach Hubble to the Shuttle while the astronauts are
18 working on it.

19 One of these carriers utilizes an advanced design
20 and composite materials to save weight, so that we can
21 carry even more equipment.

22 We will install two new science instruments on

1 Servicing Mission 4, the Wide Field Camera 3 or with C-3 as
2 the bullet, and the Cosmic Origins Spectrograph, the COS.

3 We will also change out all six gyroscopes,
4 change out all six batteries and add thermal insulation
5 under the mobiles, new outer blanket layers, to three of
6 the electronic space on Hubble.

7 One of the fine guidance sensors will be
8 replaced. An over-voltage protection device will be
9 installed, and a flow-up capture mechanism will be attached
10 to the back end of Hubble to facilitate the future
11 deorbiting of Hubble at the end of its useful life.

12 Finally, we will attempt to repair one of
13 Hubble's science instruments, the Space Telescope Imaging
14 Spectrograph, or STIS, which stopped working in August of
15 2004.

16 Servicing Mission 4 will be similar to the
17 previous Hubble servicing missions.

18 Servicing Mission 4 will require four astronauts
19 to do spacewalks over a period of 5 days. The astronauts
20 will operate in teams of two, and the teams will go
21 spacewalking on alternate days, so that they can rest up
22 between spacewalks.

1 Each day of spacewalking is planned for 6 hours
2 of work, upgrading and maintaining Hubble. This work is
3 physically and mentally very demanding, and it requires a
4 lot of planning, training, and other preparations.

5 Each day after the new hardware is installed and
6 the astronauts return to the Shuttle's crew cabin, the
7 Hubble team on the ground of Goddard Space Flight Center
8 will perform tests to ensure that everything works
9 properly.

10 Although we do a lot of astronaut training at
11 Goddard using the actual tools of a flight, hardware to
12 training the astronauts, we also do a lot of training at
13 the Johnson Space Center and facilities such as the neutral
14 buoyancy laboratory, which you can see here on the video,
15 to simulate the environment in orbit.

16 The attempt to repair the STIS (Space Telescope
17 Imaging Spectrograph) will be the first time on Hubble that
18 we will open up a complex piece of hardware to fix it in
19 situ. This will be a very intricate piece of work. The
20 Hubble team has built some clever and unique tools to meet
21 the challenges of this task, and the astronauts worked
22 closely with the Hubble servicing engineers to make it

1 feasible to repair STIS within one EVA day.

2 When SM-4 is completed in 2008, Hubble will be
3 ready to probe deeper into the universe to make more
4 exciting discoveries.

5 Here to tell us about the exciting future science
6 that will be enabled by Servicing Mission 4 is Hubble's
7 Chief Scientist, Dr. David Leckrone.

8 DR. LECKRONE: Thank you, Preston. Good
9 afternoon.

10 The first thing I want to do on behalf of the
11 literally thousands of scientists around the world who used
12 the Hubble Space Telescope regularly is to give our thanks
13 and gratitude to NASA, to the Administrator, to the Johnson
14 Space Center, to the people who work on the whole program
15 every day, to have stuck with it so long and to have worked
16 so hard to lead to this very happy day for the future of
17 science.

18 I would like to briefly talk about why we are
19 doing Servicing Mission 4 from a scientific point of view,
20 what is important about this.

21 It is pretty simple to answer that. Up to this
22 point, Hubble has not approached the limits of what it is

1 capable of doing, and with the new instruments that we are
2 going to put on board and the restoration of this fifth
3 instrument and buying ourselves an additional 5 years of
4 lifetime, I think we can begin to approach the ultimate
5 limits of what Hubble was originally designed to do.

6 At the conclusion of this mission, after SM-4 is
7 over and the astronauts have left, Hubble will literally be
8 at the apex of its capabilities. There will not have been
9 a time in its history when it has been as capable as it
10 will be at that moment.

11 For one thing, we will have on board, we hope,
12 six fully functioning scientific instruments for the first
13 time since 1993. The Hubble is a general purpose, public
14 facility, observatory, and it takes a whole tool bag full
15 of instruments to provide the various tools needed to
16 attack all kinds of astronomical problems, and that is the
17 way Hubble has worked, and that is one thing that has made
18 it so successful, its versatility.

19 Two of those six instruments will be the crown
20 jewels, in my opinion, of this mission, the two new
21 scientific instruments, the Cosmic Origins Spectrograph and
22 the Wide Field Camera 3.

1 Cosmic Origins Spectrograph is the most sensitive
2 ultraviolet spectrograph when placed behind Hubble optics.
3 You have to have a good telescope in front of it, but then
4 it is the most sensitive scientific spectrograph ever flown
5 in space.

6 The Wide Field Camera 3 will enhance our survey
7 capabilities on Hubble. Today, we have a wonderful camera,
8 the Advanced Camera for Surveys, that was designed, as its
9 name suggests, for that kind of objective, and I will show
10 you an example coming from that in a moment, but it is
11 limited primarily to visible and red light, light that you
12 can see with your own eyes. With the Wide Field Camera 3,
13 we will take that same surveying power and efficiency and
14 extend it all the way from the ultraviolet out through the
15 near infrared, of about 200 nanometers up to 1,700
16 nanometers, continuously covered with wide field, very
17 precise, very high-resolution imaging.

18 What I would like to do in closing here is to
19 offer two examples, one from each of these instruments, of
20 what will make them so powerful.

21 So we have a video to show here for a moment, and
22 if you in your mind's eye could take an imaginary trip

1 across the present-day universe, this is what it would look
2 like. You notice it is not uniform. It is not homogenous.
3 It has the structure of what we call a web or cosmic web,
4 sort of like a three-dimensional spider web. This is what
5 is observed in surveys from the ground and elsewhere. This
6 is what it theoretically predicted is coming out of the big
7 bang.

8 What you see are filaments and filaments
9 intersecting each other and voids amongst the filaments in
10 this web-like structure.

11 The underlying skeletal structure here is
12 composed of dark matter, and it is the gravity of dark
13 matter that creates this structure, but coming along for
14 the ride gravitationally is voluminous matter, galaxy,
15 stars, and the gas between, and basically, the Cosmic
16 Origins Spectrograph is intended to probe through the
17 individual filaments or through the voids in the cosmic web
18 and determine not only physical properties, but how those
19 properties have changed with time, chemical composition,
20 the pressures, densities, velocities, and the overall
21 underlying structure of the cosmic web.

22 It will use the background, as background light

1 sources, quasi-stellar sources or quasars, essentially as
2 cosmic flashlights shining through the cosmic web, and then
3 on this, the receiving end, that Earth, COS will take that
4 light and spread it out and analyze it.

5 The second example -- and this pertains to the
6 Wide Field Camera 3 -- I mentioned our Advanced Camera for
7 Surveys, and we have a graphic that shows a fairly,
8 routinely acquired survey field with the ACS in the sky,
9 but embedded in this particular field in the sky is the
10 deepest image across the cosmos ever made by humans, and it
11 is the very famous Hubble Ultra Deep Field.

12 If we could bring that up next, I would
13 appreciate it. There we are.

14 As I say, this is the deepest core sample ever
15 acquired, and it required about a million seconds of
16 exposure time, about 400 orbits of telescope time to
17 acquire this, but this is a visible light image. To go
18 deeper than this image goes requires this to go to the near
19 infrared.

20 Let me show you an example of what has come just
21 from the visible light image. This is the same field with
22 the little blue squares in it, and you see the little

1 postage stamp images to the right there.

2 Over 500 proto galaxies have been discovered in
3 this image that emitted their light when the universe was
4 between 700 million and a billion years old. They are
5 among, or probably are, the most distant proto galaxies
6 ever observed.

7 Again, to go deeper than that, we would have to
8 go from visible light to infrared light because the
9 universe is expanding, stretching the wavelength of light
10 and so forth.

11 This final graphic is a little cartoon that
12 illustrates why this is important. Think of time as moving
13 from the right to the left, and over on the right, just off
14 scale, is the big bang, which I haven't tried to display
15 here, and then there is a period of time in the history of
16 the universe where it was basically a dark place. It was
17 filled with warm hydrogen and helium gas, but no sources of
18 light as yet, and in any event, light had not passed
19 through that very opaque medium, so Dark Ages as we call
20 them.

21 You will notice on the bottom label, two times,
22 400 million years after the big bang and 700 years after

1 the big bang. Right now, those are our best estimates of
2 the interval of time during which the universe went from
3 darkness to light. Essentially, that is the period of the
4 nursery phase of the universe, when the very first stars
5 were being formed, congealing into galaxies, into clusters
6 of galaxies, and those stars and galaxies heated up the
7 surrounding material, allowing it to go from being opaque
8 to being transparent, so that light could pass through, and
9 then we see the luminous, brilliantly lit universe that we
10 see today.

11 Now, those 500 images that I showed you a sample
12 from in the prior illustration from the Hubble Ultra Deep
13 Field really sampled going from left to right to about that
14 700-million-year point. So we are just starting with
15 Hubble to push into this region of re-ionization. We want
16 to go deeper, and with the WF-3, the Wide Field Camera 3
17 and its infrared channel, we will be able to push even
18 deeper, poking into this region.

19 Now, why is this very important? We are the
20 scouting parties in the James Webb Space Telescope. The
21 latter part of 2013, James Webb will fly. It will be a
22 significantly larger telescope than Hubble, and it will go

1 even farther into the infrared than Hubble can go. What we
2 are hoping to do with this kind of a survey with Hubble is
3 to set the foundation, set the stage to help guide the
4 James Webb as it goes off into this uncharted region.

5 Thank you.

6 MR. ACOSTA: Thank you, Dave. Good lesson there
7 for us.

8 All right. Let's start off with questions, and I
9 ask that you please identify yourself and who your question
10 is for.

11 We will start off on the front row here and go
12 right here.

13 QUESTIONER: Hi. Jeff from [inaudible] magazine.
14 I guess this question may be for Bill Gerstenmaier.

15 You described what is involved with having a
16 longtime need capability. Are we talking about having a
17 second Shuttle sitting on another pad, or what exactly are
18 you going to have to do?

19 MR. GERSTENMAIER: Okay. In terms of Launch on
20 Need, our current planning is we would have another orbiter
21 ready to fly on Pad B, on 39-B, ready to launch in case
22 some problem occurred that was not repairable or needed the

1 rescue flight, and then it is available in a standby mode.

2 If it is not used for the rescue flight, then
3 that same orbiter would then be outfitted with Space
4 Station equipment and would then fly to Space Station kind
5 of its normal mode. So that is kind of the ground
6 configuration.

7 The challenge is that we were planning on passing
8 that pad off to the exploration group to start modifying
9 that pad for Constellation for a demonstration flight in
10 2009. So we are still going to do that. We are still
11 going to give them the pad to start working on, but they
12 are going to have the constraint now that they have to keep
13 that pad essentially in launch-ready configuration.

14 So some of the modifications they were going to
15 do, they are going to have to cut back and not make some of
16 those mods, but they will still be able to do a lot of the
17 work to get prepared for the exploration missions that are
18 coming down the road.

19 So, again, we are still doing the planning for
20 that, but that is the basic concept that we are putting
21 together for the Launch on Need pad configuration.

22 MR. ACOSTA: Seth?

1 QUESTIONER: Seth Borenstein, Associated Press,
2 for Preston Burch.

3 I know you have had the crew here oftentimes
4 during the summer. Can you go through what kind of
5 training you have done informally, how you have managed to
6 do that sort of without having a mission technically, and
7 has that saved any time to allow you to do it in May of '08
8 as opposed to whenever it would have been otherwise?

9 MR. BURCH: Okay. Basically, what happens is in
10 the course of putting these missions together, we go
11 through two phases with regards to the astronauts. We go
12 through a development phase, and then we go through a
13 training phase. The phase that we are currently in and
14 starting to approach completion of is the development
15 phase. So this is where we are trying out new tools and
16 new techniques for the first time, and we are trying to
17 make major advances in terms of getting the amount of time
18 that it takes to particular tasks reduced.

19 When we get into the hard-core training part, at
20 that point the emphasis shifts from developing new stuff to
21 refining the choreography and making fine adjustments.

22 During the past several months, we have been

1 focused pretty heavily on the STIS repair work and a lot of
2 these tools that need to be developed for that, a pretty
3 complex task. These are things that are all well within
4 the capabilities of astronauts and humans to do on orbit,
5 but we have had to develop some new tools and techniques,
6 and there were some things that we needed to understand
7 better about the electronics box on the STIS and its
8 serviceability. So we are doing some new stuff here, but
9 it has really come along very well.

10 And it has also given the astronauts an
11 opportunity to get themselves reacquainted with the
12 facilities here and some of the standard tools and
13 facilities that we have at Goddard for servicing missions.
14 The last mission was back in March of 2002. Folks get a
15 little rusty, you know, after a 4-year layoff like that.
16 So that has been what we have been focused on.

17 Did that answer the question? Okay.

18 MR. ACOSTA: Bob?

19 QUESTIONER: Bob Zimmerman, Freelance.

20 Actually, I have two questions. One, could you
21 go into the details of what is the difficulty of preparing
22 STIS, what the astronauts have to do?

1 And secondly, I would like to understand a little
2 bit about the two gyro mode. Are you using the two gyro
3 mode here on in, even after the gyros are replaced, or is
4 there going to be some policy with intention of going to --

5 MR. BURCH: Okay. Let me start with the last
6 question first.

7 We are currently in two gyro science mode.
8 Hubble has six gyros that are capable of supporting the
9 science mode. Two of them have failed, and they are off.
10 Two of them are powered off and are spares, and so we are
11 operating on two gyros at the present time.

12 Our plan is to continue to operate on two gyros
13 from now until the Hubble Servicing Mission is
14 accomplished. Actually, we are working on something that
15 is kind of neat at one gyro science mode, and that will be
16 ready in the spring of next year. We expect performance of
17 that to be very comparable to what we have today with the
18 two gyro science mode, but it hasn't been proven or
19 demonstrated yet on orbit.

20 Once servicing is accomplished, then we have six
21 brand-new gyroscopes, and we have them all checked out. At
22 this time, the plan is to go back to three gyro science.

1 Three gyro science does give us more flexibility in terms
2 of being able to schedule the astronomy target that we want
3 to look at with time, but we are constantly reevaluating
4 the benefits of things like that versus stretching the
5 useful life of the HST.

6 As far as the STIS repair goes, there were two
7 major challenges facing us. In order to get at the failed
8 electronics board that is inside the STIS main electronics
9 box, we need to take the cover off the box. We are very
10 fortunate in that when the astronauts open the doors to the
11 [inaudible] and they look at this instrument, that cover is
12 sitting right there in front of them. So they can get
13 relatively easy access to it.

14 The challenge was the 111 screws that are holding
15 it on, and the screws are not captivated. So they have to
16 go in there and take all these screws out. Well, you can
17 imagine what went through a lot of people's minds when we
18 first started thinking about this, you know, the 111 screws
19 floating all around inside Hubble. That was unacceptable.

20 So we came up with a very clever device called
21 the fastener capture plate which is basically made out
22 [inaudible]-type material, and this plate goes over top of

1 the MEB cover that is aligned and fastened on there. Then
2 this fastener capture plate has a series of little holes in
3 it that line up little screws. The holes are small enough
4 to allow the tool bit to go in, so you can turn the screw,
5 but they are small enough to keep the screw from falling
6 out. So, once you get all 111 screws taken care of, the
7 cover stays attached to the fastener capture plate and
8 moves the whole thing out. So all the debris and all the
9 screws is captured in there.

10 And then we have come up with an
11 astronaut-friendly replacement cover. Once we are done
12 servicing, they take the new cover, put it on, and there
13 are two latches. They just throw the latches, and Bingo,
14 it is on there, and then there is a third latch that they
15 throw that has some fingers that grab the electronics board
16 and mates them to the cover because this cover also acts
17 like [inaudible].

18 Okay. So that was challenge number one.
19 Challenge number two is actually getting the electronics
20 board -- I hope I am not scaring Mike about this.

21 Have you heard this story?

22 [Laughter.]

1 ADMINISTRATOR GRIFFIN: Actually, yes.

2 MR. BURCH: Oh, okay. I haven't gotten, you
3 know, the gaff yet.

4 So the second thing is the electronics board that
5 has the power supply on it is pulling that out, and if you
6 have ever fooled around with your desktop computer, those
7 things usually aren't much of a challenge, but the way
8 these instruments are built on Hubble, these boards slide
9 into slots in the box, but they are held in place by things
10 called wedgelocks, and the wedgelocks are designed to keep
11 the boards from rattling around, and they also provide a
12 heat path to reject waste heat up, the sides of the box, so
13 things stay nice and cool.

14 Unfortunately, these wedgelocks have a property
15 like these Chinese finger handcuffs that you may have
16 played with as a kid. You know, you put them on, and the
17 harder you pull, the tighter it gets, and you can't get
18 your fingers out. Well, the wedgelock has this kind of a
19 property, and when you loosen the bolts on them, sometimes
20 you can slide the board right out. Sometimes you have to
21 wrestle with it for a half hour or an hour to get it out.

22 We have one guy who is a technician at

1 [inaudible] Aerospace who knows how to do this really well,
2 but we are not planning on putting him in a spacesuit and
3 taking him up there. Even he, I think, is a little
4 mystified about why sometimes he has problems and sometimes
5 he doesn't. So we obviously needed a tool to overcome this
6 problem.

7 So we have a hard extraction tool that was
8 developed, and we went into a small research program to
9 see. Even if these wedgelocks jammed in their worst
10 possible way, could we pull the board out without having
11 the board disintegrate and leave a pile of debris all over
12 the place, and I am happy to report, we have come up with a
13 tool that enables us to do exactly that.

14 So those were the major challenges. There are a
15 few other things that need to be done too, before you take
16 the 111 screws out. There is an astronaut handle and
17 another small plate.

18 MR. ACOSTA: Okay.

19 MR. BURCH: Well, there are four different types
20 of screws, and I have forgotten offhand what they are, but
21 I think they are like [inaudible] and some slotted
22 [inaudible]-type screws, that sort of thing.

1 I think right now, the astronauts can actually
2 get all 111 screws out in 45 minutes.

3 PANELIST: Forty-five minutes.

4 One of the early worries was how long would it
5 take and would they get the teeth and so forth, and now we
6 envision a NASCAR wheel-changing operation.

7 MR. BURCH: Yeah. I actually bumped into a guy
8 who is sort of, I would say, a minor wheel on the NASCAR,
9 but anyhow, this guy is heavily involved. So I explained
10 to him the problem. There are some similarities there when
11 you see these guys come into the pit and they can take five
12 lug nuts off in 2-1/2 seconds and refasten it. It is
13 pretty impressive. So that is the goal, sort of.

14 MR. ACOSTA: NASCAR will appreciate the plug, I'm
15 sure.

16 [Laughter.]

17 MR. ACOSTA: Let's go over here. I think Brian
18 Berger has a question.

19 QUESTIONER: Brian Berger with Space News and
20 Space.com.

21 I have a question of Mike, but anybody can chime
22 in. That's fine.

1 You described a \$900-million cradle-to-grave
2 estimate for this mission as a lot of that being some cost.
3 How much of that \$900 million is money left to be spent,
4 and what is the rough breakdown between Bill's budget and
5 Mary's budget?

6 ADMINISTRATOR GRIFFIN: I don't have a breakdown
7 for you between Bill's and Mary's budget, but cost to go is
8 basically from now until May of '04 at 12 million a month
9 to support the SM-4 team, so, what, 18 months, that is 12
10 million a month. I don't have my calculator. Plus, of
11 course, the marginal cost of expended Shuttle hardware, as
12 I said, around \$100 million for a tank and a set of SRBs
13 (solid rocket boosters) --I am not trying to put too many
14 decimal places on it -- and then the cost of the actual
15 launch processing operation itself, and then, of course,
16 there will be a certain amount of processing for the Launch
17 on Need Shuttle.

18 Now, notice that the Launch on Need Shuttle is
19 not [inaudible] Station Shuttle. So you don't get to
20 book-keep the cost of the Shuttle to the Launch on Need
21 Mission, but any extra processing that would go on for the
22 Launch on Need operation that in the end will not take

1 place would be book-kept against Hubble.

2 So I hope that helps you, Brian.

3 MR. ACOSTA: Go ahead.

4 Brian has a follow-up.

5 QUESTIONER: As a quick follow-up, there was talk
6 before about grandfathering in this plan that Space
7 Operations Mission Directorate would pay for the Shuttle
8 for this mission. Is that still NASA's plan, or will the
9 Science Mission Directorate pay for the use of the Shuttle?

10 ADMINISTRATOR GRIFFIN: We will talk to you later
11 on the breakdown between space ops and space science.
12 Okay?

13 MR. ACOSTA: Okay. Next question. Let's keep it
14 right in the next row, Mark Kaufman.

15 QUESTIONER: Hi. Mark Kaufman with The Washington
16 Post. This is for the Administrator.

17 ADMINISTRATOR GRIFFIN: I enjoyed your article
18 the other day, by the way. Good job.

19 QUESTIONER: Thank you.

20 The first question is about we know that there
21 was a long process here of looking at whether or not to do
22 the Hubble mission and that you had a lot of your people in

1 on Friday, and I was wondering were there people who
2 recommended against going.

3 ADMINISTRATOR GRIFFIN: Well, first of all, they
4 are not my people. They are people who work for NASA, and
5 we got together to discuss it. I didn't have them in.

6 [Laughter.]

7 ADMINISTRATOR GRIFFIN: I try really hard not to
8 behave in that fashion. My mother would have disapproved.

9 We did have two meetings last week that were, to
10 be honest with you, deeply technical in nature and somewhat
11 extended. They were good discussions.

12 I don't believe I have talked to anyone in the
13 agency, from flight crew to flight ops managers to even
14 budget guys -- I don't believe I have talked to anyone who
15 thinks that we shouldn't do this.

16 There was an earlier question, was there
17 unanimity in the astronaut office about Launch on Need.
18 No, there wasn't, and there is not unanimity among the
19 engineers about Launch on Need or among the policy folks.

20 Part of why I exist is to resolve situations
21 where there is not unanimity. If there is unanimity, you
22 don't need me. So we resolved that by deciding that we

1 would do the Launch on Need, and there will be some extra
2 cost for that. I mean, life is hard, but you got to do
3 what you got to do.

4 QUESTIONER: And a second question, the Senator
5 was saying today that she was going to be sponsoring and
6 promoting the idea of a billion-dollar supplemental, where
7 you have been saying here the cradle-to-gave for Hubble is
8 900 million, or very close obviously, and I know that they
9 are kind of apples and oranges to some extent, but when
10 this is discussed in Congress, will there be an argument
11 that this money is in some way needed for Hubble?

12 ADMINISTRATOR GRIFFIN: It is well above my pay
13 grade to say what it is that Congress should or will
14 discuss.

15 I serve the President and support the President's
16 budget. Senator Mikulski was making a point that there was
17 in the aftermath of Challenger a supplemental to pay for at
18 least part of the cost associated with that loss, but there
19 has not yet been one for Columbia, and many other programs
20 in NASA were damaged to pay for Columbia, and it is her
21 desire to make up for some of that.

22 If her amendment, her and Senator Hutchison's

1 amendment passes, I am sure there will be some guidance
2 from the Congress on how to apply it, but I don't have any
3 comment on matters that at best could be speculative at
4 this point.

5 MR. ACOSTA: All right. We are going to do one
6 more question here before we go to Johnson Space Center,
7 and of course, we will come back to here at Goddard. So
8 bear with us. We will go right here for the last question
9 before we go to Johnson.

10 QUESTIONER: Hi. This is Mark Mathews with the
11 Orlando Sentinel. The question is for Preston.

12 What shape do you expect the Hubble to be in?
13 How many gyroscopes will be working? How many batteries
14 will be in order? What is the general shape of the Hubble?

15 MR. BURCH: Okay. Again, this is a little
16 speculative too about where we will be with gyros.

17 ADMINISTRATOR GRIFFIN: Don't speculate.

18 MR. BURCH: Yes.

19 According to our best estimate of gyro
20 availability, we think that the two gyro science mode is
21 viable up until approximately October of 2008. So the
22 planned window for this servicing mission is consistent

1 with our ability to continue to operate in two gyro science
2 mode.

3 Now, we are dealing with small-number statistics.
4 Nobody can tell exactly when the next gyroscope is going to
5 fail and which one it will be. Sometimes you get lucky,
6 and sometimes not so lucky.

7 So the point in time that I was giving you the
8 October 2008 is the 50/50 probability point, but inasmuch
9 as we have got potentially a one gyro science mode coming
10 online, I think we will be able to continue to keep science
11 going until we get the service.

12 As far as the batteries go, for the first 2 years
13 following Servicing Mission 3-B in March of 2002, we
14 experienced an extremely deep loss of battery-charged
15 capacity. We embarked on a program to try to understand
16 and try to improve or try to bring into operation improved
17 techniques that would ameliorate that.

18 Our measurements over the last 2 years seem to
19 point to success in this area. The loss of charge capacity
20 has pretty well flattened out over the last 2 years, but
21 once again, it is like you trying to predict when is the
22 battery in your car going to die. Batteries are basically

1 a big electrochemistry experiment, and there are a lot of
2 things going on in there that people don't fully
3 understand, but we think right now the batteries will be
4 fine, well into 2009 and probably longer. And then there
5 is a lot of other stuff going on, but I think those are the
6 main things that we are most concerned about.

7 MR. ACOSTA: Thank you, Preston.

8 All right. We are going to go to Johnson Space
9 Center. I believe we have a couple of reporters with
10 questions.

11 QUESTIONER: Mark Carreau from the Houston
12 Chronicle for Mike Griffin.

13 I noted the challenge in the assembly of the
14 Space Station and those missions. If you run into
15 difficulties and that schedule gets stretched out, how long
16 can you put off the Hubble mission until you would have to
17 sort of give up on it? Do you have any idea on that, and
18 what is your thinking on what takes priority?

19 ADMINISTRATOR GRIFFIN: Well, clearly, in both
20 Presidential policy direction and law, the Space Station
21 takes priority.

22 We strongly believe that we can accommodate the

1 Hubble mission without -- frankly, without significant
2 stress to the Station manifest or we wouldn't be doing it.

3 We are trying for a launch in early May of '08.
4 You know as well as I that missions slip, and as Preston
5 just indicated, even if we slipped to October of '08, we
6 will probably still be doing science operations, and if we
7 are not, it is not a big deal because the telescope is
8 emphatically not in any danger of degradation simply
9 because if it would be unable to do science operations and
10 is waiting for a repair, it is perfectly healthy and
11 perfectly stable, and that is a real concern.

12 So we have a 6-month window there that we are
13 planning in the manifest, and we are pretty sure we can get
14 a flight off in that 6 months. If it would go longer, we
15 would just do the mission at a later time.

16 It is very, very unlikely at this point that the
17 telescope would become damaged before we can get to it, and
18 that is part of our plan.

19 MR. ACOSTA: Gerst, do you want to add to that?

20 MR. GERSTENMAIER: Yes. The only thing I would
21 add a little bit is we are kind of targeting for spring of
22 '08 through fall of '08, and today, we currently sit on

1 Discovery, but we may very well move to a different orbiter
2 as the manifest plays out. So we reserve the right to kind
3 of change and move the manifest around as the Shuttle
4 flights and Station flights occur and as the situation
5 deems necessary on Hubble. So, if some things move on
6 Hubble that are going and turning in a wrong direction, we
7 will move some things around in the manifest to try to get
8 the Hubble with more assuredly than we would if we were
9 naturally in that sequence.

10 If we see some things on Station move around, you
11 will see some things move around.

12 So I think at this point, you need to give us the
13 degree of freedom that our manifest is going to move
14 around. So don't think that there is a particular vehicle
15 and a particular flight that is aimed. We want the ability
16 to be flexible over the next year or so to kind of move
17 that manifest around to optimize the needs of Hubble and to
18 optimize the needs of Station. So, again, we will react to
19 what life gives us, and we will change the manifest as
20 accordingly to optimize for both programs.

21 MR. ACOSTA: Mark, I believe you had a follow-up.

22 QUESTIONER: Thank you very much for the first

1 answers.

2 And for Dr. Leckrone, could you talk about what
3 the improved telescope could do with dark matter and why
4 that is sort of an important area of investigation at this
5 point?

6 DR. LECKRONE: Right. As I indicated, the --
7 well, let me back up a second with something I wanted to
8 say and didn't even take the time to do it.

9 The reason the Nobel Prize in physics that went
10 to John Mather and the COBE (Cosmic Background Explorer)
11 team here at Goddard and also to their colleagues in
12 California was that work led directly to this picture of
13 the universe as having kind of a cosmic web structure, a
14 skeletal structure. That is dark matter. You can't see
15 it. You don't know what it is, but you do know that it is
16 about 24 percent or so of the mass energy budget of the
17 whole universe. Whereas, the ordinary hydrogen and helium
18 and other matter and so forth that make up you and me are
19 roughly 4 percent of the mass energy budget of the whole
20 universe. So this is a dominant source of ordinary gravity
21 in the universe.

22 There is that ordinary gravity, starting at the

1 time of the initial big bang and the tiny fluctuans in the
2 cosmic microwave background that evolved into this skeletal
3 structure that looks like a three-dimensional cosmic web.

4 So the short answer belong that extended prolong
5 is that we will not learn directly very much at all about
6 the nature of dark matter. That is going to be up to the
7 highest energy particle accelerators and so forth. No one
8 knows as of today what dark energy us -- sorry -- dark
9 matter is. We don't know what dark energy is either, but
10 what we will learn is more about the structure and
11 particularly how it changed with time and particularly how
12 its composition changed with time.

13 So the dark matter, I would like to think of it
14 like a Christmas tree where the branches of the tree are
15 dark in a darkened room, but the illuminated stars and
16 galaxies are like the Christmas tree bulbs that you see.
17 So they give you the shape of the dark matter, the shape of
18 its structure, although you don't see the dark matter
19 directly.

20 MR. ACOSTA: All right. I believe there is
21 another question at Johnson.

22 QUESTIONER: Gina Sunseri, ABC News, for Bill

1 Gerstenmaier.

2 Bill, could you go into a little more detail? I
3 am still not quite sure how far along processing will go on
4 a Shuttle for Launch on Need. It will be on the launch pad
5 as SM-4 launches when you have gone through TCDT (Terminal
6 Countdown Demonstration Test)? Give me a little, if you
7 can talk me through that process a little more, please.

8 MR. GERSTENMAIER: Okay. I think we are still
9 putting all of the details together, but if you think about
10 it just basically, we have about 25 days of on-orbit
11 lifetime of the Shuttle that is up at the Hubble Space
12 Telescope. So, if we need a Launch on Need, we have to be
13 essentially within 25 days or so of being able to go launch
14 that next Shuttle. So we will take it fairly far along in
15 the count. We will probably be out at the pad. We will
16 probably be fueled at this point. We probably will have
17 done some kind of TCDT to be ready, and then we will be in
18 that process.

19 And again, if you think about it, we are
20 protecting for a fairly low probability event, but we are
21 doing it in kind of a clever way. We are not going to be
22 wasting that orbiter in any way, shape, or form. All of

1 this activity we have done, all of the fueling, the TCDT,
2 all of that activity will apply directly to that next
3 Station mission. So all we will have is an empty cargo
4 bay. That is a difference between a Station mission and
5 this rescue mission is in a Station mission, you will have
6 the Station cargo in the bay, and in this mission, you will
7 just have an empty bay, and you will be going to a
8 different inclination. So we will have to do the flight
9 planning for a different inclination, but we will take it
10 pretty far along in the sequence, in the processing.

11 The exact details of that will come over the next
12 couple of months as we see how all of this folds together
13 and we get more detailed planning together and see where it
14 is, but, again, the goal is to optimize the use of this
15 Launch on Need with the next mission that is coming, which
16 would be the Station mission. So we will find lots of
17 commonality between the two and minimize the impacts of
18 this Launch on Need mission.

19 MR. ACOSTA: Plus, I think we should remind folks
20 that we made the announcement of the decision today. We
21 still work details. That is not unusual to --

22 ADMINISTRATOR GRIFFIN: Let me add a comment.

1 That is exactly the point.

2 What we required in order to make this
3 announcement was to see a path through the wilderness to
4 get -- as I had said earlier in my remarks, to get the yes,
5 to get to a point where we knew we could do this.

6 The guys have got 18 months between now and when
7 we launch to figure out smarter ways, and we expect them to
8 do so, and we want a pat on the back for that, not a
9 criticism if we change, you know, puppy to small dog.

10 We expect people at NASA to be flexible,
11 creative, intelligent, and bold in their thinking on ways
12 to be efficient and at the same time be careful, and I
13 think that is what you will see.

14 MR. ACOSTA: All right. From Johnson, now we
15 will go to Kennedy Space Center for a question.

16 QUESTIONER: Thanks. This is Todd Halvorson of
17 Florida Today for Preston Burch.

18 Preston, what percentage of the original charge
19 capacity do the batteries have at this point, and how fast
20 would systems freeze up and a servicing mission become
21 undoable if you have lost charge capacity completely?

22 MR. BURCH: Right now, it is approximately

1 one-half of what we had -- it is a little more than
2 one-half of what we had at launch 16 years ago.

3 Current Hubble system capacity is approximately
4 297 amhours. I think we had somewhere around 550 amhours
5 at launch.

6 The batteries obviously are needed to support
7 Hubble's electrical load when we are in the nighttime
8 portion of the orbit which is approximately one-third of
9 each orbital period, or about 35 minutes.

10 What happens when the batteries die is that we
11 will most likely lose control of the observatory. We won't
12 be able to -- systems won't be functioning, and things will
13 start getting cold, and once that happens, in approximately
14 2 days, we estimate, temperatures will drop to a critical
15 point where the metering truss, which has hardware mounted
16 to it using titanium pads that are bonded on there, those
17 bonds will become compromised. So we will not be able to
18 guarantee the alignment of the various instruments and
19 guidance sensors and other optical pieces on the metering
20 truss. So we have about 2 days once we lose power on the
21 observatory before it will probably become useless as a
22 fine precision instrument.

1 MR. ACOSTA: All right. I believe, Todd, you
2 have a follow-up.

3 QUESTIONER: Yes. Thanks.

4 For Dr. Leckrone, will this mission bridge the
5 gap between Hubble and the James Webb Space Telescope, and
6 can you talk about or elaborate on what you think the two
7 or three cosmic mysteries you will be able to address are
8 when you have six fully functional instruments?

9 DR. LECKRONE: Sure. Well, as I tried to imply
10 earlier, yes, it absolutely will bridge a gap. It is good
11 to know a little bit in advance about the territory in
12 which you are embarking for the first time, and so we are
13 hoping to provide basic information for the sciences who
14 will be using the James Webb-based telescope that will
15 allow them to ask the most intelligent questions or the
16 most penetrating questions about the subject matter, and in
17 particular, a major objective of the James Webb Space
18 Telescope program is to thoroughly explore and understand
19 this nursery of the universe when the first stars and
20 galaxies were just coming into being, or proto galaxies.

21 So, yes, we will provide a bridge. I should
22 comment, of course, we are not only slated here for a

1 5-year extension of our mission. Typically, in the space
2 program, space observatories don't fail precipitously.
3 They generally fail rather gracefully, and I think there is
4 some hope that we might have a year or two as overlap with
5 the James Webb Space Telescope, and that would be just
6 terrific because we still do ultraviolet and optical in the
7 James Webb that an infrared telescope is not able to do,
8 and I think we could hold hands very nicely as a tandem set
9 of scientific tools.

10 Let's see. Your other question, what are the
11 major things we could do. I haven't said very much about
12 the Space Telescope Imaging Spectrograph, but if we can
13 repair it and bring it back online, it is a bit of a
14 different instrument than the Cosmic Origins Spectrograph.

15 The Cosmic Origins Spectrograph is intended to
16 focus on getting as far across the universe as quickly as
17 possible, to be as sensitive as possible.

18 The STIS spectrograph is a much more versatile
19 instrument. It has many different tools and capabilities
20 for different kinds of science.

21 The STIS has the honor of having been the first
22 instrument to directly detect and measure the atmosphere of

1 a planet around another star, and we did that. We detected
2 sodium. We detected oxygen. We detected hydrogen in one
3 of the so-called "hot Jupiters" around a particular star.

4 Shortly thereafter, this failed, and it turns out
5 now that there is a rather long list of candidate stars or
6 candidate proto planets or candidate planets around other
7 stars, XO planets, that scientists would really like to
8 have a look at with STIS to see if they can do similar
9 analyses and at least similar detections of their
10 atmospheres.

11 In fact, in a very short time, just within a
12 couple of weeks or so after STIS failed, Dave Charbonneau
13 of CalTech, who is a principal investigator in this kind of
14 work on Hubble, was lamenting, "Hey, I have got two more I
15 was planning to look at with STIS, and now it's gone."

16 So I really believe that we should be able to get
17 up to perhaps 10 or 12 XO planets whose atmospheres have
18 been sensed if we can repair this instrument.

19 I think we will continue to make significant
20 progress on the dark energy front. It is a statistics
21 game. You just need to observe more and more of the
22 [inaudible], the standard flight sources, the standard

1 candles to tell you how far away a galaxy is in order to
2 sample over different periods of time what has the
3 expansion rate of the universe been, and simply the more
4 systems like that you can detect, the tighter your error
5 bars will be on the parameters that we are currently using
6 to describe dark energy, and in particular, is dark energy
7 changing with time or is it truly a constant, as Einstein's
8 cosmological constant implies? And I think we can help
9 answer that question with Hubble.

10 There are plans in the works for future dark
11 energy missions. I think those are very valuable. I think
12 those will be paradigm-shifting. Hubble has already
13 shifted the paradigm by being a major contributor to the
14 discovery of dark energy and helping to narrow down the
15 range of possibilities for what dark energy is, but I think
16 it is important that those future missions ask the right
17 questions, and I really do believe we can push this field
18 significantly farther with Hubble, so that when the time
19 comes to build a dark energy mission, we are actually
20 asking the right questions.

21 MR. ACOSTA: All right. We have one last
22 question from Kennedy, and then we are going to go to

1 Headquarters, and then we will come back to Goddard.

2 Kennedy?

3 QUESTIONER: This is Stephen Young with
4 SpaceFlightNow.com for Preston.

5 I am just wondering in the 18 months between now
6 and the launch if you see any unexpected failures. Do you
7 have any capacity to incorporate any additional equipment,
8 repairs on this mission, or would it be a case of having to
9 drop something else if you needed to add another task?

10 MR. BURCH: I think it depends on what fails. We
11 have had this happen on two prior missions.

12 The particular device that failed or became
13 anomalous was a reaction wheel, and we have a spare handy,
14 and we were able to put that on our manifest late in the
15 game and get it ready.

16 We do have some spare time available to us at the
17 moment in the EVA (extravehicular activity) schedule. If
18 you look at the time over the 5 days, there is potentially
19 additional time to do some small tasks.

20 We do also have potentially additional room and
21 additional weight capacity to take additional hardware up
22 to Hubble. So if it is something very large and very heavy

1 and very time consuming, then we would probably be faced
2 with making some tough choices. We haven't had that happen
3 to us in the past, but we are certainly well aware that
4 that sort of thing could happen.

5 So we have an expensive long-term trending
6 program analyzing performance of all the components and
7 subsystems on the observatory, and at the moment, there is
8 no indication of anything like that.

9 Hubble also has tremendous internal redundancy in
10 all of its systems, and that would be another tradeoff of
11 how much redundancy do we want to keep in a particular area
12 versus the benefit to be gained.

13 MR. ACOSTA: Thank you, Preston.

14 Let's go to Headquarters. Traci?

15 QUESTIONER: Traci Watson, USA Today, for the
16 Administrator. I'm sorry I missed your earlier discussion
17 before this press conference started about the cost of the
18 mission, and I am wondering about that \$900 million.

19 Do you have that in hand? Where would that come
20 from?

21 ADMINISTRATOR GRIFFIN: That was asked and
22 answered. We have not in detail identified the particular

1 sources within the agency, but obviously, it comes from the
2 existing science astrophysics program and existing Shuttle
3 space flight lines, and how that will be broken down is not
4 yet determined.

5 MR. ACOSTA: All right. Let's come back here to
6 Goddard. Any questions?

7 QUESTIONER: Thank you, and a follow-up for
8 Preston Burch.

9 What does it mean to you to have veterans like
10 Mike Massimino and John Grunsfeld working on this mission?

11 What do they bring to it?

12 MR. BURCH: Well, obviously, they bring a wealth
13 of experience. John has been up there twice before. Mike
14 has been up there once before.

15 In addition to being familiar with Hubble, they
16 are familiar with all the tools, the facilities that we
17 have here, and so that really helps us to make the training
18 more efficient. They can focus more effectively on the new
19 things, the new challenges that are unique to this
20 particular mission, and they can also act as mentors to the
21 rookies who will be flying with them.

22 So we have generally tried to do that on previous

1 servicing missions to have at least one or two astronaut
2 that are doing the EVAs, the spacewalks, to be on the next
3 mission to reduce the risk and to make things more
4 efficient.

5 So we are extremely pleased to have John and Mike
6 back with us. They have been working with us for the past
7 several months on the STIS repair as well as other things
8 we will be doing on this mission. So we think that is a
9 great decision.

10 ADMINISTRATOR GRIFFIN: Well, I think also you
11 would want to point out the value of having Scott Altman as
12 the Shuttle Commander who has also been to Hubble, is
13 familiar with the unique proximity operations concerns that
14 are associated with flying around Hubble.

15 MR. BURCH: Yes. That is certainly very true,
16 and somebody was reminding me this morning about all of the
17 non-piloting stuff that Scott does in supporting servicing
18 once we are on orbit and Hubble is docked to it.

19 Scott gets in there and does a whole range of
20 other duties that are directly in support of servicing. So
21 he knows Hubble very well.

22 MR. ACOSTA: Great. All right. Now I am

1 serious. Now we are coming back to Goddard. All right.
2 So now let's go. Who hasn't been able to ask a question?
3 Let's go to the back row, and then we will go over to Alan
4 after that. We have about 7 or 8 minutes.

5 QUESTIONER: David Gaines, Independent.

6 A question for Dr. Cleave and also Dr. Griffin.
7 An opportunity for you, I think, now to maybe respond to
8 the scientists and the scientific community who have often
9 expressed their skepticism over the past few years over the
10 priorities and commitments to science by the Presidential
11 administration and then by extension, of course, to NASA's
12 budgets as the funding pie, if you will.

13 DR. CLEAVE: Well, this is a great demonstration
14 I think of the agency commitment to science. I think it
15 would have been probably easier for everyone just to stay
16 the course with not doing a Shuttle servicing mission, and
17 it took real commitment on behalf of the agency in order to
18 put this science extension back on the table, and we are
19 hopeful we will get it done. So it is a very good example
20 of that.

21 Mike?

22 ADMINISTRATOR GRIFFIN: Thanks. I will follow

1 up.

2 I am bothered by the basic premise of the
3 question. The fact is that 32 percent of NASA's budget
4 right now is science. It is \$5.3 billion this year or this
5 past year. We are doing pretty well.

6 There may be a few individual scientists whose
7 hope for influence over the program will not materialize in
8 the way that they would have personally hoped, but the
9 influence of the science community as a whole at NASA and
10 among the Nation's policy-makers is enormous.

11 The premise of the question is that the
12 administration has not been friendly to science, and yet, I
13 had the full backing of the administration in changing our
14 plans for the use of the Shuttle in its fly-out phase to do
15 this mission.

16 There is absolutely nothing at NASA that is not
17 science-friendly. We exist in part to do the kinds of
18 things that you can't do without being in space, but notice
19 that I said exist in part. There are other things also
20 that we do at NASA.

21 We have a robust aeronautics program, and we
22 would like it to be more robust. We have international

1 commitments going back 20 years to finish assembling the
2 Space Station. That simply must be done, if we are able to
3 do it.

4 We want, once again, to go beyond low-Earth orbit
5 in exploring with human beings, and if we are to have a
6 future, as pointed out by the Gehman Commission, that too
7 must be done.

8 The art is in balancing all of those things. No
9 one enterprise at NASA gets all at once, of course, and yet
10 the art is in balancing it to see that no one enterprise
11 within NASA is unduly harmed in relationship to others.

12 It would be nice. It would be very nice if I
13 could see throughout the entire space community the kind of
14 teamwork that I have seen displayed between and among the
15 mission directorates in trying to pull this Hubble
16 servicing mission together.

17 MR. ACOSTA: All right. Let's go over here to
18 Alan.

19 QUESTIONER: Alan Boyle with MSNBC. I had a
20 question for Bill Gerstenmaier.

21 There had been discussion during the buildup to
22 the recent Space Station mission that this was going to be

1 the most complicated series of operations of spacewalks
2 that we have seen in recent times. How would you compare
3 the spacewalks required for the Hubble repair mission with
4 the Space Station spacewalks? Would you say that Hubble is
5 actually more complicated?

6 MR. GERSTENMAIER: Well, it is really hard to
7 compare the two. I think they are both complicated, but
8 maybe in different ways.

9 In the Station world, the complication is it is
10 not one mission. The missions are all linked together. So
11 the four EVAs or three EVAs that we are going to do on this
12 next mission are then followed by three-stage EVAs when the
13 Shuttle is gone, staged out of Space Station. So the
14 complexity is the Shuttle EVAs then impact the EVAs that
15 are done during the period when the Shuttle is not there,
16 and then those EVAs are then necessary for the next Shuttle
17 flight to go fly.

18 So the complexity is not only in the individual
19 mission, but it is in the assembly of missions and the way
20 they are interlinked.

21 Hubble as a mission itself is very complicated
22 with the five EVAs and the planning, but, again, I think

1 they are complementary.

2 What I see here that is kind of neat is that we
3 have been doing this assembly stuff in Station. We have
4 been doing the inspection techniques. We have been using
5 the boom. So the things that we are going to have to do on
6 the Hubble mission, we have already demonstrated to a small
7 extent on the couple Station flights we have done so far.
8 We will do more of those before the mission flight. So it
9 is a neat kind of complementary action back and forth.

10 The other thing I find very interesting is when
11 Preston described to you the EVA to go replace the card
12 outside. If you look at what we are doing on Station,
13 doing very similar things on the inside, you know, I
14 repaired a gyro where I removed 70 fasteners and popped off
15 a bearing instead of replacing the 75-kilogram instrument,
16 very analogous to what Preston is going on the outside.

17 What I find intriguing is how we are
18 complementary. He has this clever technique to capture
19 screws. Well, this is a neat technique. We are going to
20 use that on Station. So you are going to see this little
21 plastic overlay show up on some Station stuff, where we
22 have not capped the fasteners.

1 So we are going to learn synergistically or
2 complementary with what they are doing on Hubble and fold
3 it right back into Station. As we learn new techniques on
4 Station, we will fold those right into Hubble. So they are
5 not distinct or competing, but I think we have got a
6 tremendous chance here to keep learning and moving forward.

7 MR. ACOSTA: A good point to highlight.

8 All right. Anybody that hasn't had an
9 opportunity to ask a question that wants to ask a question?

10 All right. We have time for just a couple more. Let's go
11 up front here. We will go to Bob first and then Seth, and
12 that will be the last question.

13 QUESTIONER: Bob Zimmerman, Freelance again.

14 I asked this question of the Administrator
15 earlier, and you punted, probably rightly, but can anyone
16 give me any specific information about at least your goals
17 or your plans for this soft capture mechanism? Is it going
18 to be comparable to a docking mechanism that is on ISS, and
19 what will its capability be? At least your goal as to what
20 you are installing, anything you can tell me at all would
21 be appreciated.

22 DR. LECKRONE: Do you want me to field that one?

1 Okay. The soft capture mechanism utilizes the
2 low impact docking system technology that is being
3 developed for CEV. So what we would like to have is a
4 system that is very general purpose that could be used by
5 either a CEV (Crew Exploration Vehicle)-type vehicle coming
6 back to Hubble to install the propulsion module to the
7 orbit Hubble or potentially by an unmanned robotic vehicle
8 that would be launched on an expendable launch vehicle.

9 Right now, Hubble is a tough vehicle to grasp
10 onto. It was really designed to be launched, serviced, and
11 retrieved by the Space Shuttle, and it has these grapple
12 fixtures on it, but they are located in an area that is
13 difficult for anything other than the Space Shuttle to use.

14 And then we have these, what we call "towel
15 bars," these small racks that are on the [inaudible] heads.
16 Those are, again, not made to be grappled by an arm or by
17 an automated device. So, by attaching the soft capture
18 mechanism to the back of Hubble, we are going to make it a
19 lot easier, a lot less costly, and a lot less risky for a
20 future mission to go to Hubble when it comes time to go to
21 Hubble when it comes time to deorbit it.

22 MR. ACOSTA: All right. Thanks, and then we will

1 go --

2 DR. LECKRONE: And one last point, it is a
3 totally passive system. There aren't any -- once it is
4 installed, there are no active devices on it. So whatever
5 comes up to it would have the active side to it.

6 MR. ACOSTA: Okay. Seth?

7 QUESTIONER: Seth Borenstein, AP, for Dr.
8 Leckrone.

9 Just a little more specifics. The furthest back
10 you have been able to gaze with Hubble now and how much
11 further in light years, however you want to put this, will
12 you be able to, if everything works out well -- you talked
13 about what you will get, but in terms of the actual
14 quantity, distance.

15 DR. LECKRONE: Okay. Well, as I mentioned, right
16 now, the universe is estimated to be 13.7, plus or minus,
17 .2 billion years old. That's a result primarily out of W
18 map. So, when we are talking about something that emitted
19 its life when the universe was, say, only 700 million years
20 old, that was about 13 billion years ago.

21 And right now, we think we have detected objects
22 in visible light that emitted their light at around that

1 time, around 700 million years ago, redshift to 6 or
2 greater, if you are familiar with that jargon.

3 How much more deeply we could go depends on the
4 properties of what the universe was like at that time, how
5 opaque was the material, how expensive was the re-
6 ionization process, had it fully completed or only
7 partially completed or so forth, and I really can't give
8 you a quantitative answer. You will really have to do the
9 experiment to see.

10 All I can say is that in principle, since we are
11 going now from visible light to near infrared light, we are
12 going to, in principle, be sensitive to objects that emit
13 most of their light in the infrared that is not really
14 sensed very easily or at all as visible labeling, and that
15 corresponds to higher redshifts, maybe 7, maybe 8. If we
16 are very lucky, the most luminous objects there, the
17 brightest output objects, we might even get to a redshift
18 of 10, but that is common-order speculation, and we really
19 have to do the experiment to see.

20 I guess if I were setting a goal for us to get to
21 a redshift to detect one object in a redshift to 10, that
22 would be an interesting goal.

1 Well, I am talking about -- again, I can't do the
2 calculations off the top of my head, but I am talking about
3 the first stars coming into being, about 400 million years,
4 and the end re-ionization period, about 700 million years.
5 So you are not talking about a very large -- I mean, things
6 compress in time during that period. Very large delta Z
7 corresponds to very small delta T, and so we might be going
8 back to 700 million years from the origin, maybe 4-, 5-,
9 600 million years. Again, it depends on the nature of the
10 object, the opacity, how easily light transmits through the
11 material, how extensive re-ionization has been, and
12 frankly, those are all questions. Those aren't statements.
13 Those are things that we don't know that we want to find
14 out about.

15 ADMINISTRATOR GRIFFIN: Well, I think part of the
16 point is that without regard to how far back you go or
17 whatever, it is a very dynamic period --

18 DR. LECKRONE: Yes.

19 ADMINISTRATOR GRIFFIN: -- when things are
20 changing rapidly, and to probe a little further back into
21 it gives you an enormously advantaged concept or view of
22 what was going on.

1 DR. LECKRONE: Right. Could you help me write a
2 proposal?

3 [Laughter.]

4 ADMINISTRATOR GRIFFIN: You probably, actually
5 wouldn't want that.

6 [Laughter.]

7 MR. ACOSTA: Well, on that note, we are going to
8 go ahead and close out our press conference today. Any
9 final thoughts or comments?

10 [No response.]

11 MR. ACOSTA: I didn't think so.

12 Just a reminder that the Shuttle crew press
13 conference will be coming up at 2:30 Eastern. So you will
14 have that ability to watch on NASA Select.

15 Also, all the video and images that were shown
16 today during our press conference will be available right
17 after this press conference. We will have it on NASA
18 Select, and of course, if you wanted any more information,
19 go to www.NASA.gov/Hubble for all the information on
20 today's announcement and about Hubble.

21 We thank you very much, and have a great
22 afternoon.

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